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Contributors

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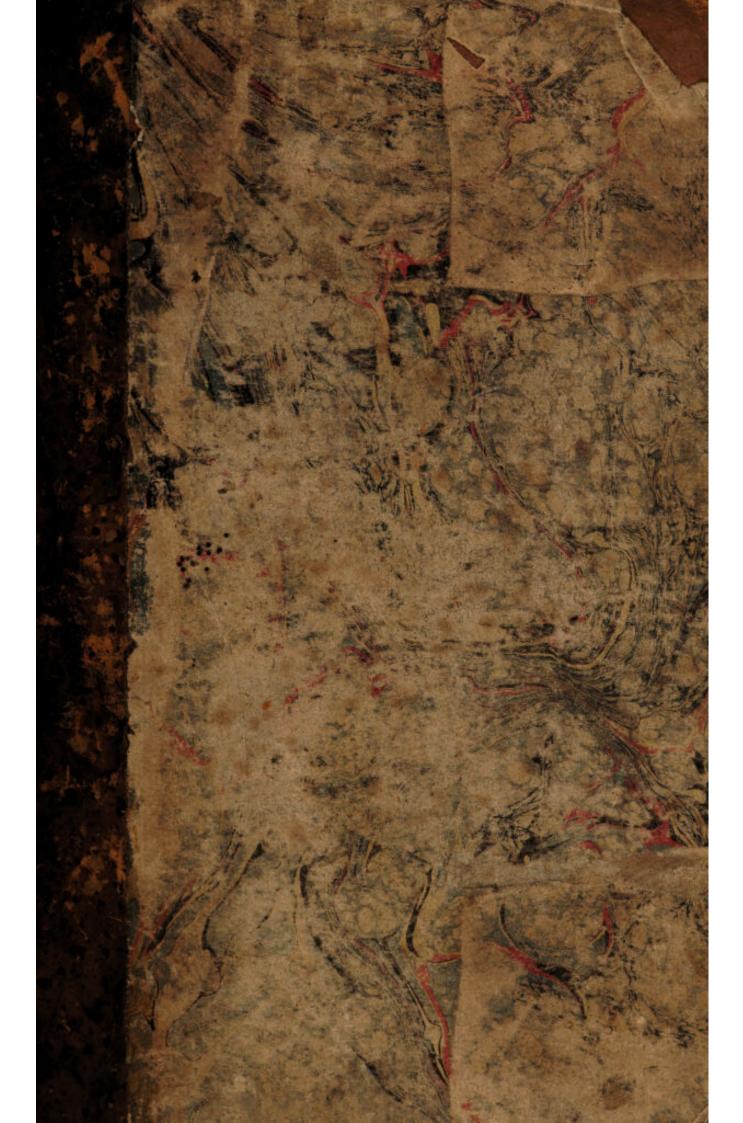
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GEORGE R.

EORGE the Second by the Grace of God, King of Great-Britain, France and Ireland, Defender of the Faith, &c. To all to whom these Presents shall come Greeting: Whereas Samuel Dunn Teacher of the Mathematics, and William Owen, Citizen and Bookseller of the City of London, have by their Petition humbly represented unto Us, that they have with great Study, Care, Labour, and Expence invented, delineated, and perfected a Work entitled,

The Universal Planispheres, or Terrestrial and Celestial Globes in Plano;

BEING

Compleat Substitutes for a large Pair of Globes, with their Uses in the Solutions of all the Capital Problems in Geography, Astronomy and Navigation.

Which Work the Petitioners humbly apprehend will be of great Benefit to the Public, and tend to the promoting of the Knowledge of Geography, Aftronomy and Navigation. The Petitioners therefore pray Us, to grant Our Royal Licence to them for the fole Printing, Engraving, Publishing and Vending the faid Work, for the Term of Fourteen Years, according to the Statute in that Cafe made and provided; We, being willing to give all due Encouragement to Works of this Nature which may be

Of Public Use and Benefit,

Are graciously pleased to condescend to their Request, and do therefore by these Presents, as far as may be agreeable to the Statute in that Behalf made and provided, grant unto them the faid Samuel Dunn, and William Owen, their Executors, Administrators and Assigns, Our Royal Privilege and Licence, for the fole Printing, Engraving, Publishing and Vending the said Work for the Term of Fourteen Years, to be computed from the Date hereof, strictly forbidding and prohibiting all Our Subjects within our Kingdoms and Dominions, to reprint, abridge, or translate the same; either in the Like, or any other Volume or Volumes whatfoever, or to import, buy, vend, utter, or distribute, any Copies thereof, reprinted beyond the Seas, during the aforefaid Term of Fourteen Years, without the Consent and Approbation of them, the faid Samuel Dunn and William Oaven, their Executors, Administrators, or Assigns, by Writing under their Hands and Seals, first had and obtained, as they and every of them, offending herein, will answer the Contrary at their Peril: Whereof the Commissioners and other Officers of Our Customs, the Master, Wardens and Company of Stationers of Our City of London, and all Our Officers and Ministers whom it may concern, are to take Notice, that due Obedience be rendered to Our Pleafure herein fignified.

Given at Our Court at Kenfington, the Seventh Day of November, 1757, in

the Thirty-first Year of Our Reign.

By his Majesty's Command.

DESCRIPTION and USE,

OFTHE

UNIVERSAL PLANISPHERES;

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TERRESTRIAL and CELESTIAL

GLOBES in PLANO.

BEING

Compleat SUBSTITUTES for

A LARGE PAIR of GLOBES.

WITH

Their Applications, in the SOLUTIONS of all the Capital PROBLEMS

IN

GEOGRAPHY, ASTRONOMY, and NAVIGATION.

Invented and Delineated

By SAMUEL DUNN,

Master of an ACADEMY for boarding and qualifying young Gentlemen, in Arts, Sciences and Languages, at CHELSEA.

The Second EDITION, with Additions and Improvements.

LONDON:

Printed and Published (according to an Act of Parliament)

For the Author, and W. Owen, at Homer's-Head, near Temple-Bar, Fleet-street.

MDCCLIX.



RECOMMENDATIONS of the Universal Planispheres.

R. DUNN, Teacher of the MATHEMATICS in London, having requested my Opinion of his UNIVERSAL PLANISPHERES, or terrestrial and celestial GLOBES in Plano; I have carefully examined those ingenious and compleat Substitutes for a large Pair of Globes; I have duly considered their Use in the Solutions of the capital Problems in Geography, Astronomy, and Navigation; and humbly apprehend, that nothing of that Nature, yet published, can be so conducive to the Assistance of Teachers, and Advantage of their Pupils, as what Mr. Dunn's Invention appears to me: And I hope, that my good Intentions, in recommending them to such as have not seen them, will not be construed into Presumption; since the Merit of the Performance is so obvious, that it cannot but gain the Approbation of those, who may be hereafter induced to enquire for them.

Carey-street, Lincoln's-Inn-Fields, Dec. 23, 1757. Stephen Demainbray, L. L. D. Teacher of Philosophy to his Royal Highness the Prince of Wales.

Onsieur Dunn, Maitre de Mathematiques à Londres, m'ayant demandé mon Sentiment de sa Planisphére Universelle C. D. de ses Globes terrestres & celestes in Plano. Je les ay trouvé après un examen scrpuleux, très propres à représenter les Globes, j'ay fais mes Reslections sur l'utilité dont ils pourroient être dans la Solution des Problèmes principaux de la Géographie, de l'Astronomie, & de la Navigation, & rien de ce qu'on a publié jusqu'icy dans ce Genre, ne ma parût si propre tant à faciliter la Connoissance de ces Sciences à ceux qui s'y appliquent qu'à soulager ceux qui les enseignent; c'est pourquoy je me slatte que ce que je viens d'en dire d'avantageux ne me sera point imputé comme prévenu en sa faveur; puisque je n'ay d'autre dessein que de recommander, un ouvrage si utile à la Societé, & dont on reconnoitra le Merite dès qu'on l'aura vû.

Londres, de Carey street, Lincoln's-Inn-Fields, Dec. 23, 1757. Etienne Demainbray, L. L. D. Precepteur en Philosophie de son Altesse Royale Monseigneur le Prince de Galles. Have examined Mr. Dunn's Universal Planispheres, and look upon them to be an useful and very ingenious Invention, founded upon unexceptionable Principles; by Means whereof the most important Problems in Geography, Astronomy, and Navigation, are solved with great Accuracy, Ease, and Expedition.

London, Dec. 1757.

T. Simpson, F. R. S. Member of the Royal Academy of Sciences at Stockholm, Mathematical-master in the Royal Academy at Woolwich.

J'ay examiné les Planisphéres Universelles de Monsieur Dunn, & les ay trouvé être une Invention trés utile & trés ingenieuse, fondée sur des Principes infallibles; par le moyen des quelles on peut resoudre les Problemes les plus importans, soit en Geographie, Astronomie, ou Navigation, avec une grande exactitude, & une grande Commodité, & Promptement.

Londres, 10bre 1757.

T. Simpson, F. R. S. Membre de l'Academie Royal des Sciences de Stockholm, Maitre Mathematicien de l'Academie Royal de Woolwich.

Certify that I have examined Mr. Dunn's new Planispheres, and am satisfied they are constructed on very true Principles; that by them all spherical Problems may be readily projected, and solved by Inspection, without the Help of Scale and Compasses, Pen and Ink, and that with more Facility and Accuracy than by any Globes of equivalent Dimensions.

London, Dec. 1757.

J. Bevis, M. D. Member of the Royal Academy of Sciences at Berlin.

Le certifie avoir examiné les Planisphéres de Monsieur Dunn, & en être satisfait elles sont construites sur de vrays Principes; & par leurs secours tous Problemes Sphériques peuvent être projettés & promptement resous, sans le secours, d'échelle, n'y de Compas, de Plume n'y d'Ancre, & ce la avec plus de facilité & d'Exactitude que par aucun autre Globe d'une Dimension equivalente.

Londres, Dec. 1757.

J. Bevis, M. D. Membre de l'Academie Royale des Sciences a Berlin.

Have seen and examined Mr. Dunn's Universal Planispheres, and find them to be four large stereographical Maps, with a transparent Index placed over each Map, whereby the Circles of the Sphere are instantaneously projected on the Plane of the Meridian for any Latitude, and the Problems of Geography, Astronomy, and Navigation, wrought with the same Certainty and Ease, as by the Globes themselves. And as I have read Lectures above 20 Years, and travelled through most Parts of England, and beyond the Seas; and many of my Pupils may be willing to be certified of the Utility of this Performance; I am desirous of assuring them, that the Principles upon which these Planispheres and their Operations do depend, are undeniably true, the Use of them easy and general, and in my Opinion, they must be of great Use in the Study and Practice of the above Sciences.

London, 26 Dec. 1757.

William Griffis, Lecturer in experimental Philosophy.

J'ay vû & examiné les Planisphéres de Monsieur Dunn, & les ay trouvé être quatre grandes Mappes stereographiques, avec un Index transparent, & placé au haut de chaque Mappe, par le moyen de quoy, les Cercles de la Sphére sont projetté trés vite, sur le Plan du Meridian pour toute Latitude, & les Problemes de Geographie, Astronomie, & Navigation, resous avec la même certitude & facilité que par les Globes mêmes. Et comme je lis depuis passé 20 ans, & ay voyagé par presque toute l'Angleterre, & delà la Mer; plusieurs de mes Eleves peuvent être assurés de l'utilité de cette ouvrage; je veux les assurér aussy, que les Principes sur les quels ces Planisphéres sont sondées & leurs Operations, sont indubitable, leur Usage facile & general, & selon mon Opinion, d'ungrand Usage pour la Pratique & l'Etude des Sciences cy-dessus mentionnées.

Londres, le 26 Dec. 1757.

Guillaume Griffis, Lecteur en Philosophie experimentale, WE whose Names are hereunto subscribed, Teachers of the Mathematics in and near London, having seen and examined the Construction and Use of the Universal Planispheres invented by Mr. Dunn; do find them to answer in the most easy and general Manner all the Problems usually answered by the Globes, with equal Certainty and Expedition, and therefore it is our Opinion, that this Work will be of great Use to the Public.

J. L. Cowley, Mathematician, Ludgate-bill.
J. Welch, Teacher of Navigation, Hatton-garden.
Charles Brent, Author of the compendious Astronomer.
Thomas Adams, Teacher of Mathematics, Sobo-square.
Edward Moore, Teacher of Navigation, and Globe-maker,

Deptford.

Peter Ruffel, private Teacher of Geography, and the Use of

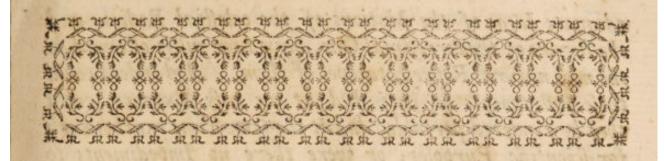
the Globes.

OUS dont les noms sont souscris, Maitres Mathematiciens dans & aux environs de Londres, ayont vû & examiné la Construction & l'Usage des Planisphéres Universelles, Inventées par Monsieur Dunn, les trouvons propres à répondre avec la plus grande facilité, & selon la generale manière, à tous les Problemes aux quels on peut répondre par les Globes, avec autant de certitude que de promptitude, & à cause de ce la nôtre Opinion est, quelle sera d'une grand Utilité au Public.

J. L. Cowley, Mathematicien à Ludgate-bill.
J. Welch, Precepteur pour les Navigations, Hatten-garden, Charles Brent, Autheur de l'Astronomie abrêgée.
Thomas Adams, Precepteur en Mathematiq. Soho-square.
Edward Moore, enseignant les Navigations, & faiseur de

Globes, Deptford.

Peter Russel, enseignant la Geographie, & l'Usage des Globes.



PREFACE.

类类类类 EEING that from a Variety of terrestrial and celestial Observations, it is certain that the S the Earth and Sea together do make one round Ball, and that the same is nearly globular or # A A fiberical, and that therefore the Sciences of Geography, Navigation and Hydrography, cannot be understood without a Knowledge of the Properties of a round Body or Globe. Seeing that this earthly Ball doth either turn round its own Axis from West to East every 24 Hours, or otherwise the celestial Bodies do move round it from East to West in the same Time; and that there is the highest Probability for the former of these Motions, notwithstanding Phænomena seem to contradict the same. Seeing also, that the fixed Stars, Sun and Planets are so far distant. from us, and their Motions under other Directions, are at all Times to exceedingly flow in Comparison with the apparent diurnal Motion, as to bear scarce any sensible Proportion therewith. Seeing also, that these Motions are apparently circular, and that Time or Duration itself is measured by them: From these Particulars, it naturally follows, that the Sciences of Astronomy, Chronology and Dialling, do principally depend upon the Lines and Properties of a Sphere.

Hence Globes, Spheres, Orreries, and other Machines, peculiarly adapted and designed for conveying distinct and adequate

adequate Ideas of the Motions of the celestial Bodies, and the Figure of the terrestrial One, have been justly esteemed as the most natural Representatives of them. And indeed could these Instruments be furnished at an easy Expence, could they be rendered portable, and at the same Time exact enough for scientific Purposes; or were there no Impediments to binder them from being of general and ordinary Use; nothing else of this Kind need be enquired after. But as the Acquisition of these Difficulties bath bitherto been thought impossible, and is likely to continue so; therefore various artificial Delineations and Representations of the Sphere on a Plane, have been introduced by Mathematicians, and made use of occasionally, some for conveying Ideas of the Lines and Properties of the Earth and Heavens, others for making of Operations and deducing Conclusions, as the Nature of the Subject bath been able to admit, all of which have been of Utility, and some of them more ready and useful than Models themselves.

Thus Geographers, Astronomers and Mathematicians in all Ages, have been able to exhibit their Thoughts concerning Bodies and Spaces of three Dimensions, by linear Delineations, Diagrams and Schemes on a Plane; different indeed from the Magnitudes themselves, yet so as to afford proper Ideas of all Positions and Dimensions whatsoever; to express large Parts of the Earth's Globe by Maps; to represent the Rumb-lines in Mercator's Chart by Right-lines; to delineate solar Eclipses geometrically; to exhibit Landscapes in a Piëture by means of optic Glasses; and to draw sceno-

graphically or in perspective.

But that Part of this Doctrine applicable to our present Purpose, and upon which, as a Basis, our Universal Planispheres are founded; is the stereographic Projection of the Sphere in Plano, or the Delineation and Measurement of the Circles of the Sphere on a Plane, as they would appear to an Eye placed in the Superficies of a Sphere, in that Point of the solsticial Colure which is at the Equator: The

Practice

Practice of which, feems to be almost coeval with the Doctrine of the Sphere itself; as being plentifully diffused throughout innumerable Treatifes of mathematical Literature; generally practifed by Men of indisputable Genius and Capacity; and universally received as the best intermediate Means, without the Globe itself, for attaining distinct Ideas of spherical Lines. Nor is it to be unobserved, that the Antients have been very peculiar in choofing the stereographic Method for describing these Lines; when at the same Time, if a mere arbitrary Delineation only would have anfivered the same End, innumerable other Ways much easier and readier for Practice might have been proposed. And berein we may distinguish the Judgment of the antient Geometricians, although Antiquity be filent on this Head. For supposing the Eye to be placed at a Point in the Surface of the Sphere, and projecting the Circles of the Sphere on a Plane, these Circles will also be Circles in the Projection, a pyramidical Solid is hereby formed, the geometrical Propofitions, relating to similar Triangles, are immediately applicable, a great Variety of curious and useful Theorems are readily investigated, and the whole System of spherical Geometry, becomes demonstrable by Euclidean Principles. Whereas in a mere arbitrary Description of the Sphere on a Plane, without Regard to any such Principle, not one of these Advantages can be expected, nor can any others be introduced in their Stead, no geometrical Comparison can be made, and every Enquiry and Determination will be vague and erroneous.

Hence our Universal Planispheres are ample Instruments for geographical, astronomical, and nautical Purposes, bave but sew hidden Sources of Error, but what is common to all Maps, and the Gores of Globes themselves, the Contraction and Expansion of Paper by drying and wetting, which although somewhat considerable upon a large Sheet, whereas the Gores of Globes, as supposed to be printed singly on smaller Pieces of Paper, may be thought to receive less

Errors hereby, yet we shall do the best in our Power to remove this Inconvenience; by carefully mounting such as pass

through our own Hands.

Whereas, if it be considered what unexact Instruments common Globes are, and how many more Sources of Inaccuracy they admit of; that if the Horizon should differ ever so little from a perfect Plane, if the Centre of the Globe should be ever so little out of the Plane of the Horizon, if the Axis Should not pass through two Points diametrically opposite to each other in the brass Meridian, or the two Extremities of the Axis should not be perpendicular to the Surface of the Globe; if there should be an Imperfection in the Figure of the Surface, or Errors in making or pasting on the Gores; if the Globe once happens to Slide on its Axis, as it frequently doth by changing the Poles, the Quadrant of Altitude be. not duly extended, or the Parallax of the Eye be not carefully avoided; in either of these Cases, and in many others of the like Nature, the Exactness is lost, the Determinations by them are quite uncertain, and it is easy to see that these Planispheres having none of these Imperfections, with Respect to the Accuracy in Johning Problems, may exceed the largest Globes usually made. Notwithstanding which Globes themselves are very useful, and may be said to be almost absolutely necessary for giving the Generality of Persons an Idea of the Lines and Properties of the Sphere, they are curious Ornaments for the Cabinets of the Curious, and shew the Motions of the Earth round its Axis, or the apparent Motion of the Heavens in a most adequate and exact Manner, but are subject to many Cafualties and Accidents on Account of their Spherical Form. Hence Mr. Moore's new Globe, 31 Inches Diameter, the Work of more than 7 Years ingenious Application, now publishing by Subscription, delineated in a most exact and elegant Manner, according to the latest geographical Discoveries, had like to have been accidentally broken to Pieces, and required much Time and Care for the Reparation: And our own 28 Inch celestial Globe,

Globe, in removing from London to Chelsea, received a Fracture of more than a Foot in Diameter. Nor can all the Accidents which are daily happening to these Machines be ennumerated, as is well known to Globe-makers, who sometimes take the Charge of sending their Goods to a great

Distance, and receive great Damages thereby.

In the first Edition of this Work, being convinced that the Method proposed of solving Problems by the Planispheres, was founded on mathematical Principles, we did not think ourselves obliged to enter into any Proof on that Head, but being told by some Friends, whose Opinion and Learning we respect, that we had not done ourselves and our Invention justice, and that it might be proper to consirm the Truth of the Method, by such an Authority, which would be particularly useful to those who were less versed in Science, we yielded to their Reasons; and shall here present the Reader with a brief Illustration of the said Principles, and without entering mathematically into the Nature of Projection, shall explain it in a Manner which seems most easy and familiar to Conception.

And here it is to be observed, that since the general Phanomena of the heavenly Bodies depend not upon their absolute Distances, but on their relative Positions to one another, and to certain imaginary Circles, called great Circles of the Sphere; therefore, for the greater Ease of Conception, and a readier Explanation of the Phanomena, Astronomers refer the Sun, Moon, Planets and Stars to the Surface of one common Sphere, which they call the celestial Sphere, having the Eye of the Observer for its Centre, or rather as being concentrical to the Globe of the Earth, the Magnitude of which is inconsiderable with respect to that of the imaginary

celestial Sphere.

In this Surface, the celestial Bodies are supposed to be placed, like so many golden Nails, in a Roof, and herein also their respective Motions, whether real or apparent, are conceived to be performed.

From

From hence, is immediately derived the practical Use of the artificial celestial Globe, it being a direct Representation of the imaginary celestial Sphere; as the practical Use of the artificial terrestrial Globe is derived from its being a direct

Representation of the true terrestrial Sphere.

It may seem therefore, at first Sight, as if nothing but artificial Globes could folve the Problems of the Sphere, fince nothing but a Globe can represent the Position and Distance of Points on a Globe, in their true and proper Relation to each other. Nevertheless, if the Spheres be projected, as they are in these Planispheres, according to a regular and mathematical Projection, it is certain, that the Distances and Positions of Places and Stars (though different from what they really obtain on the real Spheres) will bear such a certain and determined Relation to each other, as, with the Help of proper Rules, to enable any one to solve all the Problems of the Sphere, the same as he would do on the artificial Globes. And this is done principally in these Plani-Spheres, (and herein the Peculiarity of our Invention consists) by Means of a Projection of the great Circles, and Parallels of the Sphere, called the Projector; and by the Motion of the same, according to the Nature of the Problem, the Answer is brought out. For the Projector which lies above the Planisphere, being printed on a very fine Paper, and being oiled, the black Circles, Lines and Points in the Planisphere underneath, are clearly visible through the Projector, without any trouble or straining of the Eye. Of these black Circles and Lines on the Projector, some will represent the Horizon, Meridian, vertical Circles, &c. according to the Nature of the Problem to be folved.

The Planispheres, in most of the Problems, are rectified for use, in a similar Manner to the Rectification of the Globe, and the Place of the Sun, Moon or any Planet, or Comet may be pointed out, by a Mark or Patch, as is usual in solving Problems upon the Globes. The following Comparison, which occurs to us, may serve to shew the

Con-

Conformity between these Planispheres and common Globes, and that the first may solve the Problems as well as the latter. But we defire it may be taken merely as an Illustration, and not as an Inference to derive Honour to our Invention from so advantageous a Comparison which it by no Means deserves, and which the Inference is not designed to convey. Common Globes, lead the Artist directly to the Point in View; thus, in great Circle-failing, the Artist contemplates his Course in a great Circle upon the artificial Globe, as if it was on the real One; and these Plani-Spheres lead him to the same End, though more indirectly, yet as readily and easily; the Places indeed in the Planispheres are put out of their natural Order and Position, (and so they are in Mercator's Chart) yet, as long as this is done according to a certain and regular Principle, adapted to the respective Nature of each Construction, the Artist attains the same End, and perhaps with greater Exactness. Hence it is presumed, in Favour of these Plani-Spheres, that, though the Nature of the Subject doth not admit of their conveying Juch clear Ideas as the Globes do, yet, in Point of Practice and Exactness, they exceed even Globes of a much larger Diameter. The whole external Surface of a Sphere, being equal to the Area of four great Circles of the Sphere, and each Sphere in these Planispheres being projected only into two great Circles, it follows, that a Globe having the same Diameter with these Planispheres, which is somewhat more than 21 Inches, will contain double the Quantity of Surface to what the two Planispheres, representing the same Globe, do contain; thus the Surface of the Globe will contain 350 Superficial Inches, and the Planisphere of the same Diameter 175 superficial Inches. But, the Sides of similar Planes, being as the Square-roots of their Areas; and the Square-root of 350 (18.70) being to the Square-root of 175 (13.23) as 21.2 to 15. it follows, that these Planispheres ought, (cæteris paribus) on this Supposition, at a Mean, only to equal, in the Exactness of solving

ing Problems, Globes of 15 Inches Diameter. Again, seeing that the Degrees near the Centre of the Planisphere
(where, all great Circles being insensibly different from
Right-lines, the Distances and Positions of Places may be as
readily measured as on a Plane, without sensible Error) are
but half the Length of those in the Circumference, and those
Degrees in, and near the Circumference, are equally large
with those on a Globe of the same Diameter with the Planispheres; it follows, that, at a Mean, these Planispheres
will be equally exact with Globes of 16 Inches Diameter, and
but little inferior to common Globes of 17 Inches Diameter.

Furthermore, it is very well known, that in the stereographic Projection, which is that here made Use of, the Eye is Supposed to be placed in the Surface of the Sphere, looking at the Concavity of the opposite Hemisphere, and thereby, projects the faid Hemisphere upon the Plane of a great Circle, perpendicular to a Line drawn from the Eye to the Centre of the Sphere, and that this is done by the Intersection of Lines drawn from the Eye to each Point of the Hemisphere, and cutting the said great Circle's Plane. It is Evident, therefore, that every great Circle, passing through the Centre of the Projection, will be a Right-line, and therefore, if any Place be found in the Circumference of the Plani-Sphere, that Diameter of the Planisphere which intersects the Periphery, at the Distance of 90 Degrees from the said Point, on each Side, measured along the Circumference of the Planisphere, or that Diameter, which is perpendicular to a Line joining the said Point and the Centre, will be the rational Horizon of that Place in the Circumference, and every Point in this Line, representing the Horizon, will be 90 Degrees distant from the Point in the Circumference representing the Place. Hence from the Nature of the stereographic Projection, the greater and lesser Circles of the Sphere, are here shewn by circular Arcs, which, in any other Projection, would be Curves of a much more difficult Description; and a greater Equality of Proportion is also kept up thro' the

the whole, than in the Orthographic, Gnomonic, or any other Projection bitherto practifed. Therefore, seeing that a Place given in the terrestrial Planisphere, may be referred to the Circumference, along the same Parallel of Latitude, and a Place given in the celestial Planisphere, may be referred to the Circumference, along the same Parallel of Declination, the Place being referred to the Circumference, the Circumference will become the Meridian of the Place; the Sun, Moon, and every other heavenly Body, keeping in the same Parallel nearly for 24 Hours, will perform their apparent diurnal Revolutions perpendicularly over that Parallel of Latitude, which is equally distant from the terre-Strial Equator; as the Luminary or Star, at that Time, is distant from the celestial Equator, that Place on the Planisphere, to which the Star is Perpendicular, will be found, in some Part of that Parallel of Latitude, which answers to the Declination of the Star. When the Luminary comes over that Part of the Parallel of Latitude, which cuts the Circumference, then the Star is upon the Meridian, and to Places in greater Degrees of North Latitude is faid to be South; as the Star moves uniformly along the Parallel of Latitude, and comes to the Horizon, it Jets; the rest of its Course will be along the Parallel, below the Horizon, if it be a Star that ever sets: The Number of Degrees contained between that Part which cuts the Circumference when the Star fouths, and that which cuts the Horizon, will be the semidiurnal Arc or the borary Angle between the fouthing and rifing or fetting, and is turned into Time at the Rate of 15 Degrees to an Hour for the Sun, 360 Degrees to 23 Hours 56 Minutes for the Star, and 360 Degrees to the Interval of Time between the Moon's fouthing from one Day to another. The seminosturnal Arc may be found in like Manner, by converting the Remainder of the Parallel below the Horizon, into Time, or elfe by subtracting the semidiurnal Arc, from half the Time of diurnal Revolution of the Star. When the first Point of Aries

is perpendicular to the Centre of Projection, (which always happens in 24 Hours) the Ecliptic will be projected into a Right-line as it is marked on the terrestrial Planispheres; whence, the Sun's Declination, and Right-ascension, are found, having his Longitude given, and also any one of the three given, the other two are found. The next Thing which may seem to require Explanation, is the Projection of the Meridians and Parallels, which are dotted off to Degrees on the Projector, also, the Manner and Reason of its solving the Problems.

As the Planispheres are stereographical Projections of the Places or Stars on the terrestrial and celestial Spheres; so the Projector is a stereographic Projection of the Meridians and Parallels of the Sphere to the same Radius; therefore, when the Points N. and S, or the North and South-poles on the Planisphere, are sitted to the two Poles of the Projector, the Meridians of the Projector represent Circles of Latitude or Circles of Declination, and the Parallels represent Parallels of Latitude or Parallels of Declination, according as the terrestrial and celestial Planisphere is used. But, as it is by the Motion of the Projector, round the Centre, over the Planisphere, that most of the Problems are performed (which Motion answers generally to the setting of the Globe to the Latitude of the Place) it may be proper to show how this is done.

It has been said before, that the Meridians and Parallels of the Projector, when the Points Z and N are brought to the Poles of the Planisphere, do represent the Meridians and Parallels of the Sphere; and then all Places or Stars, laying under the same Meridian of the Projector, are found upon one and the same great Circle upon the Globe. But if the Points Z and N are put upon other Points of the Circumference, then, for the same Reason, the Meridians of the Projector, will still represent great Circles of the Sphere, and all Places, found under the same Meridian of the Projector, do lie in the same great Circle upon the Globe, which

which we may easily apprehend, if we consider, that, all great Circles, which pass at an equal Distance, all round, from that Point which is the Centre of the Projection, will be alike projected, all round upon the Plane of the Projection at equal Distances from the Centre of the Projection. Hence, an easy Solution is given, of that Problem, which seems most likely to embarass us, in a Planisphere, which is, to find the Distance of two given Places, or two given Stars from one another, in the Arch of a great Circle, for it is but turning the Projector about, 'till one of the Meridians of the Projector, or at least one row of its Dots, either falls over the two given Points, or at least is parallel to them, and the Distance in Degrees, upon the Projector, between the two Points, will be the Distance required. To give another similar Instance, let it be required to find either the Zenithdistance, or the Altitude of any celestial Phænomenon from the Declination and Time of fouthing of the Phænomenon given. Find the Point upon the Parallel of Latitude, anfivering to its Declination at the given Time, and move the Slider round, 'till the Pole of the Planisphere is turned from the Horizon of the Projector, fo many Degrees as is the Latitude of the Place of Observation; then the Horizon of the Projector is the rational Horizon, and the Number of Degrees contained upon a Meridian of the Projector, between the Place of the Star on the Parallel of Latitude and the Pole of the Projector, is the Zenith-distance of the Star, as the Number of Degrees between the Place of the Star on the Parallel, and the Horizon, measured upon the same Meridian, passing through the Place of the Star, is the Star's Altitude.

It may, perhaps, conduce to the better Understanding of the Nature and Grounds of the working of the Problems, to shew the direct Analogy between the Manner of performing them by the Planispheres, and the Globe. And, this may be done by supposing a Semi-globe or Sphere, which shall here answer to the Projector, with all its Circles and Lines

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drawn thereon, to be placed upon an horizontal Plane, and at the same Time the Eye of an Observer, or luminous Point, to be placed at the Distance of a Semi-diameter, above the Centre, in the Perpendicular to that Plane; then, it is certain, that all the Lines, which naturally subfift on the Surface of the Semi-globe or Sphere, will be artificially represented on the Plane, and whatever is formed in the One will also, in a perspective Manner, be formed in the other. Now, if we suppose another Semi-globe, having the Delineations, whether of Lands and Seas, or Stars, and which shall here answer to the Planisphere placed under the Projector, and the Eye of the Observer, or luminous Point, also the Superficies of the inward Semi-globe to be at rest, whilst the Cap turns horizontally round its Centre, when the Cap is thus in Motion, the same Parts of the Cap or moveable Superficies, will always coincide with the Same Parts of the immoveable Superficies, alike on the Globe as on the Projection respectively, and hence, the Rotation of the Cap will be represented by that of the Plane. Therefore, seeing that the Problems of the Globes, do require no more than such a Revolution, the Use of the Globes may be easily answered by a Plane. Hence, as in most of the Problems of the Globe, the Place is brought to the Zenith and to the Horizon, by elevating the Pole to the Latitude of the Place, and by bringing it to the Meridian, so in like Manner, in the Planispheres, the Pole in the Circumference or Meridian of the Projector, is brought to the Zenith on the Planisphere, by which Means, the Horizon of the Projector, becomes the rational Horizon, and answers to the wooden Horizon of the Globe, and the Pole becomes elevated above the Horizon of the Projector, answerably to the Latitude of the Place. So that the Planisphere, in these Circumstances is an exact projected Representation of the Globe, in its usual Position, when it is set to the Latitude of the Place, and the Place brought to the Meridian. For if the

the Eye was placed in the western cardinal Point of the wooden Horizon of the Globe, and so as to touch the same, then the Eye, being in the Plane both of the Equator and Horizon, they would both be projected into Right-lines, as they are found upon the Planisphere; the Places also being in the Zenith and in all Parts of the Meridian, would be seen in the Circumference of the Projection, and therefore they are placed in the Circumference of the Planisphere, so that we may refer them thither by the Parallels of Latitude, which answers to bringing the Place to the Meridian on the Globe, by the Motion of the Globe round its Axis. And thus would the Eastern Planisphere, with all its vertical Circles and Almicanthars, be viewed in Perspective, and

would appear like the Planisphere with the Projector.

Indeed, the Planispheres are not projected to answer the outward, but inward perspective View of the Globe; for as the Concavity of the opposite Hemisphere is seen by the Eye in Perspective, every Thing on the Planisphere takes a contrary Situation to what it would obtain, was the Eye to view the external Superficies of the Globe; therefore, tho' the Perspective is reversed in Planispheres, with the relative Politions of Places to one another, yet the Places are always set right again in their proper Situations, as if the Observer was to view the external Surface of the Globe. Otherwise the Eastern Planisphere is projected, to answer the perspective View which the Eye would have of the Eastern Hemisphere, if the Eye was placed in the eastern cardinal Point, touching the Globe, and the Eastern Hemisphere was depressed, so as to depart from the Eye, and penetrate the Solidity of the Globe, and its Convexity should yield inward, and be turned into a like Concavity, and be applied or fitted to the Concavity of the opposite Western Hemisphere, the perspective View, which the Eye still retaining its Place, would have, of the Surface of the Eastern Hemisphere, thus transferred, would answer exactly to the ProProjection of the Eastern Planisphere, when set to the La-

titude of the Place by Means of the Projector.

Having in the foregoing Explanation, endeavoured to give the Reader a general Idea of the Nature of the Plani-Ipheres, and of the Manner in which the Problems may be folved by them, as well as of the Principles on which the Justness of the Solution is founded, we shall now introduce bim to another and very different Method of folving Problems by the Planifpheres, from what was delivered in the first Edition of the Precepts, or than has been yet hinted at in the foregoing Pages; which will also be exemplified in the following Precepts; a Method which we flatter ourfelves will be looked upon as very extensive in Use, ready in Practice, and as conveying Ideas not much inferior to what the Globes convey, either in Clearness or Precission: (For in Point of Exactness of the Solution of Problems, it has been allowed by the judicious and impartial that the Planispheres exceed even Globes of a much larger Diameter.)

Hitherto, we have taken the Liberty to suppose the Place of the Observer in each Operation to be transferred to the Meridian, or the Circumference of the Planispheres, and by this Method, all the Circumstances, depending on the diurnal Motion, are readily and conveniently folved, when the Declination and Time of fouthing of the Planet or Star are given. But, as the Artist may think it a more natural and universal Method to solve the Problems without removing the Places on the terrestrial, or the Stars on the celestial Planispheres, from their Situation as there laid down; and, as in many Cases, it may be useful and agreeable, to see at one View, the relative Position of the Stars, to different Places of the Earth at the same Time, or the relative Positions of various Stars at the Same Time both to one another, and to the Horizon and Verticals of the Place in the same Manner as is done by the Globes, it seemed that it would be no inconsiderable Addition to, and Improvement of our Invention,

vention, if we could investigate such a Method and present it to our Readers.

Upon a little Consideration, and an Attention to the Nature of some of the principal Problems, which might be proposed to be solved, this seemed more easy to us than we could bave expected, and so obvious when found out, that we were almost surprized that we had not fell upon it before. This Method, we shall briefly point out, in the Cafe where the celestial Planispheres are used, where indeed it is of the greatest Use, though it also obtains in the Use tf the terrestrial Planispheres. This depends then chiefly, upon finding what we call the Place of the Zenith upon the celestial Planisphere, or that Part of the celestial Planisphere, the corresponding Point to which in the Heavens is, at that Time, in the Zenith of the given Place. The given Place of the Observer, is supposed to be projected upon the celestial Plani-Sphere, or to be referred to that Parallel of Declination, which answers to the Parallel of Latitude of the given Place. The given Place, or, if you please, the Zenith, or Place of the Zenith, is conceived to move uniformly along its Parallel of Declination, every Day, from West to East, in like Manner as any Point upon the Earth's Surface, by its diurnal Rotation, really describes a parallel Circle to the Equator, or as the artificial Globe represents the same.

To find the Place of the Zenith at Noon, find the Sun's Place in the Ecliptic, on the Planisphere, and that Point of the Parallel of Declination, which lies under the same Meridian with the Sun, is the Place of the Zenith at Noon. If the given Time be before Noon, remove the Place of the Zenith along the same Parallel so many Degrees, to the East of its Place at Noon, as answer to the apparent diurnal Motion of the Heavens in the given Number of Hours, which is at the Rate of 15 Degrees to each Hour. If the given Time be after Noon, remove the Place in like Manner to the Westward of its Place at Noon. The Place of the Zenith at the given Time, being thus found, a compleat Re-

Representation of the apparent Situation of the Heavens, is exhibited: For the Place of the Zenith being found, relative to the celestial Sphere, and consequently the Meridian of the same, or the Meridian passing through the Place of its Zenith; it is evident, that the Sun and all the Stars, are found in their true relative Position, with respect to the Zenith, no less than they are with respect to each other. Any Meridian of the Projector passing through the Place of the Zenith, is a vertical Circle, and the Meridian of the Projector, which is every Way 90 Degrees distant from the Zenith, is the Horizon. To find the vertical Circles, turn the Planisphere round the Centre, and all the Meridians of the Projector, which successively pass over the Zenith, are the vertical Circles. Whence, it is evident, that if the Zenith and any Star, be both brought under the same Meridian of the Projector, that Meridian will be the Vertical passing through the Star, and the Number of Degrees intercepted thereon, between the Zenith and the Star is the Star's Zenith Distance, the Compliment of which, to 90, is the Star's Altitude. The Horizon is found by turning round the Projector, 'till the Zenith comes under that Diameter of the Projector which is called its Horizon (though in this Case it doth not answer to that Name) then count 90 along the Horizon of the Projector, and the Meridian passing through the 90th Degree, will be the Horizon; for this Meridian will be at the Poles 90 Degrees distant from the Zenith, because the Zenith is now in the Horizon of the Projector, and every Point of this Horizon is 90 Degrees from the Poles: And a great Circle which is 90 Degrees distant from any Point in three Places, will be a great Circle, having that Point for its Pole. Here, all Stars upon this Horizon, will be either rifing or setting; if it be the eastern Part of the Horizon, they will be rifing; if it be the western Horizon, they will be setting; all Stars lying between the Horizon, or nearer the Zenith than the Horizon, will be above the Horizon; and and all Stars lying without the Horizon, will be below the Horizon.

Thus the Horizon and vertical Circles may be found and shewn by Inspection, and the Problems depending upon them, clearly and readily solved, for the Performance of which, we have given ample Instructions in the following Pre-

cepts.

Furthermore, for making this second Edition more compleat, we have given a Description of the folar System, as it is well known to be constituted, from the Discoveries which have been made by the best Philosophers and Astronomers; which we hope will not be unacceptable, it being introduced in a Short and Summary Way, for the Information of such as have neither Opportunities nor Dispositions for making deeper Enquiries into the natural Causes of those Phænomena. To have omitted an Article of this Nature in this Treatise, we think would have been improper, for notwithstanding scarce any thing but the apparent diurnal Motion can be represented by the Globes, the Obliquity and other Properties of the Planet's Orbits being of a more astronomical and physical Nature, and not to be represented by circular Machines, yet as most good Treatises of the Globes, do contain something of this Kind, we have complied with Custom in this Part of our Work.

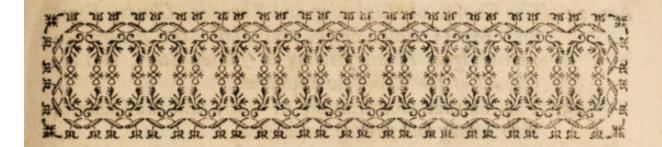
For these Reasons we have also given a short Account of the Principles of Geography, together with several Tables for understanding the most useful Particulars relating to the principal Parts of the known World. Also other Tables of a more scientific Nature, for the Purposes of Calculation and Inspection, Use and Amusement. The vacant Space left after each Problem, is to be filled up by the Practitioner, with such Examples as himself or his Preceptor shall choose to set, in the doing of which, we recommend a careful and regular Arrangment of the Data and Quæsita, by the Help of slight penciled Lines, as the Book when so filled out, will not only contain a great Variety of Problems, with their Solutions

by the Pupil himself, but will then be a more valuable Repository than at first. The blank Column in the Table of the
Sun's Place, may be filled out with such remarkable Days of
the Year as each respective Reader shall choose to insert.
Should the small Piece of Linnen, Silk, or Vellum, at the
Centre of the Slider get loose, it may be fastened by melted
Glue.

Several other Uses may be made of these Planispheres, beside what we have introduced in the following Treatife, and which we presume our more ingenious Readers will readily discover; fuch as the stereographic Projection on the Plane of the Meridian for any Radius less than that of the Planispheres, by the Help of a Pair of proportional Compasses, whereby a great Variety of Cases in Spherical Geometry will be reduced to almost a single Process; even the orthographical Projection itself may be made hereby, and all the Parts of an Ellipsis truly delineated and measured, the Paths of the Planets may be described among the fixed Stars, and the Appearances of Eclipses at different Places on the Earth's Globe may be known, the Discussion of all which would be unnecessary; wherefore I shall here conclude this Preface, and use my best Endeavours for the Satisfaction of such Persons as choose to apply to me for these Things.

The Academy at Chelsea, Dec. 1, 1758.

S. Dunn.



THE

CONTENTS.

Art. 1. THE Order of this Treatise.
2. The Universe defined.

3. The ultimate Defign of natural Philosophy.

4. The Laws of Nature defined.

Gravity defined.Elasticity defined.

Margin. Sir Isaa'c Newton's Conjecture concerning the Creation of the World.

7. A probable Reason, that the fixed Stars may be so many Suns.

8. The Planetary fiftem defined.

9. The Sun described.

10. The folar Maculæ described.

11. The Orbits of the Planets described.
12. The several Parts of an Orbit defined.

13. The Earth's Orbit defined.

14. The Motions of the Planets defined.

15. The Laws of the Planet's Motions in their Orbits.

16. The apparent Motions of the Planets defined.

17. Farther Definitions concerning the Planet's Orbits.

18. The Comet's Orbits defined.

19. The Laws of the secundary Planets defined.

20. The Planet Mercury described. 21. The Planet Venus described.

22. The annual and diurnal Motion of the Earth, and the Cause of the Vicissitudes of the Seasons explained.

23. The Nature and Properties of the Earth's Atmosphere.

explained.

Art. 24.

Art. 24. The Causes of Lightning and Thunder, Clouds, Fogs, and Mists; Rain, Snow, and Hail; the Aurora Borealis, Ignes Fatui, shooting or falling Stars, and the Colours of the Rain-bow explained.

25. The natural Constitution of Seas, and the Cause of

their ebbing and flowing explained.

26. The Causes of accidental, constant, and periodical Winds, Sea and Land Breezes explained.

27. The Origin of Rivers, Rivulets, Springs, and mineral

Waters explained.

28. The different Kinds of Earths, Fossils, and Stones explained.

29. Salts, Minerals, and Metals explained.

30. Conjectures concerning the internal Srusture of the Earth.

Margin. Dr. Halley's Opinion concerning the internal Constitution of the Earth's Globe.

31. Vegitables and Planets described, and the Manner how

they are produced.

32. Animals confidered according to their Classes, and the Manner how they are sustained.

33. Sensation, and the Powers of the Mind explained.

Margin. A short and comprehensive Account of the Operations of Nature, from Sir Isaac Newton.

34. The Moon described.

35. The Planet Mars described,

36. The Planet Inpiter, and his Satellites described.

37. The Planet Saturn, and his Satellites described.

38. Comets described.

1 2 July

39. An Account of the fixed Stars.

40. A Table, shewing the greatest, mean, and least Distances of the primary Planets from the Sun, with their Eccentricities in Miles.

41. A Table, shewing the Diameters of the Sun and Planets in Miles, and also the Surface, Magnitude and Quantity of Matter in each, the Surface and Magnitude of our Earth being 1, or Unity.

42. A Table, shewing bow many Miles the Planets move in their Orbits, per Day, Hour and Minute, at a Mean.

Art. 43. A Table, shewing the Planets daily mean Motions, with the Number of Miles they are drawn from their Restilinear Directions per Hour, their Densities, Light and Heat, compared with our Earth.

44. A Table, shewing the Times in which the Planets move

round the Sun, and also round their own Axes.

45. A Table of the Inclination of the Planet's Orbits, to the Plane of the Ecliptic, and the Inclination of their Axes to their own Orbits.

46. A Table of the Sun and Planet's greatest and least ap-

parent Diameters, as viewed from the Earth.

47. A Table, shewing the Places of the Planet's Notes and

Apheliums.

48. A Table, shewing the mean Distances of Jupiter's Satellites from his Centre, in Diameters of his Body, and the Times of their Revolutions round him.

49. A Table, shewing the Weights of equal Quantities of

Matter, on the Surface of each Planet.

50. A Table, shewing the mean Distances of Saturn's Sateltellites from his Centre, in Diameters of his Ring; and the Times of their Revolutions.

51. A Table, sherving in how long Time each of the Planets

are conjoined with the Sun.

52. A Table, shewing how many Degrees the Sun is usually under the Horizon, before the Planets and Stars appear.

53. A Table, shewing the Increase of Heat toward the Equinostial Line, and its Decrease toward the Poles, when the Sun is in the Equator; according to the Obliquity of the solar Rays.

54. A Table, shewing when Venus and Mercury may be seen in the Sun's Disque, and at what Distance from his

Centre.

55, 56, 57, 58, 59, 60. Geographical Definitions.

61, 62, 63, 64, 65, 66, 67. 68, 69. Astronomical Definitions.

70. A Table, shewing how many Days must be added to the

Julian Account for reducing it to the Gregorian.

Afcension, Declination, and Equation of Time for every Day in the Bissextile-year, 1760, also his apparent Diameter, hourly Motion, Time in transiting the Meridian, and Distance from the Earth in Miles, every tenth Day.

Art. 72.

Art. 72. A Table, shewing the Variation of the Sun's Declination to every Hour, or 15 Degrees of Longitude from the Meridian of London, his daily Variation being known.

73, 74. A Table, shewing the Dip of Horizon to 50 Feet Height above the Surface of the Earth and Sea; also the Refraction of the Atmosphere from the Horizon to the Zenith, for the Parallel of London and the Torrid Zone.

75. A Table, shewing the Number of Days from any Day in any Month, to the same Day in any other Month.

76. A Table, shewing the Gregorian Sunday Letters.

77. A Table of the fixed Holy-days, and remarkable Days throughout the Year.

78. A Table of New and Full Moons.

79. A Table, shewing how many Hours and Minutes the Moon is past the Meridian, when it is High-Water at the following Places, viz.

80. A Table, shewing the Right Ascensions, Declinations and Magnitudes of the principal fixed Stars, calculated

for the Year 1760.

81. A Table, shewing how many Hours and Minutes of Time, are equal to any Number of Degrees of Right Ascension.

82. A Table, shewing how many Degrees of Right Ascension, are equal to any Number of Hours or Minutes in

Time.

83. A Table, shewing the Sun's Right Ascension and Decli-

nation for the three first Signs, vo, &, II.

84. A Table, shewing the Latitudes and Difference of Meridians from the Royal Observatory at Greenwich, of

the following Places, viz.

85. A Table of Horizontal Distances for Horizontal Dials, shewing also the Distance of each Hour-line from the Meridian upon direct North or South Planes; calculated to every Degree of Latitude.

86. A Table of Angles which every Rhomb (or Point of the

Compass) makes with the Meridian.

87. A Table, shewing what Declination the Sun must have, when the Day is any Number of whole Hours longer or shorter than 12 Hours in any Latitude.

88. A Traverse-Table.

Art. 89. A Table, shewing the Comets Nodes, their Perihelia, and Log. of their Perihelion Distances from the Sun; the Radius of the Earth's Orbit being 1, &c.

Margin. Useful Observations concerning the aforegoing Tables, and also for restoring the Transparency of the

Slider.

90, 91, 92, 93. A Description of the Eastern terrestrial Planisphere.

94, 95, 96. A Description of the Eastern celestial Plani-

Sphere.

97. A Description of the Western terrestrial Planisphere.

98. A Description of the Western celestial Planis here.

99. The Principles of Geography.

100. An Estimate of the Magnitudes of Countries compared with each other.

101. An Estimate of the Number of Inhabitants in the follow-

ing Countries.

ployed in Time of Peace in the following Countries.

these of War and Commerce, belonging to the following Countries.

Par of Exchange, and the real Species of the Chief trading Countries of Europe.

105. A Table, shewing the Par of Exchange with the Chief

trading Places of Europe.

106. A Table, shewing the foreign Gold and Silver Coins.

107. A Table of the Climates from the Equator to the Polar Circles.

108. A Table of the Equality of Weights, at the most considerable Places in Europe.

109. A Table of the Equality of Long Measures, at the most

considerable Places in Europe.

110. A Table of English Coins, Weights and Measures.

111. A Table of the GENERAL POST.

xxiv		The	CONTENTS.
Art.	112.	Prob. I.	To find the Latitude of a Place given.
	113.	2.	To find the Difference of Latitude between
	3-		two Places given.
-	114.	3.	To find the Longitude of a Place given.
	115.	4.	To find the Longitude of a Place when
	-		the Meridian of London is the first Meri-
The same			dian.
	116.	5.	To find the Difference of Longitude be-
-ital		the water	tween two Places.
	117.	6.	To find a Place whose Longitude and La-
		CONTRACT TO ST	To find what o'Clock it is at one given
	118.	7.	Place, when it is a given Hour at ano-
		-	ther given Place.
. 500,000	710	8	To find the Number of square Miles in a
	119.	Of Albert	Country.
-	120.	g.	To find how many Miles go to a Degree
+300%	A con	d Francis	of Longitude in any Parallel of Latitude.
	121.	10.	To find how many Degrees of Longitude
			go to a Degree of Latitude or Distance
700		cing to the	in any Parallel.
	122.	11.	To find how many Miles an Hour a Per-
		tot almos	fon moves in any Parallel of Latitude,
	A POR	20 201	by the Motion of the Earth round its
			Axis. To find the Autori of a given Place
	123.		To find the Antaci of a given Place. To find the Periaci of a given Place.
	124.		To find the Antipodes of a given Place.
	125.		To find what Climate a given Place is in.
	127.	-	Given the Month and Day to find the Sun's
· Trins	4-7.	Same of	Place in the Ecliptic.
	128.	17.	Given the Month and Day to find the Sun's
	301 2	s similar	Right Afcention.
	129.	18.	Given the Month and Day to find the Sun's
	Acres		Declination.
	130.	19.	To find all those Places where the Sun will
	-		be vertical on any Day at Noon.
	131.	20.	To find where the Sun is vertical on a given
	A Elm		Day, and at a given Hour.
	132.	21.	To find how many Degrees the Sun will be
			North or South of the Zenith of a given
			Place, on a given Day.
-			Art. 133.

2150	CONTENTS. XXV
Art. 133. Prob. 22.	To find the longest Day and shortest Night
	at a given Place.
134. 23.	To find what Hour the Sun will rife or fet
	at any given Place on a given Day.
135. 24.	To find the Length of the Day or Night,
dependence and a	at a given Place on a given Day.
136. 25.	To find what Day the Sun begins to appear
.30.	above the Horizon, at a given Place in
and the Paris of the last of	the Frigid Zone.
137. 20.	To find on what Point of the Horizon the
	Sun rifes or fets at a given Place, and
pur firms and an	on a given Day.
138. 27.	To find that Part of the Equation of Time
Emoxibel and low	which depends upon the Sun's Place in
our by the Plant	the Ecliptic, and his Right Ascension.
120 28.	To find the whole Equation of Time.
	To find what Hour a Planet, Comet, or
140. 29.	
	New Star will rife or fet; its Right Af-
	cension and Declination, and the Lati-
	tude of the Place being given.
141. 30.	To find all those Places where a Comet
	will be vertical, its Declination being
hes little lost down	given.
142. 31.	To find the Distance of two Places on the
	eastern and western Planispheres.
* 12 02	To find through what Countries a great
143. 32.	Circle no Cuth in leading from an air
	Circle paffeth, in leading from one gi-
	ven Place on the Planisphere, to ano-
	ther given Place on the same.
144. 33.	To find the Distance of two Places, one of
Con need distance in	which is in the Eastern, and the other
	in the western Planisphere.
145. 34.	To find those Parts of the Earth and Sea
-13. 31.	over which a great Circle will pass, in
	leading from one Place on one Plani-
	fphere, to another Place on the other
	Planifphere.
146. 35.	To find the Position of two Places with the
	Meridian, when both Places are on one
	Planifphere.

xxxi	The	CONTENTS.
Art.	147. Prob. 34.	To find the Position of Places with the Meridian, when one Place lyeth in one Planisphere, and the other Place in the other Planisphere.
	148. 37-	To find where a great Circle, passing over the Earth's Surface, will come nearest to the Pole.
		To draw a Meridian, or North and South Line at any Place on the Earth's Globe.
	150. 39.	To find the Variation of the Magnetic Needle.
	151. 40.	To find a Meridian Line by the rifing and fetting of the Sun.
	August August	To find on what Point of the Horizon a terrestrial Object bears by the Planifiphere.
10-,1	153. 42.	To find all those Places of the Earth where the Sun is either rising or setting, at a given Time.
200 m	154. 43.	To find those Places of the Earth, which are in the enlightened Hemisphere, at a given Time.
adi a	155. 44.	Sea, where a folar Eclipse can be seen, the Day and Hour when it will happen being given.
	156. 45.	To find all those Places of the Earth and Sea, where a lunar Eclipse may be seen the Day and Hour when it will happen being given.
ane d		To find all those Places which have two Summers and two Winters in one Year.
		To find the Latitude of a Place of the Earth
		. To find the Longitude of a Place on the Earth or Sea.
*	160. 49	of the Place, to find the Hour of Begin ning and End of Twilight.
	161. 50	Day of the Year the Twilight is shorted at that Place.

Art. 162. Prob. 51. Given the Day of the Month, Latitude of the Place, and Altitude of the Sun, to find the Hour of the Day.

163. 52. Given the Latitude of the Place, the Declination and Altitude of the Sun, to find the Azimuth.

ference between the Altitude of the Sun taken at one Time of the Day, and the Altitude taken at another Time of the fame Day; also the Difference between the two correspondent Azimuths taken at the same Times the Altitudes were taken, to find the Latitude of the Place.

165. 54. To find the Latitude of a Star given on the Planisphere.

166. 55. To find the Longitude of a Star given on the Planisphere.

167. 56. Given the Latitude and Longitude of a Comet, Planet, or Star, to find its Right Afcension and Declination.

168. 57. To find the Distance of two Stars that are both on one Planisphere.

which is on the Eastern, and the other on the Western celestial Planisphere.

170. 59. Given the Direction of any celestial Phenomenon passing through any Part of the Heavens in the Arch of a great Circle, to find the Path of the same.

Latitude, Longitude, Right Ascension and Declination, for any Year past, prefent, or to come.

Hour a Star, Planet, or other celestial Body, whose Place in the Heavens is known, will come to the Meridian.

Day and Hour, to find what Point of the Ecliptic is on the Meridian.

The CONTENTS. XXVIII Art. 173. Prob. 63. Given the Latitude of the Place, and Hour of the Day, to find the Altitude of Medium Cali. 64. To find the oblique Ascension of any cele-175. stial Body. 65. To find the oblique Descension of any ce-176. leftial Body. 66. To find the cosmical Rising and Setting of 177. a given Star. 67. To find the anchronical Rifing and Setting 178. of a given Star. 68. Given the three Sides of spherical Trian-179. gle, to find the Angles by the Planiiphere. 69. Given the three Angles of a spherical Tri-180. angle, to find the three Sides by the Planisphere. 70. Given two Sides of a spherical Triangle, 181. and the Angle between them, to find the other Side and two Angles. 71. Given two Angles of a spherical Triangle, 182. and the Side between them, to find the two other Sides and Angle. 72. Given two Sides of a spherical Triangle, 183. and an Angle opposite to one of those two Sides, to find the other Side and two Angles. 73. Given two Angles of a spherical Triangle, 184. and a Side opposite to one of them, to find the other Sides and Angle. 185.

74. To know the Latitude of the Moon, Planets, Comets, or other celeftial Bodies.

75. To know the Longitude of the Moon, Planets, Comets, or other celeftial Bodies.

186.

187-

76. Given the Bearing of the Sun by the magnetic Needle (commonly called the magnetic Azimuth) the Latitude of the Place and Sun's Declination, to find the Va-Mation of the Compais.

Art. 188. Prob.	77.	Given the Latitude of a Place in the tor- rid Zone, to find the Sun's greatest Azi-
		muth on a given Day, when the Sun is nearer the Pole than the Place given;
		also the Altitude of the Sun at the same
		Time.
189.	78.	Given the Sun's Declination, Altitude, and greatest Azimuth, to find the Latitude.
190.	79.	To draw an horizontal Dial for any Lati-
Carlo Shell on Ann	,	tude by the Planisphere.
191.	80.	To draw an upright fouth or north Dial, by the Planisphere.
192.	81.	To draw a declining or reclining Dial, by
Good the shows Va		the Planisphere.
de de maria		To draw Hour-lines on any Surface what- foever.
194.	83.	Having the Latitude of your Habitation,
distrigues a stone	grik	and Meridian-line, to determine the Hour of the Day by the Sun.
195.	84.	Required the following Answers, for the Latitude of 50° North, on May 16.
196.	85.	Required the following Answers, for the
A Charles Santon	1000	Latitude of 50° North, on November 23.
197.	86.	Required the following Answers, for the Latitude of 35° South, on November 10.
198.	87.	Required the following Answers, for the Latitude of 45° South, on July 29.
199.	88.	To find if it be Leap-year, or how many Years after.
200.	89.	To find the Sunday Letter for a given Year, from 1800 to 1899 inclusive.
201.	90.	To find the Golden-number for a given Year of our Lord.
202.	91.	To find the Cycle of the Sun for a given Year.
203.	92.	To find the Roman Indiction.
204.		To find the Epact, New-style, for a given
101 1 2 2 101		Year.
205.	94.	To find the Age of the Moon.

Art. 206. Prob. 95. To find the Time of the Moon's fouthing. 207. 96. To find the Time of High-water at a given Place.

208. 97. To find Easter-day.

209. 98. To conduct a Ship from one given Place to another, along the Arch of a great Circle.

Place at Sea, where the Variation Lines are nearly North and South, to find the Variation of the magnetic Needle at that Place for a given Year.

the Longitude where the Variation Lines

run nearly North and South.

two Places, whose Difference of Latitude and Difference of Longitude are given, to find the Latitude and Longitude of each of those Places.

Margin. A new and useful Hint concerning the Angle of the Sun's Position, to the Pole and Zenith.

ERRORS of the Press, which the Reader is requested to correct with the Pen. Note, p. signifies Page, l. Line, f. for, r. read.

PAGE 1. l. 26. f. or forward, r. or Motion forward; p. 5. l. 3 and 4. f. Apsides, r. Apsis; l. 38. f. always inversely, r. always nearly inversely; l. 40. f. always as, r. always inversely as; p. 6. l. 18 and 19. f. the Planet, r. the Distance of the projected Place of the Planet on the Ecliptic from the Earth; p. 7. l. 22. f. little more dense, r. little less dense; l. 29. f. of that, r. of the Earth and that; p. 27. l. 25. f. 40,000, r. 400,000; p. 29. l. 14 and 27. f. Axis, r. Axes; p. 32. l. 26. f. June 10, r. June 12; l. 27. f. June 8, r. June 10; l. 28. f. Dec. 13, r. Dec. 16; l. 29. f. Dec. 11, r. Dec. 14; p. 55. l. 1. f. 73, r. 73 and 74; p. 58. l. 37. f. Saxony, r. Saxons; p. 59. l. 11. after Westminster, r. when 9th is a Sunday; (l. 13. after 10th in some Copies, r. when 9th is a Sunday;) p. 87. l. 15. f. Scytian, r. Scythian; l. 8. f. valient, r. valiant; p. 89. l. 6. under Religion for Ditto and English, r. Ditto; l. 6. under Manners for Ditto and English, r. Ditto; p. 142. l. 16. p. 143. l. 20. and p. 144. l. 6. f. Gnomen, r. Gnomon.

As the Limits to which we have been necessarily confined for rendering the Contents of this Treatise within the Compass of an ordinary Pocket Volume, would not admit of being so ample on several important Problems, as we could have been, of which, Problem 48, may be reckoned among those of greater Utility, we hope it will not be unacceptable to our Readers, if we particularize the same by Way of Specimen here.

PROBLEM.

To find the Longitude of a Place, on Land, or at Sea, by the celestial Bodies?

OBSERVATION.

Of all the Methods, that have been hitherto proposed, for finding the Longitude, either on Land or at Sea, by the celestial Bodies, that by the Eclipses of Jupiter's Satellites, hath been generally found to be the best for Land, and that by the Ap-

pulses of the Moon to the fixed Stars, best for Sea.

Now if it be known by Calculation, that an Eclipse of one of Jupiter's Satellites, or an Appulse of the Moon to a fixed Star, will happen at a certain Hour and Minute under a known Meridian, and the same Eclipse or Appulse, is observed under another Meridian, the Difference between the computed and observed Time, being converted into Degrees and Minutes, is the Difference of Longitude between the two Meridians. But farther, supposing a Table to be exactly calculated, shewing the Times when fuch Eclipses or Appulses will happen under a known Meridian, it will still be necessary to know nearly the Times when fuch Eclipses or Appulses are to be expected, at a Place whose Longitude is either known pretty nearly, or not at all. Now that a Pair of Globes, and therefore our Planispheres, are exceedingly useful for these Purposes, will appear from the following Confiderations. 1. Because the Right-ascension and Declination of the Planet (whether it be Jupiter or the Moon) is known in Degrees and Minutes, and confequently the Place among the fixed Stars, at the Time of the Eclipse or Appulse. And because the Observer may take the Latitude he is in, and thereby find the Time of the Planet's Continuance above the Horizon; therefore, if the Longitude be pretty nearly known, by folving the Problem by the Planispheres, they will shew whether it is above or under the Horizon at the Hour and Minute when fuch Eclipse or Appulse happens. 2. If the Observer is quite unacquainted with the Longitude he is in, the first clear Evening will discover to him the Position of the Satellite or fixed Star to the Planet, whether it be Jupiter or the Moon, and its Distance from the Planet will help the Observer to determine whether or no fuch particular Eclipse or Appulse will happen whilst the Planet is above the Horizon in that Place of Observation. 3. If an Appointment be made previous to a Series of fuch Observations, at the principal Places throughout the known World, the geographical Situation of Places may be more exactly fettled, than they can ever be expected to be by most other Ways. 4. To find all those Places, where any Eclipse or Appulse will be visible by the Planispheres. Find that Place of the Earth where the Planet will be vertical at the Time of the Eclipse or Appulse, and 90° Distance therefrom, all round the Place where the Planet is vertical, the Eclipse or Appulse may be seen if the Sky is clear and no Part thereof comes into the Sun's enlightened Hemisphere. Therefore, find where the Sun is vertical at the fame Time, and also all that Space within 90 Degrees Distance round the Place where the Sun is vertical; for to that Part of the Sun's enlightened Hemisphere which interferes with the Planet's illuminated Hemisphere, the Eclipse or Appulse will be invisible, but to all the other Parts visible. 5. What has been here observed with Respect to the Eclipses of Jupiter's Satellites, and Appulses of the Moon to the fixed Stars, is also applicable to the Appearance of a Comet, when its Place in the Heavens is known.

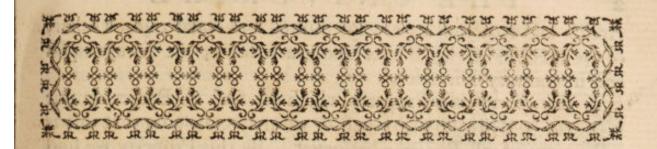
Several other Problems might be enlarged upon; but we think it quite needless, having aimed at Perspicuity without Prolixity, throughout this Treatise.

And, whereas, several Inconveniencies have arisen since the first Publication of these Planispheres for Want of the Author's having convenient Opportunity to sit them up himself (notwithstanding which Disadvantages, they have been highly approved of by eminent Mathematicians), therefore, for the surve, they are to be sitted up under the Author's own Direction, in the following Manner: On Paper or Cloth, for the Pocket, Library, or at Large, at a reasonable Expence for colouring, or other additional Ornaments.

1. Each Planisphere will have a String thro' its Centre, for Problems which require no more than the Application of a Right Line from the Centre to the Circumserence, such as related the Hour of Sun-rising and setting, Length of the Day and Night, Azimuth, and Altitude at Slis, &c. 2. Every Sett of Planispheres will have two Sliders of a permanent, glossy, transparent Paper, but little inferior to Glass itself, when pressed following a great Number of problems, some without looking thro' the Slider, and others were near its Edge; the other whole Slider to be divided in the Manner, if any Actident should happen to the divided one.

3. The divided Slider will be cut from Z to N, reserving one half for Problems which require more than 90 Degrees of the Arch of a Meridian or Great Circle, and the other Half cut into two equal Parts or Quadrants, one of which Slider Quadrants will be cut along the Parallel of 18 Degrees for Problems of the Twilight, Zodiac, or within 18 Degrees of the Horizon; and the other Part of this Quadrant will have a Centre fixed to it for Problems from the Zenith or Nadir to 18 Degrees above or below the Horizon, such as the Nature of the Problem shall require; and the other Quadrant to be cut in like Manner if wanted.

THE



THE

SOLAR SYSTEM.

Universal Planispheres, two of which together do represent the terraqueous Globe, and the other two the starry Firmament or Heaven. We think it may not be improper to introduce an Account of the Universe in general, and of the solar System in particular, of which these Planispheres do

reprefent a Part.

2. By the Word UNIVERSE we here understand the whole Frame and Extent of Nature, from the most minute Particles of Bodies which we can perceive to exist by our Senses, throughout all those various Classes of Matter, and the innumerable Multitude of Stars or Suns, which are too far distant from us for

our Understandings to comprehend.

3. Seeing that when we have extended our Enquiries into the intrinsic Nature of Things, as far as we are able, we find all that Variety which is perceived in the Universe, to arise from the Properties of Bodies, and the Laws with which they are endowed; it must therefore be allowed as quite satisfactory to have reduced any natural Phenomenon to its primitive Laws and Constitutions in Nature.

4. The Laws which we find all Bodies endeavouring to obferve, are, 1st, To continue in their present State, whether of Rest or forward uniformly in a right Line, until they are compelled to change that State, by some other Force acting on them, and this Endeavour is sometimes called the VISINERTIÆ of Matter. 2d. The Mutation or Change of Motion from one Direction to another is always proportional to the Force engaged in changing the same, and is made in the Direction of that right Line in which the Force is impressed; which if toward some Centre is called Gravity, and sometimes Attraction. 3d. Action

Action and Reaction are always mutual and equal to each other, or the Actions of two Bodies upon each other, are always equal

and in contrary Directions.

Earth to be destitute of, therefore we may conclude, that this is a Property of all Bodies whatsoever. By this Property we find that Bodies have Weight, that when they are let fall they descend freely through Space toward the Earth; by this Property we also find that the minute Particles of Matter coming very near to each other, do absolutely join and stick fast together, so as not to be easily separated, and that the great Globes of the Sun and Planets, although at great Distances from each other, do attract with Powers proportional to their Quantities of Matter, from which we may reasonably infer that this Law holds not only throughout our solar System, but also through the Universe.

6. ELASTICITY OF REPULSION is another Property which we find in some particular Bodies, but cannot discern its Origin; by this we see that when elastic Substances have been compressed, and are let loose, they immediately return to their former Situation, retaining the same Force with which they yielded. Indeed we know that this Property becomes very powerful in oily, airy, and watery Substances, when they have been potently heated by Fire; and that the Substances of Earths and Salts receive no such Activity from the same Agent, but why such different Effects should be produced by the same Cause, the most accute Naturalists have not been able as yet to understand.*

7. Ic

* It feems probable to me (fays Sir Isaac Newton,) that God in the Beginning formed Matter in folid, maffey, hard, impenetrable, moveable Particles of fuch Sizes and Figures, and with fuch other Properties, and in fuch Proportion to Space as most conduced to the End for which he formed them, and that those primitive Particles, being Solids, are incomparably harder than any porous Bodies compounded of them, even fo hard as never to wear or break in Pieces, no ordinary Power being able to divide what God himself made One in the first Creation.

While the Particles continue entire, they may compose Bodies of one and the same Nature and Texture in all Ages, but should they wear away or break in Pieces, the Nature of Things depending on them would be changed. Water and Earth composed of old worn Particles and Fragments of Particles, would not be of the same Nature and Texture now, with Water and Earth composed of entire Particles in the Beginning And therefore, that Nature may be lafting the Changes of corporal Things, are to be placed only in the various Separations, and new Affociations and Motions of thefe

Multitude of folar Systems, of which the fixed Stars appear to be so many Centres; for the Light of these Stars is of the same Kind with that of our Sun, and was he removed to so great a Distance from us as they are, he would appear like a Star as they do, and all the other Bodies in our solar System, being opake Ones, would not be visible at so great a Distance; and therefore these Systems cannot interfere with one another, whilst they are duly regulated by the Laws of Matter, although they be innumerable, and their Bounds and Extent too great for human

Reason to comprehend.

8. We find that the folar, mundane, or PLANETARY SYSTEM, contains Seventeen great Bodies nearly spherical, one of which only shines by his own native Light, the other Sixteen being opake and appearing only by the Light which comes from that luminous one. The intermediate or intermundane Spaces between those Bodies being either void of all Matter, or if there be any, it is so very thin and subtile, as to have made no sensible Resistance on the Motions of those Bodies from the first Ages of Time: The Planets and Comets appearing to move in free Spaces, and being actuated by no other Laws, but that of Gravitation toward each other, in Proportion to the Quantities of their Matter, and their Visinertiæ arising from a projectile Force, which they must have received from some powerful Agent.

9. The Sun is placed near the Centre of this System, and moves with a very small Agitation or Motion compared with the Motion of the Planets, and all the Planets and Comets move

2 round

permanent Particles, compound Bodies being apt to break, not in the midst of solid Particles, but where those Particles are laid together, and only touch in a few Points.

It feems to me, farther, that thefe Particles have not only a Visinertize accompanied with such passive Laws of Motion, as naturally result from that Force, but also that they are moved by certain active Principles, such as is that of Gravity, and that which causeth Fermentation and the Cohesion of Bodies. These Things I consider not as occult Qualities supposed to result from the specific Forms

of Things, but as general Laws of

Nature by which the Things themfelves are formed; their Truth appearing to us by Phenomena, though their Caufes be not yet discovered.

Now by the Help of these Principles, all material Things seem to have been composed of the hard and solid Particles abovementioned, variously associated in the sirst Creation, by the Counsel of an intelligent Agent. For it became him who created them to set them in order. And if he did so, its unphilosophical to seek for any other Origin of the World, or to pretend that it might arise out of a Chaos by the mere Laws of Nature. Thus far Sir Isaac Newton.

round him. He is a Body of Fire, partly liquid and partly folid; the liquid Part being an inexhaustible Ocean of Light, moving with firy Billows and flaming Ebullitions, continually diffusing innumerable Corpuscles of elementary Fire and Light to the utmost Bounds of the planetary System. The solid Parts appear like Continents, Islands, Mountains, and Rocks of Fire, with vast subterraneous Caverns and Receptacles, containing the sluid and liquid Parts, and which break forth in ignivomous Fumes when the Sun is more than usually hot on his Surface.

called Maculæ, their Shapes and Sizes are uncertain and irregular, fometimes appearing numerous for an Age together, and at other Times disappearing for so long a Time. They have all a Motion from the East-side of the Sun's Disque to the West, in a direct Parallelism with the Sun's Equator, from which it is concluded, that they are floating Dross on the solar Ocean, and that the Sun moves round his Axis with them, for they sometimes soon accumulate and at other Times as soon dissolve, it being very rarely that these Maculæ make more than one Revo-

lution before they disappear.

our Earth, Mars, Jupiter and Saturn, move round the Sun as their Centre, and in Orbits which are elliptical and may be thus described. Seal down a square Piece of white Paper on a state Board, and drive two Pins through the Paper into the Board, at the Distance of one, two, or three Inches from each other, near the Middle of the Paper; tie a Thread round the two Pins, so that when it is tied, and double, it may be two or three Inches longer than the Distance of the two Pins: Then hold a Pencil or Pen upright within the Thread whilst it is stretched tight, and carry round the Pencil or Pen, marking its Trace on the Paper, this will not be a Circle but an Ellipsis, like the Orbit described by a Planet whilst it is moving round the Sun.

one of which the Sun is supposed to be placed as at Centre: Half the Distance between the two Foci is called the Eccentricity of the Planet. The Line passing through the two Foci from one Side of the Orbit to the other, is called the transverse Diameter of the Orbit. A Line drawn from one Side of the Orbit to the other, and passing at right Angles through the Middle of the transverse Diameter, is called the conjugate Diameter of the Orbit: When the Planet is farthest distant from

the Sun, it is faid to be in its Aphelium, and when it is nearest to the Sun, it is faid to be in its Perihelium: The Aphelium is sometimes called the lower Apsides, and the Perihelium the higher Apsides of the Orbit. When the Planet is at its greatest Distance from the Earth, it is said to be in Apogeon, and when

nearest the Earth, in Perigeon.

Sun, is called the Ecliptic, and as the other Planets whilst they are moving round the Sun, do not move in the Plane of the Ecliptic, they must sometimes be above, and at other Times as much below this Plane. The Points in which the Orbits of the other Planets, cut the Plane of the Ecliptic, are called the Nodes, so there will be two Nodes in every Orbit, and the Line which joins these Nodes, is called the Line of the Nodes. That Node where the Planet ascends to the North from below the Ecliptic, is called the ascending Node, and that Node where the Planet descends to the South below the Ecliptic, is called the descending Node.

14. The Planets move fastest in their Orbits when they are nearest to the Sun, and slowest when they are farthest from the Sun, and whilst they are moving forward in their Orbits, are turning round their own Centres, under the same Direction as they move in their Orbits, which Motion is called in Consequentia, but a Motion contrary thereto, is said to be in Antecedentia. The Line passing through the Centre of the Planet, and about which it moves, is called its Axis: The Extremities of the Axis, are called the Poles. Those Planets which are nearer the Sun than our Earth, are called inferior Planets, as Mercury and Venus, and those Planets which are farther from the Sun than our Earth, are called superior Planets, as Mars, Jupiter, and Saturn.

the Distances from the Sun for all the primary Planets. If two right Lines be drawn from the Sun to different Points of the Orbit, the Time which the Planet has taken in passing from one of these Points to the other, will always be proportionable to the Area or plain Surface contained between these two Lines and the Arch of the Orbit. The Velocity of a Planet in any Point of its Orbit, is always inversely as its Distance from the Sun. The Force with which any Planet gravitates toward the Sun in

any Point of its Orbit, is always as the Square of its Distance from the Sun. The Axis of every primary Planet keeps nearly the same parallel Position whilst it is moving in its Orbit.

16. All

16. All the Planets appear to move from East to West, or to rise and set in 24 Hours, by the Rotation of the Earth round its Axis in that Time, but they really move from West to East. Sometimes they appear to have gained from West to East in 24 Hours, and then their Motion is said to be direct: Sometimes they appear to have gained from East to West in 24 Hours, and then their Motion is said to be retrograde, at other Times they appear to be in the same Place during 24 Hours, and are then said to be Stationary. All which must necessarily arise from the Motion of the Earth in its Orbit, and the Direction of the Earth in its Orbit, and the Orbit of the Earth in its Orbit of the Ear

tion and Motion of the other Planets in their Orbits.

Ecliptic, is called its Obliquity. The Distance from the Planet to the Plane of the Ecliptic, is called its Latitude. The Distance of a Planet from its Node, is called the Argument of Latitude. The Latitude as seen from the Sun, is called the heliocentric Latitude; and the Latitude as seen from the Earth, is called the geocentric Latitude. The true Distance of the Planet from the Sun, is called the curtate Distance. The Angle made by two right Lines drawn from the Earth to the Sun and Planet, is called the Angle of Elongation. The Angle which is made by two right Lines drawn from the Planet to the Earth and Sun, is called the parallactic Angle. And the Angle which is made by two right Lines drawn from the Sun to the Earth and Planet, is called the Angle of Commutation.

18. What hath been here faid of the Planets and their Orbits, is also observable in the Comets and their Orbits; for they are observed to move in free Spaces round the Sun, and in Orbits no Way different from those of the Planets, but very excentric, whereby their Perihelia come near the Sun, and their Aphelia are more distant from him, and therefore these Bodies

must fusier great Extreams of Heat and Cold.

Ten, namely, one Moon revolving round our Earth, four Moons or Satellites moving round Jupiter, and five Moons or Satellites moving round Saturn) do also observe the same Laws toward their primary Ones, as the primary Planets observe toward the Sun; and therefore whatever hath been mentioned of them, is also to be understood in like Manner of these.

appears red and firy, an opake Body receiving all his Light from the Sun, nearly round, having many Hills and Moun-

tains:

tains: He is fometimes feen like a full Moon, fometimes horned, and at other Times gibbous. He has feveral Spots, fome of which are light, and others duskish, which may be Mountains and Seas. On Account of his Nearness to the Sun, he hath a very great projectile Force, for his gravitating Force is very great, and his Orbit would otherwise be very eccentric, and therefore, by moving with a rapid Velocity, he is prevented from being fnatched into the folar Blaze. The Heat which he continually receives from the Sun, would make Water Boil on the Surface of our Earth, and was he not a very dense Body. would be calcined by the folar Heat. He is feldom feen by the naked Eye, and then a little after Sun-fetting or before Sunrifing, on Account of the Smallness of his greatest Elongation. and the Light of the Sun to which he is always adjacent. In his Passage between the Sun and our Earth, he appears like a dark Patch on the folar Difque, with his Circumference fomewhat red. He is a small Planet, and other Particulars concerning him, are contained in the following Tables.

21. VENUS is the fecond Planet in the folar System, she moves round the Sun at a greater Distance than Mercury, and next below our Earth, to which she is equal in Magnitude, tho? a little more dense. She is a bright and glittering Planet, the most regular Planet in her Motions of the whole solar System, which is owing to her fmall Excentricity, she receives twice the Light and Heat from the Sun that we do, is sometimes very near the Earth, and then is often feen in the Day time, and casts a Shadow in the Night like the Moon, and on Account of her Proximity to the lunar Orbit greatly disturbs the Motion of that Luminary. She is fometimes the Morning-star rising before the Sun, at other Times, the Evening flar fetting after the Sun; has the same Variety of Changes as the Moon, sometimes appearing full, fometimes horned, and at other Times gibbous. She has a rough and uneven Surface, and probably Lands and Seas like our Earth, is fometimes feen in the Face of the Sun, and then appears like a dark Patch much larger than Mercury. It is not certain that the has any Moons or Satellites accompanying her, although feveral Astronomers have thought to have discovered one: If the has any, their Surface may be too opake to reflect Rays to the Distance of our Earth. Other Particular lars relating to this Planet, are contained in the following Tables.

22. The EARTH, upon whose Surface we dwell, is the third Planet in the folar System, it is a Body nearly round, moves round the Sun in a Year, in an Orbit between that of Venus and Mars, its Orbit is called the ECLIPTIC, and is divided into 12 Signs, Aries v, Taurus &, Gemini II, Cancer &, Leo A, Virgo m, Libra a, Scorpio m, Sagittarius I, Capricorn ve, Aquarius m, Pisces x; each Sign containing 30 Degrees. These Signs are counted forward according to the Direction of the Earth in its Motion through the Ecliptic; and this is faid to be from West to East, or according to the Order of the Signs. Besides this annual Motion, the Earth hath another Motion round its own Axis once a Day from West to East, by which quick Rotation, the Sun, Planets, and fixed Stars, appear as though they did rife and fet from East to West in the fame Time. Moreover, the Plane of the Earth's Equator keeps the same parallel Position to itself, whilst it is moving round the Sun, and this Plane is always inclined to the Plane of the Earth's Orbit or Ecliptic under an Angle of 23 1 Degrees, by which Means there are two Points in the Earth's Orbit or Ecliptic, in one of which the Earth will revolve round its Axis, having its northern Regions much exposed to the Sun, and its southern Regions as little exposed at the same Time, and in the other of these Points, the Earth's fouthern Regions will be much exposed, whilft the northern Regions are as little exposed to the folar Rays; and these two Points are the Places of the Earth at the Summer and Winter Solftices respectively, as the intermediate Points of the Orbit opposite to each other are the Places of the Earth at the vernal and autumnal Equinoxes, and hence the Viciffitudes of the Seafons and the various Lengths of Days and Nights at different Times of the Year, at different Places on the Earth's Globe.

23. Our Earth hath an invisible light, fluid, thin, elastic Substance surrounding it, called its Atmosphere, which freely fills its Pores, and extends to the Height of about 50 Miles. The Density and Weight of this Atmosphere decreaseth as its Height increaseth. It is so dry and elastic as never to congeal or become stagnant as Water doth into Ice or Snow. Without this Atmosphere, sonorous Bodies could not be heard, Odours could not be smelt, nor could the Rays of Light be refracted to our Eyes in coming from the Sun. Hereby the Lungs receive a fresh Supply of that vivisying Spirit which animal Life destroys. This Atmosphere expands greatly with the Presence of

the Sun, condenseth in his Absence, produceth all these Changes of Beauty and Variety in the vegetable Creation which we have in Summer and Winter, is an universal Cement for all Bodies, and a Menstruum for reducing all Things to other Forms

by Duration and Time.

24. By the Weight, Elasticity and Rarefaction of the Earth's Atmosphere, great Quantities of Water are exhaled from the Seas, Rivers and Lakes by the Heat of the Sun, particularly where he is nearly vertical, and the Atmosphere is rare; these Exhalations, form Clouds, confifting of aqueous, fulphureous and faline Particles, which, when they take Fire in the Air, break out with that Violence of Light and Noise which we obferve in Lightning and Thunder. Of the same Composition are all Fogs and Mists near the Earth's Surface, which want nothing to make them Clouds in the higher Regions of the Air, but a more dense Atmosphere to swim in. Thus the Air or Atmosphere is a pellucid Menstruum for the infensible Parts of diffolved Bodies to float up and down in, until they come to a Medium specifically lighter than themselves, when they defcend freely, being subjected to no other Laws but that of Gravity to each other and the Earth, and therefore form themselves descending, into Drops of Rain. When the Atmosphere is more replete with nitrous Particles, and less replete with fulphureous Ones, the Attraction between the nitrous and aqueous Particles will be greater than the repulfive Power, and therefore the former will wholly overcome the latter, and the Fluid will become cristallized and stagnant, and this in Clouds or Drops of Rain descending through colder Regions of the Air, or intercepted by freezing Winds and warmer Air, first gently cristallizes, then gently thaws, and sticks together, falling in Fleeces of Snow. And thus when the Particles are more nitrous, or the Cold more intense, Hail is formed. When fulphureous and nitrous Vapours exhaled above the Clouds take Fire, they afcend upward as toward the Zenith or Point over our Heads, and form the Phenomenon of the Aurora Borealis, or northern Lights; when fuch Vapours are fired near the Earth's Surface, they wast to and fro with the Wind, and form the Ignis Fatuus, or Jack with a Lanthorn; and when they kindle above the Clouds in ferene Air, or long Trains, they form the Phenomena of shooting or falling Stars. The vivid Colours of the Rain-bow, are also formed of the primary and constituent Corpuscles of Light falling on small Drops of Water, from which they are reflected and refracted to our Eyes, and excite the Sensation of those Colours of Light in the Atmosphere.

25. The deeper Valleys of our Earth are filled with vaft Oceans and Seas, or Waters stretching round the Globe, the Extent of these Waters is much greater than that of the Lands, their Depths and Irregularities fomewhat proportionable to the Heights of the Mountains, with which they are generally found to compare. Great Plenty of Salt is found in these Waters, supposed to be dispersed over the Bottom of the Ocean, and to commix by Dilution. These Seas rise or swell above their usual Surface, twice every Day, which is when the Moon is either South or North on the Meridian. These Swells cause the Waters of the Ocean to flow up the narrow Seas, Channels and Rivers, which striking against irregular Shores that lie in their Way, are variously resisted until they have reduced those Rivers to even Heights with the general Swell, which fometimes requires feveral Hours to accomplish after the Moon is past the Meridian. All which is a necessary Confequence of the Gravitation of these Waters toward the Moon; for as she is much nearer to the Earth than any of the Planets beside, and the Particles of Water are globular and easily to be feparated, the Waters, where the Moon is vertical, will gravitate toward her with a greater Force, than those Waters which are diametrically opposite on the other Side of the Earth's Globe, the intermediate Seas will gravitate less than the former, but more than the latter, and the Waters where the Moon is vertical, and those on the opposite Side of the Earth, will be raised above their usual Surface, nearly equal; for the Centre of the Earth itself, gravitates toward the Moon, and the Gravitation of all its Parts, toward each other, are mutual and equal. When the Sun's Attraction is conjoined with that of the Moon, which is the Case when the Sun and Moon are in Conjunction or Opposition, that is, in a Line with each other, the Tides will then be higher, and these are called Spring-tides, but when the Moon is in her Quadratures, or a Quarter of her Orbit from the Sun, the Force of the Moon's Attraction will be leffened by that of the Sun, and these Tides are called Neap-tides. Currents or particular Settings of the Ocean arise from these and other like accidental Caufes.

26. The Transition of the Atmosphere from one Place to another, or which is the same Thing, the Motion of the Air, is the immediate Cause of Winds in all their Kinds. These are

occasioned by the Rarefaction and Condensation of Vapours, the Pressure of the Clouds, Descent of Rain, Snow and Hail, and other like accidental Causes. These Winds which blow in the larger Oceans, between, and near the Tropics, or where the Sun is nearly vertical at any Time of the Year, do constantly blow from East to West, because the Sun is constantly rarifying the Atmosphere that Way, and the subsequent adjacent Atmosphere is continually preffing the same Way to preserve an Equilibrium. The Moon also attracts the Atmosphere from East to West, and these Winds are called constant, or Trade-Winds. In leffer Oceans, variously environed with Lands, the Winds will be variously interrupted, and their natural Directions altered by the fuperior Heat of the Sun on the terrestrial Parts of the Earth's Globe, on which Account there are frequent Sea and Land Breezes near the Shores, and therefore the Winds in fuch Seas, do blow fome Part of the Year one Way, and the other Part of the Year another Way, and these are called Monfoons, or periodical Winds.

27. By Means of the Earth's Atmosphere, Springs and Rivers are furnished with a continual Supply, for as Rains defeend, and fill the Crevices, Channels and Veins of the Earth, and Clouds are exhaled from the Earth and Sea, and the Winds wast these Clouds to the Sides of Mountains, the Clouds are attracted by the Mountains and run down into the Vallies forming Rivers. Waters accidentially lodged under Mountains, are attracted upward through the Pores and spungy Parts of the Earth, and pressed out in Springs, which descend by their Sides, and carry various Particles of Sulphur, Nitre, Earth, or Minerals with them, and hence are formed Mineral Waters. Air condensed within the Earth, gusheth forth into a more rare Atmosphere, and brings great Streams and Torrents of Water with it. These Waters form Rivers, and run to the Ocean.

28. On the Surface of the terraqueous Globe, and below its Surface, are found a great Variety of different Kinds of Bodies, variously mixed and disposed, some of which give rooting to Plants, others are medicinal, and others curious. Among these are Sand, Gravel, Chalk and Rocks which are barren; Clays and Mould, or Garden Earth, producing most Things necessary for animal Sustentation. Below the Surface of the Earth, are found various Fossils, or Compositions partaking of Earth and Stone; of which Kind are Oker, Chalk and soft Leads. Stones of various Kinds, as Marble, Porphyry, Freestone, Flint,

C 2 Agate,

Agate, Cornelian, Pebble; Diamonds, Rubies, Amethysts, Emeralds; Limestones, Whetstones, and many others, with the Loadstone, whose peculiar Property is to direct itself nearly North and South, when suspended freely, and to give Iron or Steel the same Direction by a slight Touch, but this Direction changeth in Time.

which are common Salt, Vitriol and falt Gemma, also Minerals or Semi-metals, as Antimony, Cinnabar, Zink and Brimstone. But the Bodies of greater Value, dug out of the Earth, are usually distinguished by their Weight, Fusibility and Malleability, as Lead, Tin, Copper, Iron, Silver, Gold and Quicksilver.

30. The Structure and Constitution of the Earth's Globe toward its * Centre, is entirely unknown to us for certain, and

per-

* Dr. Halley's Opinion hereof is as follows: Now (faith he) to propose fomething that may answer the several Appearances, and introduce nothing firange in Philosophy, after a great many close Thoughts, I can come to no other Conclusion, than that the whole Globe of the Earth is one great Magnet, having four magnetical Poles or Points of Attraction, near each Pole of the Equator two; and that in those Parts of the World which lie near adjacent to any one of those magnetical Poles, the Needle is governed thereby, the nearest Pole being always predominant over the more Remote.

The Parts of the Earth wherein these magnetical Poles lie, cannot as vet be exactly determined for Want of fufficient Data to proceed geometrically. And this Hypothesis is sufficient for folving the great Variety and feeming Irregularity, which is obserwed in the Variations of the Compass. But to calculate exactly what the Variation is in any Place affigued, is what I dare not yet pretend to, tho' I could wish it were my Happiness to be able to oblige the World with fo ufefol a Piece of Knowledge; there are Difficulties that occur, that render the Thing as yet not feafible, for first, there are a great many Observations requifite, which ought to be made at

the same Time, not at Sea but ashore, with greater Care and Attention than the Generality of Sailors apply. And besides, it remains undetermined in what Proportion the attractive Power decreases, as you remove from the Pole of a Magnet, without which it were a vain Attempt to go about to calculate. There is yet a further Dissibility, which is the Change of the Variation, one of the Discoveries of this last Century, which shews that it will require some Hundreds of Years to establish a compleat Doctrine of the magnetical System.

Now confidering the Structure of our terraqueous Globe, it cannot be well supposed that a very great Part thereof can move within it, without notably changing its Centre of Gravity and the Equilibre of its Parts. which would produce very wonderful Effects in changing the Axis of diurnal Rotation, and occasion strange Alteration in the Sea's Surface, by Innundations and Recesses thereof, fuch as Hiftory never yet mentioned. Besides, the solid Parts of the Earth are not to be granted permeable by any other than fluid Substances, of which we know none that are any Ways magnetical. So that the only Way to render this Motion intelligible and possible, is to suppose it to turn

about

perhaps for ever will be. Indeed, by Sir Isaac Newton's Difcoveries, we know that if our Earth be an homogeneous Solid of equal Denfity throughout, that the Force of Gravity diminishes downward toward its Centre in Proportion to the Distance, and that at the Centre it becomes nothing at all, and therefore Bodies are there void of all Gravity with Respect to the Earth. Also, we know, that if the Earth's Globe was a hollow Sphere with an indefinitely thin Surface, a Particle of Matter placed any where within that Surface, would not be attracted toward the Centre nor under any other Direction. We also know, that the Denfity of the Earth, compared with its Magnitude, is lefs than the Denfity of some other Planets of the folar System, compared with their Magnitudes. And therefore there is no Inconfistency in supposing the Earth to be hollow toward its Centre,

about the Centre of the Globe, having its Centre of Gravity fixed and immoveable in the fame common Centre of the Earth. And there is yet required, that this moving internal Substance be loose and detached from the external Parts of the Earth whereon we live; for otherwise were it affixed thereto, the Whole must necessarily

move together.

So then the external Parts of the Globe may well be reckoned as a Shell, and the internal as a Nucleus or inner Globe included within ours, with a fluid Medium between, which having the fame common Centre and Axis of diurnal Rotation, may turn about with our Earth each twentyfour Hours; only this outer Sphere having its turbinating Motion fome fmall Matter either fwifter or flower than the internal Ball: And a very Minute difference in Length of Time, by many Repetitions becoming fenfible, the internal Parts will by Degrees recede from the external, and not keeping Pace with one another, will appear gradually to move either Eastwards or Westwards by the Difference of their Motions.

Now supposing such an internal Sphere having fuch a Motion; if this exterior Shell of Earth be a Magnet, having its Poles at a Distance from

the Poles of diurnal Rotation; and if the internal Nucleus be likewise a Magnet, having its Poles in two other Places distant also from the Axis; and thele latter by a gradual and flow Motion, change their Place in Respect of the external; we may then give a reasonable Account of the four magnetical Poles I prefume to have demonstrated before; as likewife of the Changes of the Needle's Variations, which till now hath been unat-

tempted.

The Period of this Motion being wonderful great, and there being hardly an hundred Years fince these Variations have been duly observed, it will be very hard to bring this Hypothefis to a Calculus, especially since though the Variations do increase and decrease regularly in the same Place, yet in different Places at no great Distance, there are found such casual Changes thereof, as can no Ways be accounted for by a regular Hypothefis, as depending upon the unequal and irregular Distribution of the magneticai Matter, within the Substance of the external Shell or Coat of the Earth, which deflect the Needle from the Position it would require from the Effect of the general Magnetism of the Whole.

Hence

Centre, and there possessed of Fire, Water, Vapour, or a Void.

Earth; of these there are various Kinds, as Herbs having soft and tender Stalks, but no Wood or hard Substance; Shrubs which grow near the Earth's Surface, and Trees which grow high and wide. Most of these have Root, Stalk, Leaf, Flower and Seed, and are all nourished by the Juices imbibed by the smaller Fibres of the Roots, from which they are conveyed throw very minute Tubuli to the Stalks, Branches and Leaves, from which these Juices return in like Manner through the Bark to the Root. All which must necessarily be on the Rarefaction and Condensation of the Atmosphere in Summer and Winter.

32. Ani-

Hence, and from some other of like Nature, I conclude that the two Poles of the external Globe are fixed in the Earth, and that if the Needle were wholly governed by them, the Variations thereof would be always the fame, with some little Irregularities, upon the Account I but just now mentioned: But the internal Sphere having fuch a gradual Translation of its Poles, does influence the Needle, and direct it variously according to the Refult of the attractive or directive Power of each Pole, and confequently there must be a Period of the Revolution of this internal Ball, after which the Variations will return again as before. But if it shall in future Ages be otherwife observed, we must then conclude there are more of these internal Spheres. and more magnetical Poles than four, which at prefent we have not a fufficient Number of Observations to determine.

If this be allowed me, 'tis plain that the fixed Poles are the Poles of this external Shell or Cortex of the Earth, and the other two the Poles of a magnetical Nucleus, included and moveable within the other. It likewife follows, that this Motion is Westwards, and by Consequence that the aforesaid Nucleus has not precisely attained the same Degree of Velocity

with the exterior Parts in their diurnal Revolution; but so very nearly equals it, that in 365 revolves, the Difference is scarce sensible. This I conceive to arise from the Impulse whereby this diurnal Motion was imprest on the Earth, being given to the external Parts, and from thence in Time communicated to the internal; but not so as perfectly to equal the Velocity of the first Motion imprest on, and still conserved by the superficial Parts of the Globe.

As to the Quantity of this Motion, it is almost impossible to define it, both from the Nature of this Kind of Obfervation, which cannot be very accurately performed, as also from the small Time these Variations have been observed, and their Change discovered. So that the nice Determination of this and of several other Particulars in the magnetic System, is reserved for remote Posterity; all that we can hope to do is, to leave behind us Observations that may be consided in, and to propose Hypothesis, which after Ages may examine, amend, or resource.

Only here I must take Leave to recommend to all Masters of Ships, and all others, Lovers of natural Truths, that they use their utmost Diligence to make, or procure to be made, Ob-

fer-

32. Animals are a most noble Part of terrestrial Beings, having Sensation and voluntary Motion. Of these some are aerial, or peculiar to the Air, as Birds and Flies; some are aquatic, or peculiar to the Waters, as Fishes; some are peculiar to the Face of the Earth, as Beasts of sour Feet, Reptiles of many Feet, and Serpents with none, and Insects or small Creatures, whose Bodies are almost separated by a stender Ligature; and some are amphibious, or peculiar to both Land and Water. All of which, have in general, Sensation by seeing, hearing, smelling, tasting and feeling, much after the same Manner, and some of these much more acutely than Man. These are sustained by their Food, which, when it is chewed and broken, mixes with Saliva, and is digested in the Stomach, conveyed through the Guts, the Lacteals separating the Chyle or nutritive Part, and

fervations of these Variations in all Parts of the World, as well in the North as South Latitude (after the laudable Custom of our East-India Commanders) and that they please to communicate them to the Royal Society, in order to leave as compleat a History, as may be, to those that are hereafter to compare all together, and to compleat and perfect this abstruct

Theory.

Thus, in order to explain the Change of the Variations, we have adventured to make the Earth hollow, and to place another Globe within it, and I doubt not but this will find Oppofers enough. I know it will be objected, that there is no Instance in Nature of the like Thing; that if there is fuch a middle Globe, it would not keep its Place in the Centre, but be apt to deviate therefrom, and might possibly chock against the Concave Shell to the Ruin, or at least endamaging thereof; that the Water of the Sea would perpetually leak through, unless we suppose the Cavity full of Wa ter; that were it possible, yet it does not appear of what Use such an inward Sphere can be of, being that up in eternal Darkness, and therefore unfit for the Production of Animals or Plants; with many more Objections,

according to the Fate of all fuch new

Propositions.

To these and all other that I can foresee, I briefly answer, that the Ring environing the Globe of Saturn, is a notable Instance of this Kind, as having the fame common Centre, and moving along with the Planet, without fenfibly approaching him on one Side more than the other. And if this Ring were turned on one of its Diameters, it would then describe such a Concave Sphere as I suppose our external one to be. And fince the Ring in any Polition given, would in the fame Manner keep the Centre of Saturn in its own; it follows, that fuch a concave Sphere may move with another included in it, having the fame common Centre. Nor can it well be supposed otherwise, considering the Nature of Gravity; for should these Globes be adjusted once to the same common Centre, the Gravity of the Parts of the Concave, would prefs equally toward the Centre of the inner Ball, which Equality must necessarily continue till fome external Force difturb it.

As to the leaking of the Water through this Shell, when once a Paffage shall be found for it to run thro', I must confess it an Objection feemingly conveying it into the Blood with which it concocts, and is forced through the Lungs, where it is enlivened by the inspiring Air, is conveyed into the left Ventricle of the Heart, and from thence distributed by the Arteries to all Parts of the Body, from whence it is brought back by the returning Veins, and fo takes the fame Courfe.

33. The

ingly of Weight: But when we confider how lightly great Beds of Chalk or Clay, and much more Stone do hold Water, and even Caves arched with Sand; can we then think it a hard Supposition, that the internal Parts of this Bubble of Earth should be replete with such faline and vitriolic Particles as may contribute to Petrifaction, and dispose the transuding Water to shute and coagulate into Stone, fo as continually to fortify, and if Need were, to confolidate any Breach or Flaw in the concave Surface of the Shell.

And this, perhaps, may not without Reason be supposed to be the final Cause of the Admixture of the magnetical Matter in the Mass of the terrestrial Parts of the Globe, viz. to make good and maintain the concave Arch of this Shell. Another Argument favouring this Hypothesis, is drawn from a Proposition in Sir Isaac Neauton's Principia, where he determines the Force wherewith the Moon moves the Sea in producing the Tides: His words are Denfitas Lunce est ad densitatem Terræ ut 680 ad 387 seu 9 ad 5 quam proxime. Est igitur corpus Lunæ denfius ac magis terestre quam terra noftra.

The Denfity of the Moon is to the Denfity of the Earth, as 680 to 387, or as 9 to 5 nearly. Therefore the Body of the Moon is more dense, and more earthy than the Earth itself.

Now if the Moon be more folid than the Earth, as 9 to 5, why may we not reasonably suppose the Moon, being a fmall Body and a fecandary Planet, to be folid Earth, Water, Stone, and this Globe to confift of the same Materials, only four Ninths thereof to be Cavity, within and between the internal Spheres.

And fince it is now taken for granted, that the Earth is one of the Planets, and they all are with Reason supposed habitable, tho' we are not able to define by what Sort of Animals; and fince we fee all the Parts of the Creation abound with animate Beings, as the Air with Birds and Flies, the Water with the numerous Varieties of Fish, and the very Earth with Reptiles of fo many Sorts; all whose Ways of living would be to us incredible, did not daily Experience teach us: Why then should we think it strange, that the prodigious Mass of Matter, whereof this Globe doth confift, should be capable of fome other Improvement than barely to ferve to support its Surface. But still it will be faid, that without Light there can be no Living; and therefore, all this Apparatus of our inward Globes must be useless: To this I answer, that there are many Ways of producing Light which we are wholly ignorant of; the Medium itself may be always luminous, after the Manner of our Ignes fatui. The concave Arches may in feveral Places shine with such a Substance as invests the Face of the Sun; nor can we, without a Boldness, unbecoming a Philosopher, adventure to affert the Impossibility of peculiar Luminaries below, of which we have no Sort of Idea.

Thus far from Dr. Halley, and which we have transcribed as a Memento or Remembrancer to fuch of our Readers, as shall in after Times be furnished with more ample Observations for fettling the Course and Pe-

riod of the magnetic System.

33. The Understanding is the most distinguishing Characteristic of the human Species, and is peculiar only to the Mind. For although we can perceive Bodies by our Senses, * retain full and clear Ideas of them, and compare or compound those Ideas, yet we could not know their Result without this Understanding in our Minds. The Manner in which we come to D

* The immortal Sir Isaac Newton, having given us, in his Optics, a most concile yet sublime Description of the Phenomena of Nature, worthy the Attention of the most Knowing, and particularly such as would acquire a certain Knowledge in philosophical Matters; we beg Leave to introduce an Extract thereof as follows:

If (faith he) in two large tall cylindrical Vessels of Glass inverted, two little Thermometers be suspended so as not to touch the Vessels, and the Air being drawn out of one of these Vessels, and these Vessels, and these Vessels, and these Vessels thus prepared be carried out of a cold Place into a warm one; the Thermometer in Vacuo will grow warm as much, and almost as soon as the Thermometer which is not in Vacuo. And when the Vessels are carried back into the cold Place, the Thermometer in Vacuo will grow cold almost as soon as

the other Thermometer.

Is not the Heat of the warm Room conveyed through the Vacuum by the Vibrations of a much fubtiler Medium than Air, which after the Air was drawn out remained in the Vacuum? And is not this Medium the same with that Medium by which Light is refracted and reflected, and by whose Vibrations Light communicates Heat to Bodies, and is put into Fits of eafy Reflection and easy Transmission? And do not the Vibrations of this Medium, in hot Bodies, contribute to the Intenseness and Duration of their Heat? And do not hot Bodies communicate their Heat to contiguous cold Ones, by the Vibrations of this Medium propagated from them into the cold Ones? And is not this Medium exceedingly more rare and fubtile, than Air, and exceedingly more elaftic and active? And doth it not readily pervade all Bodies? And is it not by its elaftic Force expanded thro' all the Heavens?

Doth not the Refraction of Light proceed from the different Denfity of this ætherial Medium in different Places, the Light receding always from the denfer Parts of the Medium? And is not the Denfity thereof greater in free and open Spaces, void of Air, and other groffer Bodies, than within the Pores of Water, Glass, Crystal, Gems, and other compact Bodies? For when Light paffes through Glass or Crystal, and falling very obliquely upon the farther Surface thereof, is totally reflected, the total Reflection ought to proceed rather from the Denfity and Vigour of the Medium without, and beyond the Glafs, than from the Rarity and Weakness thereof.

Doth not this ætherial Medium, in passing out of Water, Glass, Crystal, and other compact and dense Bodies into empty Spaces, grow denser and denser by Degrees, and by that Means refract the Rays of Light, not in a Point, but by bending them gradually in Curve-lines? And doth not the gradual Condensation of this Medium extend to some Distance from the Bodies, and thereby cause the Inflections of the Rays of Light, which pass by the Edges of dense Bodies, at some Distance from the Bodies?

Is not this Medium much rarer within the dense Bodies of the Sun, Stars, Planets and Comets, than in the empty celestial Spaces between them? And in passing from them to greater Distances, doth it not grow denser and

den-

understand, is by our Senses. For Seeing the Eye is an Organ, wonderfully adapted with its Coats and Humours, besides which a luminous Body, emitting indefinitely small Particles or Rays; Air or Æther to turn these Particles out of a straight Course; and Bodies or Objects themselves not pellucid, are all necessary to the Completion of Vision. Such Rays coming from the Sun and

denfer perpetually, and thereby cause the Gravity of those great Bodies toward one another, and of their Parts toward the Bodies? Every Body endeavouring to go from the denfer Parts of the Medium toward the Rarer? For if this Medium be rarer within the Sun's Body than at this Surface, and rarer there than at the hundredth Part of an Inch from his Body, and rarer there than at the fiftieth Part of an Inch from his Body, and rarer there than at the Orb of Saturn; I fee no Reason why the Increase of Denfity should stop any where, and not rather be continued through all Distances from the Sun to Saturn and beyond. And though this Increase of Denfity may at great Distances be exceedingly flow, yet if the elaftic Force of this Medium be exceedingly great, it may fuffice to impel Bodies, from the denfer Parts of the Medium towards the Rarer, with all that Power which we call Gravity.

And that the elastic Force of this Medium is exceeding great, may be gathered from the Swiftness of its Vibrations. Sounds move about 1140 English Feet in a Second of Time, and in feven or eight Minutes of Time they move about One Hundred English Miles. Light moves from the Sun to us in about feven or eight Minutes of Time, which Distance is about 70,000000 English Miles, supposing the horizontal Parallax of the Sun to be about 12 Seconds. And the Vibrations of Pulses of this Medium, that they may cause the alternate Fits of easy Transmition and easy Reflection, must be swifter than Light, and by Confequence, above 700,000 Times fwifter than Sounds. And therefore,

Proportion to its Desity is 700,000 x 700,000 (viz. above 490,000,000000)
Times greater than the elastic Force of the Air is in Proportion to its Density. For the Velocities of the Pulses of elastic Mediums, are in a subduplicate Ratio of the Elasticities, and the Rarities of the Mediums taken together.

As Attraction is stronger in all small Magnets than in great Ones in Proportion to their Bulk, and Gravity is greater in the Surfaces of small Planets than in those of great Ones in Proportion to their Bulk, and fmall Bodies are agitated much more by electric Attraction than great Ones; fo the Smallness of the Rays of Light may contribute very much to the Power of the Agent by which they are refracted. And so if any one should suppole that Æther (like our Air) may contain Particles which endeavour to recede from one another (for I do not know what this Æther is) and that its Particles are exceedingly fmaller than those of Air, or even than those of Light: The exceeding Smallness of its Particles, may contribute to the Greatness of the Force by which those Particles may recede from one another, and thereby make that Medium exceedingly more rare and elastic than Air, and by Confequence less able to refift the Motions of Projectiles, and exceedingly more able to prefs upon gross Bodies, by endeavouring to expand itself.

May not Planets and Comets, and all grois Bodies, perform their Motions more freely and with less Resistance in this ætherial Medium, than and Stars, a Candle or other luminous Substance, with a very great Degree of Swiftness, pass to the Retina or Bottom of the Eye, where they paint the Images of the Objects against which they have last impinged, and this immediately excites Vision. For Hearing the Ear is an Organ, which receives the Vibrations of the Air occasioned by a quick and tremulous Motion of the D 2

in any Fluid which fills all Space adequately without leaving any Pores, and by Consequence is much denser than Quickfilver or Gold? And may not its Resistance be so small as to be inconsiderable? For Instance; if this Æther (for so I will call it) should be supposed 700,000 Times more elastic than our Air, and above 700,000 Times more rare, its Resistance would be above 600,000000 Times less than that of Water. And so small a Resistance would scarce make any sensible Alteration in the Motions of the Planets in 10,000 Years.

. It any one would ask how a Medium could be fo rare, let him tell me how the Air in the upper Part of the Atmosphere can be above a hundred thousand Times rarer than Gold. Let him also tell me how an electric Body, can by Friction emit an Exhalation to rare and fubtile, and yet fo potent, as by its Emission, to cause no sensible Diminution of the Weight of the electric Body, and to be expanded thro' a Sphere, whose Diameter is above two Feet, and yet to be able to agitate and carry up Leaf-copper, or Leaf-gold, at the Distance of above a Foot from the electric Body. And how the Effluvia of a Magnet can be to rare and lubtile as to pass through a Plate of Glass without any Resistance or Diminution of their Force, and yet to potent as to turn a magnetic Needle beyond the Glass?

Is not Vision performed chiefly by the Vibrations of this Medium excited in the Bottom of the Eye, by the Rays of Light, and propagated through the folid, pellucid and uniform Capillamenta of the optic Nerves, into the Place of Sensation? And is not Hearing performed by the Vibrations, either of this or fome other Medium, excited in the auditory Nerves by the Tremours of the Air, and propagated through the folid, pellucid and uniform Capillamenta of those Nerves into the Place of Sensation? And so of the other Senses.

Is not animal Motion performed by the Vibrations of this Medium, excited in the Brain by the Power of the Will, and propagated from thence thro' the folid, pellucid and uniform Capillamenta of the Nerves into the Murcles for contracting and dilating them? I suppose that the Capillamenta of the Nerves are each of them folid and uniform, that the vibrating Motion of the ætherial Medium may be propagated along them, from one End to the other uniformly, and without Interruption: For Obstructions in the Nerves create Palfies. And that they may be fufficiently uniform, I suppose them pellucid when viewed fingly, though the Reflections in their cylindrical Surfaces, may make the whole Nerve (composed of many Capillamenta) appear opake and white. For Opacity arises from reflecting Surfaces, fuch as may diffurb and interrupt the Motions of this Medium.

Do not all fixed Bodies, when heated beyond a certain Degree, emit Light and shine? And is not this Emission performed by the vibrating Motion of their Parts? And do not all Bodies which abound with terrestrial Parts, and especially with sulphurous Ones, emit Light as often as those Parts are sufficiently agitated, whether that Agitation be made by Heat, or by Friction or Percussion, or Putrefaction, or by any vital Motion, or

fonorous Body, on a fudden Stroke or Shock. For Smelling, the olfactory Nerves, receiving odoriferous Particles, excite this Senfation. For Tafting, the Palate is an Organ, which, receiving the Contact of fweet, bitter, fower, and other Substances, excites this Sensation. For Touch or Feeling, the whole nervous System, diffused unequally through the Body,

any other Cause? As for Instance, Sea-water in a raging Storm; Quickfilver agitated in Vacuo; the Back of a Cat or Neck of a Horse, obliquely ftruck or rubbed in a dark Place; Wood, Flesh and Fish, while they Putrify; Vapours arising from putrified Water, usually called Ignes Fatui; Stacks of moist Hay, or Corn growing hot by Fermentation; Glowworms, and the Eyes of some Animals by vital Motions; the vulgar Phosphorus agitated by the Attrition of any Body, or by the acid Particles of the Air; Amber, and fome Diamonds, by striking, pressing, or rubbing them; fcrapings of Steel, struck off with a Flint; Iron hammered very nimbly, till it becomes so hot as to kindle Sulphur thrown upon it; the Axle-trees of Chariots taking Fire by the rapid Rotation of the Wheels; and some Liquors mixed with one another, whose Farcles come together with an Impetus, as Oil of Vitriol distilled from its Weight of Nitre, and then mixed with twice its Weight of Oil of Anifeeds.

Is not Fire a Body heated fo hot as to emit Light copiously? For what clfe is a red hot Iron than Fire? And what else is a burning Coal than red hot Wood? Is not Flame a Vapour, Fume, or Exhalation heated red hot, that is, fo hot as to shine? For Bodies do not Flame without emitting a copious Fume, and this Fume burns in the Flame. The Ignis Fatuus is a Vapour shining without Heat, and is there not the same Difference between this Vapour and Flame, as between rotten Wood shining without Heat and

burning Coals of Fire?

Do not great Bodies conferve their Heat the longest, their Parts heating

one another, and may not great dense and fixed Bodies, when heated beyond a certain Degree, emit Light fo copioufly, as by the Emission and Reaction of its Light, and the Reflections and Refractions of its Rays within its Pores to grow still hotter, till it comes to a certain period of Heat, such as is that of the Sun? And are not the Sun and fixed Stars great Earths vehemently hot, whose Heat is conserved by the Greatness of the Bodies, and the mutual Action and Reaction between them and the Light which they emit, and whose Parts are kept from fuming away, not only by their Fixity, but also by the vast Weight and Dentities of the Atmospheres incumbent upon them; and very strongly compressing them, and condensing the Vapours and Exhalations which arise from them? For if Water be made warm in any pellucid Veffel emptied of Air, that Water in the Vacuum will bubble and boil as vehemently as it would in the open Air in a Vessel set upon the Fire, till it conceives a much greater Heat. For the Weight of the incumbent Atmofphere keeps down the Vapours and hinders the Water from boiling until it grow much hotter than is requifite to make it boil in Vacuo. Alfo a Mixture of Tin and Lead being put upon a red hot Iron in Vacuo emits a Fume and Flame; but the fame Mixture in the open Air, by reason of the incumbent Atmosphere, does not to much as emit any Fume which can be perceived by Sight.

In like Manner, the great Weight of the Atmosphere which lies upon the Globe of the Sun, may hinder Bodies there from rifing up and going

away

and most powerfully at the Finger's Ends may be taken for one continued Organ, and in contact with hard, soft, smooth, rough, dry, wet, hot and cold Bodies excites this Sensation. From which it may be observed, that all Sensations are but different Modifications of Feeling, or the Contact of Bodies on some particular Part of the human Frame. By these Sensations,

away from the Sun in the Form of Vapours and Fumes; unless by Means of a far greater Heat than that which on the Surface of our Earth would very eafily turn them into Vapours and Fumes. And the same great Weight may condense those Vapours and Exhalations as foon as they shall at any Time begin to ascend from the Sun, and make them prefently fall back again into him, and by that Action increase his Heat much after the Manner that in our Earth the Air increases the Heat of a culinary Fire. And the same Weight may hinder the Globe of the Sun from being diminished, unless by the Emission of Light, and a very fmall Quantity of

Vapours and Exhalations.

Do not the Rays of Light, in falling upon the Bottom of the Eye, excite Vibrations in the Tunica Retina? Which Vibrations being propagated along the folid Fibres of the optic Nerves into the Brain, cause the Sense of Seeing. For because dense Bodies conserve their Heat a long Time, and the denfest Bodies conserve their Heat the longest, the Vibrations of their Parts are of a lasting Nature, and therefore may be propagated along folid Fibres of uniform dense Matter to a great Distance, for conveying into the Brain the Impressions made upon all the Organs of Sense. For that Motion which can continue long in one and the same Part of a Body, can be propagated a long Way from one Part to another, supposing the Body homogeneal, fo that the Motion may not be reflected, refracted, interrupted or difordered by any Unevenness of the Body.

Do not feveral Sorts of Rays make Vibrations of feveral Bignesses, which according to their Bignesses excite Sensations of several Colours, much after the Manner, that the Vibrations of the Air, according to their several Bignesses, excite Sensations of several Sounds? And particularly do not the most refrangible Rays excite the shortest Vibrations for making a Sensation of deep Violet, the least refrangible, the largest for making a Sensation of deep Red, and the several intermediate Sorts of Rays, Vibrations of several intermediate Bignesses to make Sensations of the several intermediate Colours?

May not the Harmony and Discord of Colours, arise from the Proportions of the Vibrations propagated through the Fibres of the optic Nerves into the Brain, as the Harmony and Discord of Sounds arise from the Proportions of the Vibrations of the Air? For some Colours, if they be viewed together, are agreeable to one another, as those of Gold and Indigo,

and others difagree.

Are not gross Bodies, and Light convertible into one another, and may not Bodies receive much of their Activity from the Particles of Light which enter their Composition? For all fixed Bodies being heated, emit Light fo long as they continue fufficiently hot, and Light mutually stops in Bodies as often as its Rays strike upon the Parts. I know no Body lefs apt to shine than Water; and yet Water by frequent Distillations changes into fixed Earth, as Mr. Boyle has tried; and then this Earth being enabled to endure a fufficient Heat, fhines by Heat like other Bodies.

The changing of Bodies into Light, and Light into Bodies, is very con-

form-

tions, we have strong Impressions on our Minds, called simple Ideas, which we are able to compare and compose, making of them complex Ones. From these we are able to draw abstract Ideas, less dependant on Bodies, and by Concatenation and Ratiocination deduce complex Truths.

34. The

formable to the Course of Nature, which feems delighted with Tranfmutations. Water which is a very fluid tasteless Salt, she changes by Heat into Vapour, which is a Sort of Air, and by Cold into Ice, which is a hard, pellucid, brittle, fufible Stone; and this Stone returns into Water by Heat, and Vapour returns into Water by Cold. Earth by Heat becomes Fire, and by Cold returns into Earth. Dense Bodies by Fermentation rarify into several Sorts of Air, and this Air by Fermentation, and fometimes without it returns into dense Bodies. Mercury fometimes appears in the Form of a fluid Metal, fometimes in the Form of a hard, brittle Metal, fometimes in the Form of a corrofive. pellucid Salt, called fublimate, fometimes in the Form of a tasteless, pellucid, volatile, white Earth, called Mercurius Dulcis; or in that of a red. opake, volatile Earth, called Cinnabar; or in that of a red or white Precipitate; or in that of a fluid Salt: and in Distillation it turns into Vapour, and being agitated in Vacno, it shines like Fire. And after all these Changes, it returns into its first Form of Mercury. Eggs grow from infenfible Magnitudes, and change into Animals; Tadpoles, into Frogs; and Worms, into Flies. All Birds, Beafts and Fishes, Infects, Trees and other Vegetables, with their feveral Parts, grow out of Water and watry Tinctures and Salts, and by Putrefaction return again into watry Substances. And Water standing a few Days in the open Air, yields a Tincture, which (like that of Malt) by standing longer yields a Sediment and a Spirit, but before Putrifaction is fit Nourish-

ment for Animals and Vegetables. And among fuch various and strange Mutations, why may not Nature change Bodies into Light, and Light into Bodies?

Have not the small Particles of Bodies certain Powers, Vertues or Forces, by which they act at a Distance, not only upon the Rays of Light for reflecting, refracting and inflecting them, but also upon one another for producing a great Part of the Phenomena of Nature? For it's well known, that they act one upon another by the Attractions of Gravity, Magnetism and Electricity; and these Instances shew the Tenure and Course of Nature, and make it not improbable that there may be more attractive Powers than thefe. For Nature is very confonant and conformable to herfelf. How these Attractions may be performed, I do not here confider. What I call Attraction, may be performed by Impulse, or by some other Means unknown to me. I use that Word here to fignify only in general any Force by which Bodies tend towards one another, whatfoever be the Caufe.

By Experiments compared with the great Quantity of Sulphur with which the Earth abounds, and the Warmth of the interior Parts of the Earth, and hot Springs and burning Mountains, and with Damps, mineral Corufcations, Earthquakes, hot suffocating Exhalations, Hurricanes and Spouts; we may learn, that sulphureous Steams abound in the Bowels of the Earth and ferment with Minerals, and sometimes take Fire with a sudden Coruscation and Explosion; and if pent up in subterraneous Caverns, burst the Caverns with a great shaking of the

Earth,

34. The Moon is the Second great Luminary of the Heavens. She is the nearest Planet to our Earth, and much less than either of the primary Planets, moves round the Earth between the Orbits of Venus and Mars, is disturbed in her Motion, together with the Earth, by the Attractions of the other Planets, and particularly Venus, which is large and sometimes very near. These Disturbances render her true Place and that of the Earth uncertain to a great Degree of Exactness, and the Computation of the same for any past, present, or suture Time, is a Task with the

great-

Earth, as in springing a Mine, and then the Vapour, generated by the Explosion, expiring through the Pores of the Earth, feels hot and fuffocates, and makes Tempests and Hurricanes, and fometimes causes the Land to flide, or the Sea to boil, and carries up the Water thereof in Drops, which by their Weight fall down again in Spouts. Also some sulphureous Steams, at all Times when the Earth is dry, ascending into the Air, ferment there with nitrous Acids, and fometimes taking Fire, cause Lightning and Thunder and fiery Meteors. For the Air abounds with acid Vapours, fit to promote Fermentations, as appears by the rusting of Iron and Copper in it, the kindling of Fire by blowing, and the beating of the Heart by means of Respiration. Now the abovementioned Motions are fo great and violent, as to flew that in Fermentations the Particles of Bodies which almost rest, are put into new Motions by a very potent Principle, which acts upon them only when they approach one another, and causes them to meet and clash with great Violence, and grow hot with the Motion, and dash one another into Pieces, and vanish into Air, Vapours and Flames.

As Gravity makes the Sea flow round the denfer and weightier Parts of the Globe of the Earth, fo the Attraction may make the watry Acid flow round the denfer and compacter Particles of the Earth for composing the Particles of the Salt. For otherwise the Acid would not do the Office of a Medium between the Earth and

common Water, for making Salts diffolvable in the Water; nor would Salt of Tartar readily draw off the Acid from disfolved Metals, nor Metals the Acid from Mercury. Now, as in the great Globe of the Earth and Sea, the denfest Bodies by their Gravity fink down in Water, and always endeavour to go towards the Centre of the Globe; fo in Particles of Salt, the densest Matter may always endeavour to approach the Centre of the Particle: So that a Particle of Salt may be compared to a Chaos, being denfe, hard, dry and earthy in the Centre; and rare, foft, moist and watry in the Circumference. And hence it feems to be, that Salts are of a lasting Nature, being scarce destroyed, unless by drawing away their watry Parts by Violence, or by letting them foal; into the Pores of the central Earth, by a gentle Heat in Putrefaction, until the Earth be diffolved by the Water, and separated into smaller Particles, which by Reason of their Smallness, make the rotten Compound appear of a black Colour. Hence also it may be, that the Parts of Animals and Vegetables preferve their feveral Forms, and affimilate their Nourishment, the foft and moist Nourishment easily changing its Texture by a gentle Heat and Motion, till it becomes like the dense, hard, dry and durable Earth in the Centre of each Particle. But when the Nourishment grows unfit to be affimilated, or the central Earth grows too feeble to affimilate it, the Motion ends in Confusion, Putretaetion and Death.

greatest Mathematicians of the present Age, next to insuperable. Could this be made easily practicable, the geographical Situation of Places on the Earth's Globe, and the Art of Navigation would be more certain. She is an opake Body, having her Light from the Sun, but continually prefents the same Face toward our Earth, turns round her own Axis exactly in the same Time in which she makes a Revolution round her Orbit, from which it has been thought that she abounds with magnetical Matter like our Earth, and that her Position is directed thereby. One half of her Surface is illuminated by the Sun, the other half is dark. That Side facing our Earth, hath abundance of Irregularities and Protuberances with Light and dark Spots, which have been taken for Mountains, woody or rushy Plains, Lakes and Seas of great Amplitude and Extent, also round Cavities, dark at Bottom, many Leagues wide, of great Depth, with conical Mounts in the Midst which have been taken for deep Pits, some dry and others partly filled with Water. This Planet is very irregular in her apparent Motion, receding 360 Degrees, and traverling through all the Signs, with wide Excursions from North or South in about 27 - Days, in which Time the finishes a Revolution round the Earth, her Orbit being inclined to that of the Earth, under an Angle of 5 th Degrees, and the Time of her coming to the South, about 53 Minutes later every Day than the preceding one. She appears largest when rising or fetting (as doth the Sun) but measures less in Diameter then, than when she is in the Zenith; the Cause of this is the Nature and Properties of Vision. At the autumnal Equinox she rifes about the same Hour for several Days together, which is occasioned by the Position of her Orbit, the Position of the Earth's Orbit, and the Polition of the Horizon at that Time, and as this is about the Time of Harvest, she is then called the Harvest or Hunter's Moon. She casts a very long conical Shadow, which falls on the Earth, when the is between the Sun and Earth, this makes an Eclipse of the Sun, the Earth also cafts a much larger conical Shadow, which, when it is between the Moon and Sun, falls on the Moon and makes an Eclipse of that Luminary. The Moon appears to have neither Clouds nor Vapours, and therefore is thought to have no Atmosphere, or, if any, but a very thin one, nor is it known whether the is inhabited, although generally supposed to be so, like our Earth. Other Particulars relating to this Planet, may be feen in the following Tables.

moves round the Sun between the Orbits of the Earth and Jupiter. He is much less, and also less dense than our Earth, to which he comes very near in Perigeon, and therefore sometimes appears large, red and siery, and at other Times very small. He has an Atmosphere with a large dark Spot in the Middle of his Face, and other Irregularities, which probably are Seas and inhabited Lands. Other Particulars relating to this Planet,

may be feen in the following Tables.

36. JUPITER is the fifth, and by far the largest Planet in the folar System, he moves round the Sun, between the Orbits of Mars and Saturn, and has but a twenty-seventh Part of the Heat from the Sun that we have. He has a rough Face with dark Spots, and three duskish Belts, or Girdles, crossing his Body, all which are subject to small Changes and Alterations, and are therefore thought to be Clouds and Vapours swiming in his Atmosphere; if so, this Planet must enjoy great Serenity of Weather, but if they adhere to his Body, great Diforder and Confusion of Lands and Waters, he having all the other Properties of a habitable World. Four Moons, or SATELLITES, are continually moving round this Planet, as the Moon moves round our Earth, all of which are within Jupiter's Blaze and invisible to the naked Eye, but are feen through a Telescope, in a right Line with each other and Jupiter, except when they are hid behind his Body, in his Shadow, or in his Rays, which latter Place renders them obscure. The Immersions and Emerfions of these Satellites furnish frequent Opportunities for taking the Longitudes of Places at Land wherever the Eclipses can be observed, the Immersions being always to be observed when Jupiter is more West, and the Emersions when he is more East than the Sun, and these would answer the same End at Sea, and make the Art of Navigation perfect, could good Telescopes be used so as to keep the Planet and Satellites in View on the Ocean. Other Particulars relating to this Planet, may be feen in the following Tables.

37. SATURN is the fixth and highest primary Planet in the folar System. He is much larger than either of the Planets except fupiter, and moves round the Sun in an Orbit of vast Extent, between the Orbit of fupiter below, and the fixed Stars, which are at an immense Distance beyond him. He is less dense than our Earth, and has but little more than a hundredth Part of the Heat and Light that we enjoy, and therefore his polar

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Regions must be as many Times intensly cold as at the Poles of the Earth's Globe. He appears small, whitish and dull, with an irregular Surface like the other Planets, and probably hath Inhabitants. This Planet has a large and wonderful Ring of solid Matter encompassing his Body, without any where touching it, of which extraordinary Phænomenon various Conjectures have been made, but nothing certain is yet known. Five Satellites, or Moons, also continually move round his Body, like the Satellites of Jupiter round Jupiter's Body, the Immersions of which can only be seen when he is West before the Sun, and the Emersions when he is East after the Sun. Their Uses being the same with those of Jupiter, but not so visible. Other Particulars relating to this Planet, may

be feen in the following Tables.

38. Comers, or Blazing-stars, are wandering Planets, which move round the Sun, in elliptical Orbits which are very eccentric, and in their Motions they are regulated by the same Laws as the other Planets. These Bodies are very numerous, their Orbits variously inclined, and it is quite uncertain when they will appear, their true Theory being but of late discovered, and their periodic Revolutions not yet observed, beside which they are accelerated and retarded in their Motions by the Actions of the other Planets, which will always make the Time of their Return very uncertain. They are opake Bodies, having their Light from the Sun, appear under different Forms and different Colours, generally with a long fiery Tail or Blaze, which makes a very frightful Appearance in the Heavens, in a dark Night when the Sky is clear. Some of them cross the Orbits of the Planets, descend to their Perihelia nearer the Sun than Mercury, and are there above 2000 Times hotter than red-hot Iron, from thence ascend to their Aphelia, immensly far beyond the Regions of Saturn, and are there most intensly cold. Their Tails are produced of the inflameable Matter of their Bodies and Atmospheres, which flies off in fiery Vapours by the Action of the folar Rays, and is fo exceedingly rarified as to be able to produce no bad Consequences on our Earth, should it ever happen to be invelloped therein at a confiderable Distance; which possibly may be, for these Bodies, crossing the solar System in all Directions, may strike against the Planets, and therefore our Earth, the Effects of which might be the uniting of the Bodies and their Centres of Gravity into one, the Removal of the Oceans to other Parts of the Earth's Globe, and great Deva-Station

the centripetal Force of each Body, or a repulfive Power in both, may conftruct a new Orbit for each, and the scattered Fragments be carried away with them. Such a Salute we hope will never happen to our Earth. The Origin and Use of these Bodies is not known, but they are thought to fall freely into the Sun, after making a great Number of irregular Gyrations round his Body, and that they are constant Supplies to him of fresh Fuel for maintaining his Blaze. Other Particulars relating to

Comets, may be feen in the following Tables.

39. The FIXED STARS, are those apparently small, lucid, twinkling Lights, which bedeck the Heavens when the Moon doth not shine, and the Sky is clear. They appear to rife and fet, or move round the Axis of our Earth in twenty-four Hours from East to West, like the Sun and Planets, which is occafioned by the Earth's Rotation from West to East, and are faid to be fixed, because they continually keep the same Position among themselves. The more visible Ones are but few, those invisible to the naked Eye are innumerable, when looked at through good Telescopes. They all shine with a native Lustre, have no dependance on our Sun or any Part of the folar Syftem, appear like lucid Points through the best Telescopes, and probably are fo many Suns, each having his Planets and Comets like ours. The Distance of these Stars is at least 40,000 Times greater than the Distance of the Sun from us; the Time which the Rays of Light take in coming from them to us is fix Years, at the Rate of 10,000000 Miles per Minute, and their Distance from the Sun so great, that at either of them he appears but a lucid Point, as they do to us, and Ninety-nine hundredths of the intermediate Distance is dark Night. Many bright Places are also observed in the Heavens, which, when viewed through good Telescopes, appear to be innumerable Multitudes of small Stars, of which the milky Way is one continued Band or Girdle paffing quite round the Heavens. All the fixed Stars have an apparent Motion in a Direction parallel with the Ecliptic, at the Rate of about 50 Seconds of a Degree per Annum, by which they are continually altering their Longitudes, right Ascensions and Declinations, but their Latitudes remain invariable, or nearly fo; besides this, they appear to describe a small Ellipsis round their true Places in the Heavens, which is occasioned by the Abberation of Light, or its coming from those Stars to us.

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TABLES of the SOLAR SYSTEM.

40. A TABLE, shewing the greatest, mean, and least Distances of the primary Planets from the Sun, with their Eccentricities in Miles.

	Greatest Distance.		Leaft Diffance.	Eccentricity. Miles.
Mercury	37.752.480		0 0	
Venus	58.960.710	58.590.000	58.094.010	427.770
Earth	82.402.920	81.000.000	79.597.080	1.370.520
Mars	134.771.850	123.418.890	111.941.190	11.421.000
Jupiter	441.458.910	421.289.100	400.890.060	20.290.500
Saturn	1816.757.020	772.578.000	729.254.360	44.307.000
dans le f			m our Earth.	a de i me i
Moon	258.131	240.67	223.734	1 17.016

41. A TABLE, Shewing the Diameters of the Sun and Planets in Miles, and also the Surface, Magnitude and Quantity of Matter in each, the Surface and Magnitude of our Earth being 1, or Unity.

Sun Mercury Vénus Earth	Diameters. 763.460 Miles 2.460 7.906 7.942	Surfaces. 9.241 10 10 10 10 10 10 10 10 10 10 10 10 10 1	Magnitud. 888.300 nearly I	Quant. Matt. 200.000
Mars Jupiter Saturn Moon	4.444 81.155 67.870 2.175	104 73	1066 624	214 76

42. A TABLE, Shewing how many Miles the Planets move in their . Orbits, per Day, Hour and Minute, at a Mean.

			The same of the sa
ne noneka	Per Day.	Per Hour.	Per Minute.
Mercury	2.181.818	90.909	1.515
Venus	1.576.837	65.701	1.095
Earth	1.395.307	58.138	985
Mars	1.074.235	44.759	746
Jupiter	587.189	24.466	407
Saturn	433.301	18.054	301
Moon	52.833	2.201	36
2 9 3 4 7 7 7 9 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1			

43. A TABLE, shewing the Planets daily mean Motions, with the Number of Miles they are drawn from their Restilinear Directions per Hour, their Densities, Light and Heat, compared with our Earth.

to District	Daily Motions.	Hourly Grav.	Denficies.	Light & Heat.
Sun	at rest.	at rest.	4	45.000
Mercury	4° 5′ 32″	129 Miles	1-1	7
Venus	1 36 8	361	4 5	2
Earth	0 59 8	25	I	I
Mars	0 31 57	8	7	10
Jupiter	0 4 59	71	4	27
Saturn	0 2 0	100	6	1 20
Moon	13 9 24	10	14	1

44. A TABLE, shewing the Times in which the Planets move round the Sun, and also round their own Axis.

. F. S.	Rotations round their own Axes. Days, Hrs, Min.			Revolutions round the Sun.
Sun	Bert 16.2.7			at reft.
		5		Days. Hrs. Min.
Mercury	0	6	0	87 23 16
Venus	0	23	0	224 16 49
Earth	0	23	56	365 5 49
Mars	1	0	40	686 23 27
Jupiter	0	9	56	4332 12 20
Saturn	29	10	I	10759 6 36
Moon	27	7	45	27 7 45 round the Earth.
				round the Earth.

45. ATABLE of the Inclination of the Planets Orbits, to the Plane of the Ecliptic, and the Inclination of their Axis to their own Orbits.

	Inc. to Ecliptic.	Axes Inc. to their own Orbits.
Sun	87 30	At reft.
Mercury	6 54	Not known.
Venus	3 24	Not known.
Earth	coincident	66° 30′
Mars	1 52	Perp. nearly.
Jupiter	I 20	Perp. nearly.
Saturn	2 30	31° 00′
Moon.	5 18	Right Ang. nearly.

46. A TABLE of the Sun and Planets greatest and least apparent Diameters, as viewed from the Earth.

	Diam	et. in P	erigeon.	[Diam	et, in A	pogeon.
Sun	32'	47"	0"	31'	40"	0"
Mercury	0	II	48	0	4	4
Venus	I	5	58	0	9	34
Mars	0	20	50	10	2	46
Jupiter	0	24	12	10	14	36
Saturn	0	19	40	0	14	II
Moon	29	24	0	33	40	0

47. A TABLE, shewing the Places of the Planets Nodes and Apheliums.

WAR TON	1	No	des.	1500	1	Aphe	liums.	
Mercury	8	15°	ı'	54"	1	13°	7'	54"
Venus	п	14	25	54	22	4	19	54
Earth	100	No	ne.		1/20	8	I	10
Mars	8	18	29	54	THE SET	0	31	54
Jupiter	69	7	19	54	-2-	9	9	54
Saturn	69	21	49	54	1	27	49	54

48. A TABLE, sheaving the mean Distances of Jupiter's Satellites from his Centre, in Diameters of his Body, and the Times of their Revolutions round him.

Satellites.	Diftances,	Revolutions.	Magnit.
First	25	Days. Hrs. / //	2
Second	41	3 13 13 42	3
Third	7 6	7 3 42 36	I
Fourth	1 12-3	16 16 32 9	1 4

49. A TABLE, shewing the Weights of equal Quantities of Matter, on the Surface of each Planet.

	Weights.	AND B	Weights.
Sun	25	Mars	1
Mercury	36	Jupiter	2
Venus	4 4	Saturn	1 11
Earth	1	Moon	1 3

50. A TABLE, shewing the Mean Distances of Saturn's Satellites from his Centre, in Diameters of his Ring; and the Times of their Revolutions.

The Distance of Saturn's Ring from his Body 21.000 Miles.

Breadth or Thickness of the Ring — 21.000

Diameter of the Ring passing thro' Saturn 152.000

Inclination of the Ring to the Ecliptic 31°.

Satellites.	Diftances.		Revolu	tions.	
Total Total	of white and	Days.	Hrs.	1	11
First	19	1	21	18	26
Second	I-3	2	17	41	10
Third	1-7	4	12	25	10
Fourth	4	15	22	41	28
Fifth	1113	79	7	46	0

51. A TABLE, shewing in how long Time each of the Planets are conjoined with the Sun.

Mercury, in a Conjunction of the same Kind, in 115 Days. Venus, in a Conjunction of the same Kind, in 584 Days. Mars, in Conjunction with the Sun, every 780 Days. Jupiter, in Conjunction with the Sun, every 398 Days. Saturn, in Conjunction with the Sun, every 378 Days.

52. A TABLE, shewing how many Degrees the Sun is usually under the Horizon, before the Planets and Stars appear.

Moon is often feen in the Day-time.	
Venus when horned about	50.
Jupiter and Mercury -	10°.
Saturn and Mars	110.
Stars of the First Magnitude	12°.
Second — — —	13°:
Third — — —	14°.
Fourth — —	15°.
Fifth — — —	16°.
Sixth — — —	17°.
Beginning and End of Twilight -	18°.

53. ATABLE, shewing the Increase of Heat toward the Equinoctial Line, and its Decrease toward the Poles, when the Sun is in the Equator; according to the Obliquity of the solar Rays.

Latitude.	Quantity of Hea	The Dillance of Verma a Ring for
00	20.000	Decading of Thicknets of the
11	19.696	I hameter of the King palling
20	18.796	THE TOTAL OF THE COLUMN
30	17.321	This Table will be alike useful
40	15.321	for any Time when the Sun is not
50 60	12.155	in the Equator, if due Allowance
60	10.000	be made for his Declination.
70	6.140	A STATE OF THE PARTY OF THE PAR
80	3.473	Fourth La Land
90	0.000	Titte thirt

54. ATABLE, shewing when Venus and Mercury may be seen in the Sun's Disque, and at what Distance from his Centre.

Planets.	Times.		h m	Central Distance.
Venus 1	1761	June 6,		1 4 7 c above
ALC: NO VINCE		01		4 15 above.
Venus	1769	June 6,	II o Even.	15 43 below.
Mercury	1769	Nov. 9,	10 10 Even.	7 24 below.
Mercury	1776	Nov. 2,	10 19 Even.	15 23 above.
Mercury	17820	Nov. 12,	3 44 Even.	15 27 below.
Mercury	1786	May 4,	6 57 Morn.	12 43 below.
Mercury	1789	Dec. 6,.	3 53 Even.	7 20 above.
Mercury	1799	May 7,	2 34 Aftern.	4 12 above.
Venus	1996	June 10,	1 13 Aftern.	13 36 above.
Venus	2004	June 8,	7 18 Morn.	6 22 below.
Venus	2109	Dec. 13,	2 56 Aftern.	14 36 below.
Venus	2117	Dec. 11,	4 3 Morn.	10 5 above.

TABLES of the Comets, and also Tables for finding the Moon's Place, and thereby the Tides for many Years to come, are inserted farther on.

GEOGRAPHICAL DEFINITIONS.

55. # AVING in the foregoing Part of this Treatife, H given a particular Account of the Laws, Periods, Distances, Velocities and other Affections of the * A * celestial Bodies, and shewn how the various Phenomena of the Earth and Heavens do primarily arise from the established Laws of Matter and Motion; I shall now introduce fuch Definitions as are necessary for understanding the Sciences of Geography, Astronomy, and others depending on them. And, because in the Construction of the terrestrial and celestial Globes, of which the universal Planispheres are Representatives, and in all Treatifes of their Uses, a peculiar Respect has been had to the apparent Motion of the Sun, Moon and Stars round our Earth in 24 Hours, although at the fame Time there be not wanting fufficient Evidence to evince that this Phenomenon is principally owing to the Motion of the Earth itself: I shall therefore, for the Sake of Ease and Analogy, introduce the Uses of these Planispheres upon the same Plan.

56. A Circle, great or small, is supposed to be divided into 360 equal Parts, called Degrees, each of these Degrees into 60 equal Parts, called Minutes, and each of these Minutes into 60 equal Parts, called Seconds. And when a Circle is the greatest that can be drawn on the Surface of a Globe, it is called a great Circle, but otherwise, a lesser Circle of the Sphere. The Earth hath two Points in its Surface, diametrically opposite to each other, called its Poles, the one North, the other South. A straight Line, drawn from Pole to Pole, and through the Centre of the Earth, is called the Earth's Axis, or Axis of the World. The Equinoctial, or Earth's Equator, is a great Circle, having all its Parts 90 Degrees from each Pole, and dividing the Surface of the Earth's Globe into two Hemispheres, the one North, the other South. Meridians are great Circles paffing through the two Poles, and dividing the Earth into two Hemispheres, the one East, the other West. The sensible Horizon, is that Circle which divides the upper and visible Part of

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the Heavens from the lowest and invisible. The rational Hori-

Earth's Diameter. The Tropic of Cancer is a leffer Circle, 23 \frac{1}{2} Degrees North of the Equinoctial, and parallel therewith. The Tropic of Capricorn is a leffer Circle, 23 \frac{1}{2} Degrees South of the Equinoctial, and parallel therewith. The Equinoctial, and parallel therewith. The Artic Circle is a leffer Circle, 23 \frac{1}{2} Degrees from the North Pole, and the Antarctic Circle is a leffer Circle, 23 \frac{1}{2} Degrees from the South Pole. The Poles, Axis, Equator, Tropics and polar Circles being extended into the Heavens, have like Names for the celeftial Sphere.

57. The torrid Zone, is all that Space of the Earth and Sea between the Tropics of Cancer and Capricorn. The North temperate Zone, is that Space between the Tropics of Capricorn and the arctic Circle. The South temperate Zone, is that Space between the Tropic of Capricorn and the antarctic Circle. The North frigid Zone, is that Space between the North Pole and the arctic Circle. And the South frigid Zone, is that Space between the South Pole and the antarctic Circle. The Latitude of a Place, is its Distance from the Equator on a Meridian in Degrees and Minutes. Parallels of Latitude are leffer Circles parallel to the Equinoctial. The Longitude of a Place, is the Number of Degrees and Minutes from the Place to a first Meridian, counting on a parallel of Latitude. Difference of Latitude, is the Number of Degrees and Minutes which one Place lieth more North or South than the other. And Difference of Longitude, is the Number of Degrees and Minutes of Longitude, which one Place lieth more East or West than the other.

58. Climates are certain Spaces, each of which is contained between two parallels of Latitude, the longest Day in the Middle of one Climate being half an Hour longer or shorter than the longest Day in the Middle of the next adjacent Climate. There are 48 of these, 24 from the Equator to each Pole. Besides which, there are 12 others, 6 from each polar Circle to its Pole, whose Breadths are such, that the Sun appears to one a Month sooner than it appears to the next adjacent one, in his Return from the remoter Pole. Amphiscii, are such as inhabit the torrid Zone, their Shadows being cast either North or South. Periscii, are such as inhabit the frigid Zone, their Shadows moving round them. Heteroscii, are such as inhabit the temperate Zones, their Shadows being cast only one Way. Antoecii, are such Inhabitants as live under the same Meridian, but on different Sides of the Equator, and equally distant therefrom; having the same Hour of Day or Night. Perioecii, are such

In-

Inhabitants as live in the same Parallel of Latitude, but on different Sides of the same Meridian, their difference of Longitude being 180 Degrees. Antipodes, are such Inhabitants as are diametrically opposite to each other on different Sides of the Earth's Globle.

59. A right Sphere, is when the Poles of the World are in the Horizon. A parallel Sphere, is when the Equator is in the Horizon. An oblique Sphere, is when neither the Poles nor the Equator are in the Horizon. When a Person faceth the North Point of the Horizon, his Right-hand directs toward the East, his Left-hand toward the West, and behind him is the South. These are called the four Cardinal Points, between which are feven others, making in all 32, being 11 1 Degrees distant from each other, as in the following Tables. Of the Cardinal Points, Geographers respect the North, Astronomers the South, Augurs the East, Poets the West. Bearing, is that Point of the Compass on which one Place lieth from another. Distance, is the Number of Miles or Degrees of the Equator, measuring from one Place to another. Rhombs, fometimes called Loxodromiques, are curve Lines, leading from one Place to another, and cutting all the Meridians through which they pass at equal Angles.

the Sea, having many Countries and Kingdoms. And an Ocean, or Sea, is a great and general Collection of Waters. An Island, is a Part of the Earth encompassed round with Water. And a Lake, is a Tract of Water encompassed by Lands. A Promontory, is a Part of the Continent extending itself into the Ocean. And a Bay, is a Part of the Ocean extending itself between Lands. A Peninsula, is a Part of the Earth almost surrounded by Water, but joined to the Continent by a narrow Neck of Land, called an Islands. And a Gulph, is a Part of the Ocean almost surrounded by Lands, but joined to the Ocean by a narrow Neck of Water, called a Streight. Mountains, Vallies, Morasses, Rivers, Rivulets, Springs, &c. are generally well known, and therefore need not to be placed with the aforegoing

Definitions.

ASTRONOMICAL DEFINITIONS.

61. 芦溪溪灣HE celeftial Equator, is a great Circle in the Hea-T & vens, coincident with the Plane of the Earth's Equator produced, and dividing the Heavens into *** two equal Parts, called the northern and fouthern Hemispheres. The Poles of the celestial Equator are two Points in the Heavens, diametrically opposite to each other, the one above the Horizon, and the other below the Horizon in any Latitude. The Axis of the World, is an imaginary Line drawn thro' the Earth's Centre from Pole to Pole. The celestial Meridians are great Circles croffing the Equator and meeting at the Poles. The Ecliptic, is a great Circle cutting the Equator in two opposite Points under an Angle of 23 1 Degrees, and having the 12 Signs, Aries V, Taurus &, Gemini II, Cancer &, Leo A. Virgo nx, Libra =, Scorpio m, Sagittarius I, Capricorn vo, Aquarius m, Pisces x. The Zodiac is a Zone or Belt, 18 1 Degrees broad, on each Side of the Ecliptic 9 1, in some Part of which Space the Planets are always found.

The Equinoctial Colure, is a Meridian passing through the Beginning of Aries and Libra, or where the Ecliptic cuts the Equator. And the Solfticial Colure, is a Meridian passing thro's the Beginning of Cancer and Capricorn. The Poles of the Ecliptic are two Points in the Heavens diametrically opposite to each other, 23½ Degrees distant from the Poles of the World, and 90 Degrees distant from the Ecliptic itself. The Latitude of a Star, is the Star's nearest Distance from the Ecliptic. And the Longitude of a Star, is the Number of Degrees and Minutes contained between the Star and a great Circle point of the Poles of the Ecliptic and the Beginning of Aries, counting in that parallel to the Ecliptic in which the Star is, from Aries

to Taurus, &c.

62. The Tropic of Cancer, is a lesser Circle of the Heavens parallel to the Equator, and 23 ½ Degrees North therefrom. The Tropic of Capricorn, is a lesser Circle of the Heavens, parallel to the Equator, and 23½ Degrees South therefrom. The Artie Circle, is a lesser Circle of the Heavens, 23½ Degrees distant from the North Pole. And the Antarctic Circle, is a lesser

lesser Circle of the Heavens, 23 1 Degrees from the South Pole. The Zenith, is that Point of the Heavens, directly over our Heads. And the Nadir, is that Point of the Heavens, directly under our Feet. Amplitude, is that Point of the Horizon, where a celestial Body rises or sets, counting from the East or West in Degrees and Minutes. Azimuth, is that Part of the Horizon, perpendicularly below or above any celeftial Phenomenon, counting from the North or South Points of the Horizon, in Degrees and Minutes. Vertical Circles, are great Circles, paffing from the Zenith to the Nadir, and cutting the Horizon at right Angles. Almacanthars, or parallels of Altitude, are leffer Circles parallel to each other from the Horizon to the Zenith. Parallels of Declination, are leffer Circles parallel to the Equator from the Equator to the Poles. The Latitude of a Place, is the Height of the Pole above the Horizon, or the Distance from the Equator to the Zenith. The Altitude of a celestial Body, is its Height above the Horizon in Degrees and Minutes. The Declination of a celestial Body, is its Distance from the Equator in Degrees and Minutes.

63. Right Ascension, is so much of the Equator as cometh to the Meridian with the Sun or Star, counting from the Beginning of Aries, or where the Ecliptic cuts the Equator. Oblique Ascension, is so much of the Equator as riseth with the Sun or Star. And Oblique Descension, is so much of the Equator as fets with the Sun or Star, counting as above. Ascensional Difference, is the Difference between the Right and Oblique Afcension. Parallax, is the Difference between the apparent and true Place of a celestial Body, the former as seen from the Earth's Surface, the latter as estimated from the Earth's Centre. Refraction, is the Elevation of a celestial Body, apparently in Altitude, whereby it appears to rife fooner and fet later than it really doth. Cosmical Rising or Setting of a Star, is when the Star rifes or fets when the Sun rifes. Anchronical Rifing or Setting, is when the Star rifes or fets when the Sun fets. Heliacal Rifing, is when a celestial Body hath been near the Sunbeams, and is visible before Sun-rising. And Heliacal Setting, is when the celestial Body sets inconspicuous with the Solar Rays.

64. A spherical Triangle, is a three cornered Figure, whose three Sides are Arches of a great Circle of the Earth or Heavens. A spherical Angle, is measured by the Arch of a great Circle, palling from one of the Sides next the Angle to the other, at those

those Sides produced, at the Distance of 90 Degrees from the angular Point. Projection of the Sphere in Plano, or on a Plane, is the Delineation of half a Sphere or Globe, as it would appear, was an Eye at rest, a transparent Plane interposed between the Eye and Sphere at right Angles to the Axis of Vision, and all the Parts of the half Sphere, as seen by the Eye, delineated on the Plane. A Planisphere, is a Projection of the Sphere upon a Plane, and if it be that of the Earth, it is called a terrestrial Planisphere, but if it be that of the Heavens, it is called a cele-

Stial Planisphere.

Universal Planispheres, are such as will supply the Uses of the terrestrial and celestial Globes, or Models of the Earth and Heavens, with all the real and imaginary Lines belonging to them; and solve the Problems of Geography, Astronomy and other Sciences depending on these, by the Help of proper Data. An eastern terrestrial Planisphere, is an eastern Hemisphere of the Earth, projected on the Plane of its Meridian. A western terrestrial Planisphere, is a western Hemisphere of the Earth, projected on the Plane of its Meridian. An eastern celestial Planisphere, is an eastern Hemisphere of the Heavens, projected on its Meridian. A western celestial Planisphere, is a western Hemisphere of the Heavens, projected on its Meridian. A Slider, Index, or Projector, is a Projection only of the Meridian.

ans, Equator and its Parallels.

65. Chronology, is a Science whereby we understand the Parts of Time or Duration by the Motions of the celestial Bodies. A natural Day contains 24 equal Divisions, called Hours. An Hour contains 60 equal Divisions, called Minutes. And a Minute contains 60 equal Divisions, called Seconds. A Year, is that Space of Time in which the Earth makes one Revolution in its Orbit. A Solar Year, is that Space of Time which the Sun apparently takes in departing from one Place in the Heavens and returning to it again, this is, 365 Days, 5 Hours, 48 Minutes, 57 Seconds. A siderial Year, is that Space of Time in which the Sun returns to the same Position with respect to the fixed Stars, and this is 365 Days, 6 Hours, 9 Minutes, 14 Seconds. The Lunar Year, is that Space of Time in which the Moon makes 12 Revolutions round the Earth, which she doth in 354 Days, 8 Hours, 48 nutes, 38 Seconds. The Epast, is the Space of 11 Days, Deing the Difference between the Solar and the Lunar Year. Julian Year, is 365 Days, 6 Hours, or 365 Days, three fuccreding Years, and 366 Days, every fourth Year, which is called

called Bissextile or Leap-year. A Gregorian Year, is the same as a Julian Year, only it begins the first of January, whereas the Julian Year begins the twenty-sifth of March, and, as in the Julian Account, every hundredth Year is a Leap-year, so in the Gregorian Account, every four hundredth Year is a common Year. France, Spain, Portugal, Italy, Poland; the Popish Electors in Germany, and most other Places use the Gregorian Account or New-stile, and England changed from the Julian or Old-stile to the New, 2 September 1752, calling the next Day September 14. The Anticipation of the Equinoxes, is 11 Minutes, 3 Seconds per Annum, being the Difference between the Julian and solar Year, whereby the equinoctial Points are carried backward, or the Sun carried forward a Day in 131 Years.

66. A Lunar fynodical Month, is the Space of Time in which the Sun and Moon are parted and conjoined apparently after one Revolution, and this is 29 Days, 12 Hours, 44 Minutes, 3 Seconds. A Lunar periodical Month, is that Space of Time in which the Moon returns to the fame Place of the Ecliptic, and this is 27 Days, 7 Hours, 43 Minutes, 8 Seconds. A civil Month, is from 28 to 31 Days. An artificial Day, is the Space of Time from Sun-rifing to Sun-fetting. An astronomical Day, is the Time in which the Sun apparently leaves the Meridian and returns to the same again. A civil Day begins with the Fews, Greeks, Bobemians, Italians and Chinese, at Sun-setting; with the Perfians and Syrians, at Sun-rising; with the Egyptians, Romans, English, French, Spaniards, Germans and Dutch, at Midnight; with the Arabians and modern Astronomers, at Noon; and ends at the Beginning of the fucceeding Day. Equal Time, is that meafured out by Clocks or Watches. True Time, is that measured out by the unequal apparent Motion of the Sun, or rather the Motion of the Earth in its Orbit.

67. The Cycle of the Sun, is 4 Times 7, or 28 Years, in which Space of Time the Dominical or Sunday Letters A, B, C, D, E, F, G, do make all their Varieties with the Days of the Month according to the Julian Account. But as this Cycle is 11 Minutes and 3 Seconds of Time flower per Annum, than the folar Year, the Julian Account, in every 130 \(\frac{1}{2}\) Years, must be a Day later than the Sun. Hence, although at the Nicene Council, Anno Dom. 325, the Equinox was March 21, Anno Dom. 1582, it was March 11, when Pope Gregory added 10 Days, and directed as above.

Days, and directed as above.

68. A Lunar Cycle, otherwise called the Golden-number, is the Space of 19 Years, proposed by Meton, 432 Years before Christ, and which has been continued in the Julian Account, for shewing all the Positions or Aspects which the Sun and Moon can make with each other; and for settling Easter-day the first Sunday after the first full Moon, happening after the twenty-first Day of March, or vernal Equinox. But this Number hath been found defective, by 1 Day in 304 Years; and seeing it is now near 4 \frac{3}{4} Times 304, or near 1444 Years since the Nicene Council; therefore the new Moons according to this Cycle, do happen about 4 \frac{3}{4} Days sooner than they really do in the Heavens.

69. A Chaldean Saros, is that Space of Time in which the Sun and Moon make all their various Positions or Aspects to each other, with respect to the Earth. This is performed with but little Inequalities in 223 Lunations, or 18 Years, 10 Days, 7 Hours, 43 Minutes, 30 Seconds. Therefore, add these Years, Days, Hours and Minutes, to the Day, Hour and Minute of a new Moon, full Moon, Eclipse of the Sun or Moon, and you have the Time when that Lunation or Eclipse will return with much the same Circumstances as before. Thus much may suffice for the Definitions of Time; indeed the Antients used other Periods, but they were much less certain than the Julian, whose Simplicity and Certainty is such, that for the Purposes of Astronomy, it seems to be preferable to the Gregorian itself.

70. A TABLE, shewing how many Days must be added to the Julian Account, for reducing it to the Gregorian.

1700 Leap-year 11 Days	2800 Leap-year 19 Days
1800 12	2900 20
1900 13	3000 21
2000 Leap-year 13	3100 22
2100 — 14	3200 Leap-year 22
2200 15	3300 23
2300 16	3400 24
2400 Leap-year 16	3500 25
2500 17	3600 Leap-year 25
2600 18	3700 26
2700 — 19	3800 27
A THE RESERVE AND ADDRESS OF THE PARTY OF TH	3900 28, 530.

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TABLE,

SHEWING

The Sun's Place in the ECLIPTIC, Right ASCENSION, DECLINATION and EQUATION of TIME for every Day in the BISSEXTILE Year, 1760.

ALSO,

His apparent Diameter, hourly Motion, Time in transiting the Meridian, and Distance from the Earth in Miles, every tenth Day.

Note. This Table being calculated for the Year 1760, New-Style, and the Meridian of London, may be readily used for any other Year of the present Age, adding 1' 48" for every succeeding Bissextile Year, and subtracting 14' 20" for every Year after Leap Year, the Sum or Remainder shews the Sun's Place for that Day; opposite to which, or by a proportional Difference, you have the Right Ascension and Declination required.

Numbers to be added to, or subtracted from the Sun's Place, in the following Table, from 1 to 40 Years; besides, which, for an exact Determination, a proportional Allowance should be made for any Part of a Year.

	,	n		,	"		,	11		,	11
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2 fub.	28	40	12 add	5	24	22 fub.	19	40	32 add	14	24
3 fub.	43	0	13 fub.	8	56	23 fub.	34	0	33 fub.	0	4
4 add	1	48	14 fub.	23	16	24 add	10	48	34 fub.	14	24
5 fub.	12	32	15 fub.	37	36	25 fub.	3	32	35 fub.	28	44
6 fub.	26	52	16 add	7	12	26 lub.	17	52	36 add	16	12
7 fub.	41	12	17 fub.	7	8	27 fub.	32	12	37 add	1	50
						28 add					
9 fub.	10	44	19 fub.	35	48	29 fub.	1	44	39 fub.	26	50
10 fub.	25	40	20 add	9	0	30 fub.	16	4	40 add	18	0

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72. AT ABLE, shewing the Variation of the Sun's Declination, to every Hour, or 15 Degrees of Longitude from the Meridian of London, his daily Variation being known.

Hours and Degrees of Longitude, East or West, from the Meridian of London.

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73. A TABLE, shewing the Dip of Horizon to 50 Feet Height above the Surface of the Earth and Sea. Also the Refraction of the Atmosphere from the Horizon to the Zenith, for the Parallel of London and the Torrid Zone.

Feet. Dip. Alt. R. Lond. R. Equat. Alt. R. Lond R. Equat. 1 1 8 0 33 45 27 0 41 1 2 0 51 31 58 2 17 8 15 53 43 0 58 0 46 45 5 2 33 4 10 48 10 5 45 0 54 0 44 0 36 6 4 7 6 5 48 0 48 0 49 0 47 0 39 10 3 35 9 5 22 4 50 50 0 45 0 38 0 31 12 3 56 10 4 52 4 20 51 0 44 0 36 18 4 0 36		CONTRACTOR OF THE PARTY	for the										
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75. A TABLE, shewing the Number of Days, from any Day in any Month, to the same Day in any other Month.

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Jan. 365	Dec. 334	Nov 304	O. 273	Sept. 243	Aug. 212	July 181	June 151	May 120	Apr. 90	Mar. 59	Feb. 31	Jan.
3651 Feb. 365	Jan. 334	Dec. 303	Nov.273	08. 242	Sept. 212	Aug. 181	July 15	June 120	May 89	April 59	Mar. 28	Feb.
Mar. 365	Feb. 337	Jan. 306	Dec. 275	Nov. 24.5	08. 214	Sept. 184	Aug. 153	July 122	June 92	May 61	April 31	Mar.
Apr. 365	Mar. 334	Feb. 305	Jan. 275	Dec. 244	Nov.214	04. 183	Sept. 153	Aug 122	July 91	June 61	May 30	April.
May	Apr. 335	Mar.304	Feb. 276	Jan. 245	Dec. 214	Nov.184	O.A. 153	Sept. 123	Aug. 92	July 61	June 31	May
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Aug. 365	July 334	June 304	May 273	Apr. 243	Mar. 212	Feb. 184	Jan. 153	Dec. 122	Nov. 92	08. 61	Sept. 31	August.
Sept. 365	Aug. 334	July	June	May 242	Apr. 212	Mar	Feb. 153	Jan. 122	Dec. 91	Nov. 61	08. 30	Sept.
O.A. 365	Sept. 335	Aug. 304	July 273	June 243	May 212	Apr. 182	Mar. 151	Feb. 123	Jan. 92	Dec. 61	Nov. 31	October.
Nov. 365	08. 334	Sept. 304	Aug.	July 242	June 212	May 181	Apr. 151	Mar.120	Feb. 92	Jan. 61	Dec30	Nov.
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76. ATABLE, shewing the Gregorian Sunday Letters for these Years, viz.

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77. A TABLE of the fixed Holy-days, and remarkable Days throughout the Year.

JANUARY, xxxi Days.

1 Circumcifion

4 Sir Isaac Newt bo. 1642, N S.

5 Old Christmas Day

6 Twelfth-Day, or Epiphany

8 Lucian, Anno 307

10 Princess Elizabeth born 1740

13 Hillary, 357

15 Exchequer opens

18 Prisca, 45

20 Fabian, 353. Oct. Hill. 1 ret.

21 Agnes, 304

22 Vincent, 304 23 Hillary Term begins 25 Conversion of St. Paul

27 Quind. Hill. 2 Return

30 K. Cha. I. Mart. 1648-9, O. S.

FEBRUARY, xxviii Days.

2 Purif. or Cand. Day

3 Blaize. 3 Return

5 Agatha, 252 9 Oct. Purif. 4 r.

12 Hillary Term ends

13 Old Candlemas Day

13 Valentine, 45 24 Matthias

28 Hare-hunting goes out.

MARCH, xxxi Days.

r David

2 Chad.

5 Pris. Mary of Hesse born 1722.

7 Perpet. Maur. 252 12 Gregory, 604

17 St. Patrick

18 Edward King of West Saxony

19 Prsse Louisa Ann b. 1749. Joseph Cuthbert. Day and Night equal

21 Benedict, 542

25 Annunciation, Pr. Edward born 1739. 1st Quarter Day.

APRIL, xxx Days.

1 All Fools Day

3 Rich. B. of Chich. 1218

4 Ambrose

5 Old Lady Day

19 Alphege, 1006

23 St. George 25 St. Mark

26 Duke of Cumberland b. 1721

27 Victory of Cul. 1746.

MAY, xxxi Days.

I Philip and James

3 Invention of the Cross

6 St. John ante Port. Lat. 98

19 Dunstan, 988

24 Prince Fred. William b. 1750

26 Augustine, 596. 1 Return

27 Ven. Bede, 673

29 King Charles II. Nat. & Restor.

JUNE, xxx Days.

I Nicom. 41

4 Geo. Prince Wales born 1728

5 Boniface, 725

10 Princess Amelia born 1711

11 Barnabas, 50

17 St. Alban

20 Trans. of Edward K. of W. Sax.

22 King George II. Inaug.

24 Nat. St. John Bapt. 2 Quar. Day

29 St. Peter, 68

30 Buck-hunting comes in, and continues till Holy Rood.

JULY, xxxi Days.

2 Visitation of the Virgin Mary

3 Dies Comition

4 Tranf. St. Mart.

5 Old Midfummer Day

7 Tho. a Becket

15 Swithin,

15 Swithin, 862

20 Marg. 243

22 Mary Magdalen, 84. Prís. Carolina and Matilda b. 1751, Magdalen College Election.

25 St. James, 42

26 Ann

30 Dog Days begin.

AUGUST, xxxi Days.

1 Lammas Day

6 Transfiguration

7 Name of Jesus

10 St. Laurence, 259

11 Princess Augusta born 1737

15 Affumption 21 Athanasius

24 St. Bartholomew, 73

28 Austin, 432

29 Beheading of St. John Baptift.

SEPTEMBER, xxx Days.

I Giles, 750

2 London burnt 1666, O. S.

7 Enurchus, 374 Dog Days end

8 Nat. of the bleffed Virgin Mary

14 Exal. of the Cross

17 Lambert, 733

21 St Matthew, 90 22 Equal Day and Night

26 Cyprian, 288

28 Sheriff of London fworn

29 St. Michael. 3d Quarter Day. Hare-hunting comes in, and lasts till the End of February

30, Jerom, 420.

OCTOBER, xxxi Days.

1 Remigius, 544

6 St. Faith, 290

9 St. Dennis

10 Old Michaelmas

13 Trans. of K. Ed. Conf. 1165

15 Etheldred V.

18 St. Luke

22 King George II. crowned, 1727

25 Crispin

28 St. Simon and Jude, 64.

NOVEMBER, xxx Days.

1 All Saints

2 All Souls Election at all Souls Col. Oxen

3 Cras. anim. 1 Return

5 Gunpowder Treaf. 1605, O. S.

6 Leonard, 456, Michaelma. Term begins

7 Prince Henry Frederick, born

8 Ld Mayor fworn at Westminster

9 King George II. born 1683, N. S. Birth-Day kept the 10th, Lord Mayor's Feaft-day, London

11 St. Martin, 397

12 Cras Mart. 2 Return

13 Britius, 412

15 Machutus, 500

17 Hugh, 1200

19 Octab. Mart. 3 Return

20 Edmund, 871

22 Cicilia, 225, Old Michaelmas Day

23 Clem. 92

25 Prince William Henry, born

Quid. Mart. 4 Ret. Catherine 28 Michaelmas Term ends

30 St. Andrew

Princess Dowager of Wales, b.

DODLINGS :

DECEMBER, xxxi Days.

4 Barbary

6 Nicholas, 342

8 Concept. bleffed Virgin Mary

13 Lucy

16 O Sapientia

21 St. Thomas, 35 Shortest Day

Fox-hunting comes in, and lasts till Lady-day

26 St. Stephen

27 St. John the Evangelift

28 Holy Innocents

31 Sylvester, Conf. 335.

78. A TABLE of New and Full Moons.

Anno 1760.

F. Jan. 2 4 A. F. Jan. 11 10 M. N. Jan. 3 10 M. N. Jan. 11 1 M. N. Jan. 18 6 M. N. Jan. 25 4 M. F. Jan. 19 1 M. F. Jan. 26 1 M. F. Jan. 18 6 M. N. Jan. 25 4 M. F. Jan. 19 1 M. F. Jan. 26 1 M. F. Feb. 1 6 M. F. Feb. 8 8 A N. Feb. 2 1 M. N. Feb. 9 Noon. N. Feb. 16 8 A. N. Feb. 23 11 A. F. Feb. 17 1 A. F. Feb. 24 7 A. R. Mar. 1 9 A.F. Mar. 10 7 M N. Mar. 2 5 A. N. Mar. 10 10 A. 8 M N. Mar. 25 4 A. F. Mar. 17 Midn. F. Mar. 26 11 M. N. Mar. 17 F. Mar. 31 1 M F. Apr. 8 5 A. N. Apr. 1 10 M. N. Apr. 9 10 M. 6M F. Apr. 16 9M F. Apr. 24 Midn. 5 A. N. Apr. 24 N. Apr. 15 F. Apr. 30 6 M. F. May 8 4 M N. May 1 3 A. N. May 8 11 A. N. May 15 1M N. May 23 4 A F. May 15 5 A. F. May 24 10 M. F. May 29 9 A F. June 6 5 A. N. May 30 7 A. N. June 7 Noon.

N. June 13 8 M. N. June 22 1 M F. June 13 Midn F. June 22 5 A.

F. June 28 Noon F. July 6 5 A. N. June 29 9 M. N. July 7 2 M.

N. July 12 3 A N. July 20 8 A. F. July 13 9 M F. July 22 1 M.

F. July 28 1 M F. Aug. 4 8 A. N. July 28 9 A. N. Aug. 5 6 A.

N. July 28 1 M F. Aug. 4 8 A. N. July 28 9 A. N. Aug. 5 6 A.

N. July 28 1 M F. Aug. 4 8 A. N. July 28 9 A. N. Aug. 5 6 A. N. Aug. 11 1 M N. Aug. 19 4 A. F. Aug. 11 6 A F. Aug. 20 7 M. F. Aug. 26 1 A. F. Sept. 3 Noon N. Aug. 27 7 M. N. Sept. 4 9 M. N Sept. 9 Noon. N. Sept. 17 Midn. F. Sept. 10 7 M. F. Sept. 18 3 A. F. Sept. 25 1 M. F. Od. 3 4M. N. Sept. 25 5 A. N. Od. 4 1 M. N. O. 9 2 M. N. O. 17 O.M. F. O. 9 10 A. F. O. 18 2 M. E. OA. 24 11 M F. Nov. 1 8 A.N. OA. 25 2 M N. Nov. 2 4 A. N. Nov. 7 7 A N. Nov. 15 9 A. F. Nov. 8 4 A. F. Nov. 16 F. Nov. 22 9 A F. Nov. 31 11 M. N. Nov. 23 1 A. N. Dec. 2 5 M. Dec. 7 1 A N. Dec. 15 11 M F. Dec. 8 Noon F. Dec. 16 6 M. F. Dec 22 7M F. Dec. 31 Midn N. Dec. 22 11 A. N. Dec. 31 5 A. Anno 1761 | Anno 1763. Anno 1765. | Anno 1767. N. Jan. 6 9 M N. Jan. 14 4 M. F. Jan. 7 6 M. F. Jan. 15 1 M. F. Jan. 20 6 A F. Jan. 29 Noon. N. Jan. 21 11 M N. Jan. 30 3 M. N. Feb. 5 3 M. N. Feb. 12 10 A. F. Feb. 5 11 A. F. Feb. 13 8 A. F. Feb. 19 6M F. Feb. 27 11 A. N. Feb. 19 11 A. N. Feb. 28 N. Mar. 6 7 A. N. Mar. 14 5 A. F. Mar. 7 1 A. F. Mar. 15 F. Mar. 20 7 A. F. Mar. 29 8 M N. Mar. 21 1 A. N. Mar. 29 9 A. N. Apr. 5 8 M N. Apr. 13 10 M. F. Apr. 6 1 M F. Apr. 14 F. Apr. 19 8M. F. Apr. 27 4 A. N. Apr. 20 4M. N. Apr. 29 9M. N. May 4 6 A N. May 13 iM F. May 5 10 M F. May 13 9 A. F May 18 11 A F. May 27 1 M. N. May 19 8 A N. May 27 N. June 3 1 M. N. June 11 2 A F. June 3 5 A F. June 12 8 M. F. June 17 1 A F. June 25 11 M. N. June 18 11 M. N. June 26 7 M. N. July 2 8 M. N. July 10 Midn F. July 2 Midn F. July 11 15 A. F. July 17 4 M. F. July 24 11 A. N. July 18 2 M. N. July 25 8 A. and fuly 30 3 A. N. Aug. 9 8 M F. Aug. 1 7 M. F. Aug. 10 11 M. P. Aug. 15 107 A. F. Aug. 23 1 A N. Aug. 16 4 A. N. Aug. 24 Noon. N. Aug. 28 11 A. N. Sept. 7 5 A.F. Aug. 30 4 A.F. Sept. 8 10 M. F. Sept. 14 10 M. F. Sept. 22 05 M. N. Sept. 15 4 M. N. Sept. 23 4 M. N. Sept. 128 10 M. N. Oct. 7 1 M. F. Sept. 29 3 M. F. Oct. 7 4 A. F. Od 13 11 A. F. Od. 21 10 A. N. Od. 14 5 A N. Od. 22 9 A. N. O.A. 27 10 A N. Now 5 10 M. F. O.A. 28 6 A F. Nov. 6 2 M. F. Nov. 1) 1 N. F. Nov. 21 4 A.N. Nov. 13 4M. N. Nov. 21 N. Nov. 26 2 A. N. Dec. 4 9 A. F. Nov. 27 11 M F. Dec. 5 P. D.c. 11 11 A. F. Dec. 20 9 M. N. Dec. 12 N. D. 26 8 M. F. Dec. 27 3 A N. Dec, 21 4 M. HF. Dec. 27 (M. 79 4

79. A TABLE, shewing how many Hours and Minutes the Moon is past the Meridian, when it is high Water at the following Places, viz.

illes o o	H.	M.	0 0	H.	M.
A Berdeen	0	45	Bridgwater	7	30
St. Andrews	2	15	Cape Blanco	9	45
Amsterdam	3	0	Bulloign	10	30
Armentiers	3	0	C. Cautin (Barbary)	0	0
Abroth	5	15	Calice without	I	30
Antwerp	6	0	Camfere	I	30
Archangel	6	0	Coquet	3	0
Abermorich	6	0	Cork in Ireland	4	30
Amazon river	6	0	Cape-clear, Ireland	4	30
Aldborough	9	45	Caldy	5	15
Beach	0	0	Carnarven-bay	5	15
Bajador (Barbary)	0	0	Cromer	6	0
Blacktail Beacon	0	15	Caskets without	8	15
Blackness	I	30	C.Sierre-lion Guiney	8	15
Bell Isle	1	30	Chamberness	9	45
Berwick	I	30		10	10
Bluet without	2	15		10	30
Britain fouth coast	3	0	Calice-road	10	30
Bifcay coast	3	0	Calshot	II	15
Bourdeaux river	3	0	Dunkirk	0	0
Buchaness	St. Jak	0	Dover-port	0	0
Bona Esperance	33333	0	Port Defire, America	0	0
Breft	3	45	Downs	I	30
Bass without	3	45	Dundee	2	15
Bridlington	3	45	Denby	2 3	15
Bourdeaux R. within	13	45	Dort	3	0
Brovage without	3	45	Dunbar	4	30
Baltimore	4	30	Dungarven	4	30
Bree-found	46	30	Dartmouth	4 4 5 6	15
Blackney	6	0	St. David's-head		0
Briftol	6	0	Dublin, Ireland	8	15
Briftol-key	6	45	Diep	9	45
					Dun-

	H.	M.		H.	M.
Dunwick	0	45	Hague	8	15
Dover	10	30	Harwich	10	30
Elder	0	0	St. Helens	10	30
Elve	0	0	Jutland Isles	0	0
Enchuysen	0	0	Ireland W. coast	3	0
Edam	1	30	Ireland fouth coast	5	15
Edinburgh	4	30	John de Luce	10	30
Egmon	4	30	Kentish Knock	0	0
Exwater	7	36	Killiars	3	0
Entrance of the Eme	s 7	30	Kinfale, Ireland	4	30
Flushing	0	45	Kilduyn	7	30
Finmark-coaft	I	30	Lifbon	2	15
Fountnay without	2	75	St. Lucas	2	15
Flamborough-head	3	0	London	3	0
Fair Isle	3	0	Leith	4	30
Frith	.3	30	Lawreness	4	30
Falmouth	4	30	Lyn without	5	15
Foy	- 5	15	Lundey	5	15
Foulness	6	45	Lyn	4 5 5 6 6	0
Friefland-coast	7	30	Lanion		45
Florida, Carolina	7	30	Landsend	78	30
Foreland N. and S.	9	45	Lam-bay	8	
Gibraltar-road	0	0	Leostaff	9	45
Guernsey	0	45	Lenow	9	45
Goree	I	0	Maxe within	0	45
Gravefend	IJ	30	Malden	0	45
Gascoin	3	0	St. Mark	2	15
Groy	3	0	St. Matthew point	3	45
Gorend	11	15	Mount's-bay	4	30
Hever	0	0	Milford	5	15
Hern	0	0	Moonless	5	15
Holy Island	I	30	St. Maloes	5 8	15
Hartlepool	3	0	Magnes-found		15
Humber without	4 6	30	Ifle of Man	9	0
Hull		0	Marget-road	11	15
Hamborough	6	0	Newport(life-wigh		0
		The state of	A CONTRACTOR OF THE PARTY OF TH	- 27	Nore

PLAN	IS	PHERES.		63
H.	M.	AND SHOULD ADDRESS TO	H.	M.
Nore west-end o	0	Shetland	3	0
North C. Maggero 3	0	Scilly	3	45
Nantz-river without 3	0	Scarborough	3	45
Newcastle 5	15	Sound	3	45
St. Nicholas, Russia 6	45	Severn	4	30
Normandy coast 10	30	Stockton	4 5	15
Orkneys 3	0	Spurn	5	15
Orwel 9	0	Start	5	45
Orfordness 9	45	Seyn-head	10	30
Portfmouth o	0	Senegal	10	30
Poictou S. coast 3	0	Terveer within	0	45
	0	Thanet -	1	30
Pens 3 Porthus 3	0	Terveer without	1	30
Portugal coast 3 Plymouth 6	0	Tinmouth	3	0
	0	Tees-mouth	3	0
St. Powls 6	0	Tenariff	3	0
Podesemsk, Russia 6	45	Torbay	5	15
Portland 8	15	Texel	7	30
Peterport 8	15	Ureck	0	0
Picardy coast 10	30	Uſe	3	0
Quinborough o	0	Ushant without	6	0
Quebec Canada 6	0	Isle of Wight	0	0
Rebdan	45	Winchelsea	0	45
Rochester	45	Weilands	I	30
Rumney	30	Whitby	3	0
Ramkins	30	Waterford in Ireland	4	30
Rotterdam 3	0	Weymouth	6	0
Roan Rochell without Roan-river within Ramfey 3 3 3 7 7 7 7 8 7 7 7 7 7 7 7 7 7 7 7	45	Wells	6	0
Rochell without 3	45	Weymouth-key	6	45
Roan-river within 3	45	Wieringham	7	0
	15	Winterton	9	0
Rye 11 Rhodes 11	15	Youghall (Ireland)	4 8	30
OL C	15.	Yarmouth peer		15
0 1	0	Yarmouth-peer Yarmouth-town	9	0
0 :	0	7 1 1 0	9	45
Spits	0	Zicaland Coalt	I	30
Mark - L			80	A



81. ATABLE, shewing how many Hours and Minutes of Time, are equal to any Number of Degrees of Right Ascension.

	to any Ivu		regrees of		
D.R. Af.	Time,	D.R. Af.	Time.	D.R.AC	Time.
D.F	Н. М.	D.F	Н. М.	D.I	Н. М.
I 2	o 4 o 8	31 32	2 4 2 8	70 80	4 40 5 20 6 0
3 4 5	0 12 0 16 0 20	33 34 35	2 12 2 16 2 20	100	6 0 6 40 7 20
6	O 24 O 28	36	2 24 2 28	120	8 o 8 40
7 8 1	o 32 o 36	38	2 32 2 36	140 150 160	9 20
11	0 44	40	2 44	170	10 40
12	0 48	42 43	2 48 2 52 2 56	180 190 200	12 0
14	0 56	44 45	3 0	210	13 20
16 17 18	I 4 I 8 I 12	46 47	3 4 3 8 3 12	220	14 40 15 20 16 0
19	I 16 I 20	48 49 50	3 4 3 8 3 12 3 16 3 20	240 250 260	16 40 17 20
2 I 2 2	I 24 I 28	51 52	3 24 3 28	270 280	18 0 18 40
23	1 32 1 36	53 54	3 3 ² 3 36	290 300	19 20 20 0
25	I 40	55	3 44	310	20 40
27 28 29	1 48 1 52 1 56	57 58	3 48 3 5 ² 3 5 ⁶	330 340 350	22 0 22 40 23 20
30	2 0	39 60	4 0	360	24 0

82. A TABLE, shewing how many Degrees of Right Ascension, are equal to any Number of Hours or Minutes in Time.

- 3		12 27 27			1 2 2	122	_
Hours.	Deg.	Min. of	Tu	ne.	Min. of	111	ne.
Liouis	treg.	a Deg.	1	11	a Deg.	1	"
-	A 200	1000 700		-	Thomas	-	-
I	15	I	0	15	31	7	45
2	30	2	0	30	32	8	0
3	45 60	3	0	45	33	8	15
5	60	3 4 5	1	0	34	8	30
5	75	5	I	1.5	35	8	45
1	-	-			-		
6	90	6	1	30	36	9	9
7 8 9	105	7 8	I	45	37	9	15
8	120		2	.0	38	9	39
9	135	9	2	15	39	9	45
10	150	10	2	30	40	10	0
11	165	11	2	15	41	10	7.5
12	180	12		45	41	10	15
13	195	13	3 3	15	42	10	39
14	210	14	3		43	11	45
15	225	15	3	30	44	11	1 1000
-3			3	45	45		15
16	240	16	4	0	46	11	39
17	2.55	17	4	15	47	11	45
18	270	18	4	30	48	12	Q
19	285	19	4	45	49	12	15
20	300	20	5	0	50	12	30
-			-				
21	315	21	5	15	51	12	45
22	330	22	5 5 5 6	30	52	13	0
23	345	23	5	45	53	13	15
24	360	24		0	54	13	30
1 . 0	1 050	25	6	15	55	13	45
		06	-6		-		
1	1 3	26	6	30	56	14	0
198	1 123	27	6	45	57	14	15
1	TO THE LOCAL COMMENTS	28	7	0	58	14	30
FIRE DA		29	7 7	15	59	14	45
-		30	7	30	1 60	114	0

33. A TABLE, shewing the Sun's Right Ascension and Declination for the three first Signs, V, &, H.

Note, For 55, 9, 10, subtract the Right Ascension from 12 Hours, For 2, 10, 2, add to the Right Ascension 12 Hours, For 19, 20, 11, subtract the Right Ascension from 24 Hours.

	One	e Sign.	12	T	wo	Sign	s.	T	hree	Sign	ns.
	R. Afc	2011	lin.	Charles and	Afc.		lin.		Afc.	De	clin.
13	H. M		1	H.	M.	0	1	H.	M.	0	1
1	1	4 0	24	1	55	11	50	3	55	20	23
2	7	7 0	48	I	59	12	II	4	0	20	36
3	1 100	ı i	11	2 2	3	12	32	4	4	20	47
4	0 1		35		7	12	52	4	. 8	20	59
5 6	0 1		59	2	II	13	12		12	21	10
1 POSTER	0 2		23	2	15	13	32	4	16	21	20
1 8	0 2		47	2	19	13	52	4	21	21	31
Contract of	0 2	74 P.	10	2	23	14	21	4	25	21	40
9	0 3		34	2	26	14	31	4	29	21	50
10	0 3	CO DO COURT	58	2	30	14	50	4	33	21	59
11	0 4		21	2	34	15	9	4	38	22	8
12	0 4	4 4	45	2	38	15	27	4	42	22	16
13	0 4	8 5	8	2	42	15	46	4	46	22	24
14	0 5		32	2	46	16	4	4	51	22	31
15	0 5		55	2 2	50	16	21	4	55	22	38
16	0 5		18		54	16	39	4	59	22	44
17		3 6	41	2	58	16	56	5	3	22	50
18		6 7	4	3	6	17	13	5	20	22	56
119		0 7	27	3		17	30	5	12	23	6
20		4 7	50	3	10	17	46	5	7.0 (65.4	23	2000
21		8 8	12	3	14	18	2 18	5	21	23	10
22		6 8	35	3		18		5	25	23	14
23			57	3 3 3 3 3 3 3 3	22 26	18	30	5	29 34	23	17 20
24		9 9	19	3	20		40	5	38	23	20
25	I 3	3 9 6 10	41	3	30 35	19	3	5	43	23	25
27	1 4	010	3	3	35	10	21	5		22	26
28	1 4	4 10	25 46	3	39	19	31 45	5	47 51	23	27
29	1 4	8 11	8	3	43 47	19	51	5	56	22	23 25 26 27 28
30		2 11	29	3	51	20	II	5 5 5 5 5 5 5 5 6	0	23 23 23 23	281
130	1 5	-111	-9	1 3	01			11		1-3	203

84. A TABLE, shewing the Latitudes and Difference of Metidians from the Royal Observatory at Greenwich, of the following Places, viz.

Names of Places.	D	if.IV.	ler.		A	lt. ó	f P	ole.
I AMES OF TEACES.	h.	m	. f.	100	d.	m.	f.	
Alexandria in Egypt	2	1	6	A.	31	7	0	N.
Athens in Greece	1	35	30	A.		5		The same of the same of
Babylon in Egypt, Grand Cairo	2			A.		2	30	
Berlin in Brandenburg	0	53	50	A.	52	33	0	
Cape of Good Hope in Africa	1	8	0	A.	34	15		
Cadiz in Spain	0	24	28	S.		33		
Carthagena in America	5	0	46	2		26		
Copenhagen in Denmark	0		0		55	40	45	N.
Dantzick in Poland	I			-	54	22	0	N.
Island of St. Helena	0	24		S.	15	55	0	S.
Jerufalem	2	21		A.		55		
London at St. Paul's	0	0	2-12-15/0	S.	-	30	- 1711	
Lifbon in Portugal	0	36	CONTRACTOR OF THE PARTY OF THE	S.	COTTO DE LA COMPANSION	42	-	
Mofcow in Ruffia	.2	41	20	A.	-	36		man man
Naples in Italy	0	5.8	40	A.		50		
Observatory at Greenwich	0	0	1211/200	-		28	30	N.
Observatory at Paris	0	.9		A.	77 740	50	7	
Oxford	0	25	4	S.		45		
Petersburg in Russia	2	EI	SECTION S	A.		0		
Pekin in China	7	45	20	A.	39	54	0	N.
Porto Bello in America	5	19	20	15-3-17		133	100	TO THE
Prague in Bohemia	0	59	0	A.	40	14	30	N.
Rome, the Metropolis of Italy	0	50	0	A.		54		N.
Rochelle in France	0	5	4	S.	and the same of	29		N.
Smyrna in Turkey	1	49	19	W		28		N.
Teneriff Mountain	I	6	12	S.	28	23	27	N.
Terra del Gada in Madagascar	2	58	0.	A		_		1 100
Tornea in Lapland	I	35	15	A.		50		
Venice in Italy	10	48	18			25		
Vienna in Austria	I	5	30	A.		12		
Upfal in Sweden	11	11				51		

85. A TABLE of Horizontal Distances for Horizontal Dials, shewing also the Distance of each Hour-line from the Meridian upon direct North or South Planes; calculated to every Degree of Latitude.

1	E 15	1	17	1 - 7	1.47	154.	Hay	-114	1111		18			
1	An	Ho	urs.	Ho	urs.	H	ours.	H	ours.	H	ours.	H	ours.	D
110313	2	XI.	1;				111.						1.	dir ial,
	2	1014	1 434	1	-	-	-	-			-			- 00
79733 55 5575	27.	1		-				-			-			St South
29.99	100	d.	m.	d.	m.	d.	m.	d.	m.	d.	m.	d.	m.	South titude.
146.	al	100	00	1.	00				185	17	211	1	+ 18	tb de.
-		100	100	46	10	0 "	101			-	ZIL	15		
4	0	0	0	0	0	0	0	0	0	0	0	0	00	90
	1	0	16	O	34	I	0		44	3	44	12000		89
1	2	0	32	1	9	2	. 0	1	27	7	3 27113	90	3/1	88
1	3		48	I	44	3	0		II	II		90	P-3000	87
-	-4	I	5	2	19	4	0	1	28.80	14	36	90	600	86
3	- 5	I	20	2	52	4	58	8		18	4.30 307 1.3	90		85
-	4 5 6	I	36	3	27	5	58	10		21		90		84
1		2	52	4	3	6	57	II	W 24 W 25	24	170 30 30	90	The second second	83
Ť	7 8	2	8	4	37	7		13		27	and the second	90	A COLUMN TO SERVICE AND ADDRESS OF THE PARTY	82
1	9		23	5	9	8		15	10	1.00	1000	90	82. C. C.	Sı
I	10		40	5	43	9		16		32	0.0	90		70
H	11	2	55	6	17	10	48	81	17	The state of	100	90		79
1	12	3	11	6	51	II		19		37		90	00	78
H	13	100	27	7	100 100	12		21	17	10		90		77
H	14	W27%	53	7	57	13	26	22	44	1		90		76
П	15	The state of the s	58	8	- F-51-51-7	14		24	9			90		75
П	16	4	13					25	- B-0.100	+5		90		74
1	17	10000	29	9	25	15		26	100	47	26	90		73
1	18			10	8	17		28		49		90		72
1	19	100000000000000000000000000000000000000		10		18		29		50		90		71
1	20		71 10000	II	10	18		30		51		90		60
1	21	1		11		19		31	50			90		69
	22	F 100 100 100 100 100 100 100 100 100 10	777.779	12		20		32		54		90		68
9	23			12		21		34		55		90		67
1	24		1000	13		22	8	35	10	244		90		66
1	25	1 /	28	13		22	55	35		57	-	90		65
1	26	1		14		23		37		58	13777	90		64
1	-	-	-	The state of	-	7	-	10/	-	W-2	27	2	-	-

AT	BLE	of	Horizontal	Distances,	Ec.
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1				3					136		Shi	L w	MIL R	nith i
D	An	Ho	urs.	Ho	urs.	Ho	ours.	Ho	urs.	Ho	urs.	Ho	ours.	DA
Dial,	-	XI.	I.	x.	II.	IX.	III.	VIII	.IV.	VII.	v.	v	I.	A dir Dial,
	17	-		-	-	-		-	-		10000	0.1	-	
Latitude	Harizantal			,			200	,		,		,	1	
ud	nta	d.	m.	d.	m.	d.	m.	d.	m.	d.	m.	d.	m.	South
10.			Com.									-		200
2	7	6	56	14	41	24	25	38	11	59	27	90	00	63
	8	7	10	15	10	25	9	39	7	60	17	90	00	62
	9	7	24	15	40	25	52	40	2	61	4	90	00	61
	ó	7	58	16	6	26	33	40	54	61	49	90	00	60
	I	7	50	16	34	27	15	41	44	62	30	90	00	59
	2	8	5	17	1	27	55	42	30	63	II	90	00	58
	3	8	19	17	27	28	34	43	20	63	49	90	00	57
	4	8	31	17	54	29	13	44	5	64	24	90	00	56
1 3	5	8.	44	18	20	29	50	44	49	64	58	90	00	55
3	6	8	57	18	45	30	. 27	45	31	65	30	90	00	54
3	7	9	10	19	9	31	2	46	12	66	10	90	00	53
3	8	9	22	19	34	31	37	46	50	66	29	90	00	52
	19	9	34	19	58	32	11	47	28	66	56	90	00	51
1 4	0	9	45	20	2.1	32	44	48	7	67	2.1	90	00	50
	I	9	57	20	44	33	16	48	39	67	47	90	00	49
4	2	10	10	21	7	33	46	49	12	68	II	90	00	48
	3	10	22	21	29	34	18	49	44	68	33	90	00	47
	4	10	32	21	51	34	47	50	10	68	54	90	00	46
	5	10	43	22	12	35	17	50	46	69	15	90	00	45
	.6	10	54	22	23	35	44	51	15	69	35		00	Contract to
1 4	8	11	5	22	53	30	II	51	42	69	53	90	00	STATE OF THE STATE
1 4	0	II	5 17 25 35	23	53 13 33	36 36 37	37 3 28	52 52	9 35 00	70	53 11 28 43 59 13 28	90	00	42
	9	11	25	23	33	37	3	52	35	70	28	90 90	00	41
1 5	0	11	35	23	52	37	20	53	00	70	43	90	00	40
1 5	2	II	45	24	9	37	52	53	16	70	59	90	00	39
1 5	1	11	22	24	47	28	15	53	40	71	13	90	00	38 37
1 3	3		13	25	43	28	5/	54	20	71	20	90	00	37
1 5	4	12	45 55 5 13 22	25	5 ² 9 27 43 2 18	37 37 38 38 38 38	52 15 37 58 19 40	53 54 54 54	24 46 18 29 49 9	70 70 71 71 71 71 71 72	41	90 90	00	36 35
1 3	5	12	32	25	34	39	19	55	49	72	54	90	00	35
1 5	7	12	40	25	50	39	50	55	28	72	17	90	00	34
1 3	8	12	48	26	50	40	59 81	55	45	72	41 54 5 17 28		00	33
1 3	0	-	401	2.0	2	140	-	133	43	1/-	20	190	00	32

	tal		AT	ABI	.E 0	f H	oriz	onta	l Di	istan	ices,	છ	c.	D
Dial,	An Ho	Ho xI.	urs.	Ho x.	urs.	1	urs.	THE PARTY OF	urs.	100 AND			ours.	A direst Dial, La
Latitude.	Horizontal	d.	m.	d.	m.	d.	m.	d.	m.	d.	m.	d.	m.	Latitude.
	59	12	46		20	40	36	56	3	72	38 48	90	00	31
	60	13	4	26		40	54	56	19	72	58	30	00	30
1	61	13	19	27	47 I	41	21	56	34 49	72 73	7	90	00	29 28
	63	13	26	27	13	41	42	57	3	73	15	15.000	00	27
	64	13	32	27	25	41	57	57	17	73	24		00	26
	65	13	39	27	37	32	15	57	30	73	32	90	00	25
	66	13	46	100000000000000000000000000000000000000	49		25	57	43	73	39	90	00	24
1	67	13	51	27	59	42	38	57	54	73	46	90	00	23
1	68	13	57	27	9	42	50	58	5	73	53	1000	00	22
1	69	14	3	28	19	43	2	58	16	73	59		00	21
	70	14		28	29	43	13	58	26	74	05		00	20
-	71	14	18	28	27 46	43	1000	58 58	35	74	16		00	19
1	72	14	22	28	54		24 36	58	44	A COLUMN TO A COLU	20	90	00	17
	73	14	27	28	2	43	52	59	52 €0	74	25		00	16
	74 75	14	30	29	17	44	00	59	7	74	30	90	00	15
	76		33	0.0		44	8	59		74		90	00	14
1		14	37	29	21	44	14	59	22	74	37	90	00	13
-	77	14	41	29	27	44	22	59	27	74	41	90	00	12
1	79	14	44	29	27 32	44	28	59	27 32	74 74	44	90	00	II
-	80	14	47	29	37	44	34	59	37	74	45	90	00	10
1	81	14	49	29	41	44	37	59	40	74	49	90	00	9
1	82	14	51	29	45	44	40	59	44	74	51 53	90	00	8
-	83 84	14	53	29	49	44	44 48	59	47	73	53	90	00	76 5 4 3
	84	14	55 56	29	52	44	48	59	51	74	55	90	00	5
-	85 86	14	50	29	54	44	53	59	54	74	57 58 58	90	00	3
-	87	14	57 58	29	55 56	44	55 56	59	55 56	74	50	90	00	7
1	8 ₇ 88	14	59	29	57	45	57	59 59	58	74 74	59	90	00	2
1	89	14	59	29	57 58	44	57 58		59	74	59	90	00	I
1	90	15	00	30	00	45	00	59	00	75	00	90	00	0

86. 4	TABL!	E of Ang	les which	every	Rhomb (Meridian.	or.	Point	of
1		ormpuje)		12 2120				-

	or compays		and the same			
NO RTH.	SOUTH	Loint.	D.	M	NORTH	SOUTH,
N. by East	S. by East.	1	02 05 08 11	49 37 26 15		S. by W.
N. N. E.	S. S. E.	2	14 16 19 22	04 52 41 30	011 45	S. S. W.
N. E. by N.	S. E. by S.	3	25 28 30 33	19 07 56 45	N.W. by W.	S. W. by S.
North-East.	South-Ea.	4	36 39 42 45	34 22 11 00	Nor. Weft.	Sou. West.
N. E. by E.	S. E. by E.	5	47 50 53 56	49 37 26 15	N.W.byW.	S. W. by W.
È. N. E.	E. S. E.	6	59 61 64 67	04 52 41 30	W. N. W.	W. S. W.
E. by N.	E. by S.	7	70 73 75 78	19 07 56 45	W. by N.	W. by S.
East.	East.	8	81 84 87 90	34 22 11	West.	West.



87. A TABLE, shewing what Declination the Sun must have, when the Day is any Number of whole Hours longer or shorter than 12 Hours in any Latitude.

-	inan s	4 11	oursi	n un	Laste	i uuc						
Lat.	Diff.	ı H.	Diff.	2 H.	Diff.	3 H.	Diff.	4 H.	Diff.	5 H	Diff.	6H.
-	o D	ecli.	o D	ecli.	o D	ecli.	o D	ecli	o D	ecli.	o D	ecli.
	D.	M.	D.	M.	D.	M.	D.	M.	D.	M.	D.	M.
89	0	16	0	16	0	23	0	30	0	36	0	43
87	0		0	31	0	46	1	30	I	50	2	25
86	0	32	1	47	i	56	2	co	2	27	2	50
85	0	39	I	18	-1	55	2	3	3	03	3	32
84	0	47	1	27	2	18	3	00	3	39	4	14
83	0	55	1	37	2	42	3	31	4	15	4	56
82	1	3	1	46	3	05	4	01	4	52	5	39
81	1	11	1	55	3	29	4	32	5	29	6	22
80	1	19	2	37	3	52	5	03	6	97	7	06
79	1	27	2	47	4	16	5	34	6	45	7	48
79 78	1	35	3	07	4	40	5	05	7	22	7 8	31
771	I	43	3	16	5	04	6	36	8	00	9	14
76	-1	51	3	37	.5	28	7	07	8	3.8	9	57
75	1	59	3	58	5	51	7	38	9	16	10	40
74	2	8	4	15	6	16	8	10	9	54	11	25
73	2	17	4	32	6	41	8	43	10	33	12	10
72	2	26	4	49	7	06	9	15	11	12	12	56
71 70	2 2	36	5 5	23	7 7	32 56	9	47	12	30	13	42 26
	- 10	43	-	-3	3 3 3 3 3	033	01		-	5	-	-
68	3 3 3 3	152	5	51	8	22	10	52	13	05	15	11
67	3	0 1	5	59	8	49	11	.26	13	41	15	57
66	3	10	6	35	9	15	12	59	14	52	17	43
65	3	29	6	53	9 9 10	07	13	07	15	52 28	18	15
67 66 65 		-	£ 310							_		
62	3 3 4 4	38 48 57 8	7	I Z	10	35	13	43	16	15 02	19	02
62	3	40	7	40	11	32	14	19	17	40	19	49
61	4	8	7 8 8	59	12	01	14	55	17	361	20	37
60	4	17	8	30	12	28	16	06	19	36	22	12
59 58 57 56 55	4	28	8	51	12	58	16	44	20	07	23	02
58	4	39	9	51	13	58	17 18 18	23	20	51	23	50
57	4	52 I	9	33	13	59	18	01	21	51 36 21		
50	4 4 5 5	I	9	54	14	30		40	22			
33	5	83	10	17	15	co	19	18	23	05	-	
11	p.S.Green	SOUTH STREET	Authorities 20	2000	20/20/2	-		-	-	-	-	-

1	A TABLE, shewing what Declination the Sun must have, &c															
A	Diff. 1 H. Diff. 2 H. Diff. 3 H. Diff. 4 H. Diff. 5 H. Diff. 6 H.															
Lat.		Diff.	н.	Diff.	2 H.	Diff.	3 H.	Diff.	4 H.	Diff.	5 H.	Diff.	6 H.			
-		o D	ecli.	o D	ecli.	o D	ecli.	o D	ecli.	o D	ecli.	o D	ecli.			
=		D.	M.	D.	M.	D.	M.	D.	M.	D.	M.	D.	M.			
3 3 3 3 3	3210198 76 5 4 321 0 98 76 5 4 32 1 0 98 76 5 4 3	D. 55566 6677 78888 8 9 9 9 10 10 11 11 12 12 13 13 14 15 16 17 17	M. 25 38 50 02 15 30 42 56 17 26 43 00 17 34 50 10 29 49 12 34 15 48 22 57 39 20 05 54	10 11 11 12 12 13 13 14 14 14 15 16 16 17 18 19 20 21 22 22 23	M. 40 04 27 51 15 42 09 37 04 31 02 33 04 36 09 47 24 02 39 17 09 01 53 44	15 16 16 17 17 18 19 20 20 21 22 23 23 23 16 ta th an re	33 97 40 14 48 26 93 40 18 57 40 23 96 23 66 23 66 23 66 23 66 66 67 68 68 68 68 68 68 68 68 68 68	ow we must given in a siven at the	hat have Diff Les ours	Dec we we mbe in Later ingth in fin cop of	clina hen ers of od the of the	tion the le; f between the lis Die Ta	the Day ours inft, reen Day offe-ble.			
10	2 2 3 7	18 19 20 21 23 24	46 46 53 07 26			under which, in the same Co- lumn, opposite to the Latitude of the Place, is the Declination required.										

88. A TRAVERSE-TABLE.

Din	1 4 F	oint.	1 1 P	oint.	3 P	3 Point.		I Point	
Distance.	Lat.	Dep.	Lat.	Dep.	Lat.	Dep.	Lat.	Dep.	
1 2 3		0 0	1 0 2 0 3 0	0 1 0 2 0 3	1 0 2 0 3 0	0 1 0 3 0 4	1 0 2 0 3 9	0 2 0 4 0 6	
4 5 6 7 8	4 0 5 0 6 0 7 0 8 0	0 2 0 2 0 3 0 3 0 4	4 ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° ° °	0 4 0 5 0 6 0 7 0 8	4 º 4 9 5 9 6 9 7 9	0 6 0 7 0 9 1 0 1 2	3 9 4 9 5 9 6 9 7 8	0 8 1 0 1 2 1 4 1 6	
9 10 15 20 25	9 0 10 0 15 0 20 0 25 0	0 4 0 5 0 7 1 0 1 2	9 0 9 9 14 9 19 9 24 9	0 9 1 0 1 5 2 0 2 4	8 9 9 9 14 8 19 8 24 7	1 3 1 5 2 2 2 9 3 7	8 8 9 8 14 7 19 6 24 5	1 8 1 9 2 9 3 9 4 9	
30 35 40 45 50	30 0 34 9 39 9 44 9 49 9	1 5 1 7- 2 0 2 2 2 4	29 8 34 8 39 8 44 8 49 8	2 9 3 4 3 9 4 4 4 9	29 7 34 6 39 6 44 5 49 5	4 4 5 1 5 9 6 6 7 3	29 4 31 3 39 2 44 1 49 0	5 8 6 8 7 8 8 8 9 8	
55 60 70 80 90	54 9 59 9 69 9 79 9 89 9	2 7 2 9 3 4 3 9 4 4	54 7 59 7 69 7 79 6 89 6	5 4 5 9 6 9 7 8 8 8	54 4 59 3 69 2 79 1 89 0	8 1 8 8 10 3 11 7 13 2	53 9 58 8 68 7 78 5 88 3	10 7 11 7 13 7 15 6 17 6	
100 200 300 400 500	399 6	4 9 9 8 14 7 19 6 24 5	99 5 199 0 298 5 398 0 497 5	9 8 19 6 29 4 39 2 49 0	98 9 197 3 296 7 295 6 494 5	14 7 29 4 44 1 58 8 73 5	98 1 196 2 294 3 392 4 490 5	19 5 39 c 58 5 78 0 97 5	
Diftan.	Dep. 7 3 P	Lat.	Dep. 7 ½ P	Lat.	Dep. 7 4 F	Lat.	Dep.	Lat.	



A TRAVERSE TABLE.

Din	2 1/4	Point.	2 1]	Point.	2 3	2 3 Point.		oints.
Diftance.	Lat	Dep.	Lat.	Dep.	Lat.	Dep.	Lat.	Dep.
1 2 3	0 9 1 8 2 7	0 4 0 9 1 2	0 9 1 8 2 6	0 5 0 9 1 4	0 9 1 7 2 6	1 0	0 8 1 7 2 5	06
4 5 6 7 8	3 6	1 7	3 5	1 9	3 3	2 1	3 3	2 2
	4 5	2 1	4 4	2 4	4 2	2 6	4 2	2 8
	5 4	2 6	5 3	2 8	5 0	3 1	5 0	3 3
	6 3	3 0	6 2	3 3	6 8	3 6	5 8	3 9
	7 2	3 4	7 1	3 8	6 6	4 1	6 6	4 4
9	8 1	3 8	7 9	4 2	7 5	4 6	7 5	5 0
10	9 0	4 3	8 8	4 7	8 3	5 1	8 3	5 6
15	13 6	6 4	13 2	7 1	12 5	7 7	12 5	8 3
20	18 1	8 5	17 6	9 4	17 6	10 3	16 6	11 1
25	22 6	10 7	22 0	11 8	21 7	12 8	20 7	13 9
30	27 I	12 8	26 5	14 1	25 7	15 1	24 9	16 7
35	31 6	15 0	30 9	16 5	30 0	18 0	29 1	19 4
40	36 2	17 1	35 3	18 9	34 3	20 6	33 3	22 2
45	40 7	19 2	39 7	21 2	38 6	23 1	37 4	25 0
50	45 2	21 4	44 1	23 6	40 9	25 7	41 6	27 8
55	49 7	23 5	48 5	25 9	47 2	28 3	45 7	30 6
60	54 2	25 9	52 9	28 3	51 5	30 8	49 9	33 3
70	63 3	29 6	61 7	33 0	60 0	36 0	58 2	38 9
80	72 3	34 2	70 5	37 7	68 6	41 1	66 5	44 4
90	81 4	38 5	79 4	42 5	77 2	46 3	74 8	50 0
100		42 7	88 2	47 1	85 7	* 51 3	83 1	55 6
200		85 4	176 4	94 2	171 4	102 6	166 2	111 2
300		128 1	264 6	141 3	257 1	153 9	249 3	166 8
400		170 8	352 8	188 4	342 8	205 2	332 4	222 4
500		213 5	441 0	235 5	428 5	256 5	4 ¹ 5 5	278 0
Distan.	Dep.	Lat.	Dep.	Lat.	Dep.	Lat.	Dep.	Lat.
0	5 3 P	oint.	5 1 Point.		5 ‡ P	oint.	5 Poi	ints.

A TRAVERSE TABLE.

Dift	3 4 1	Point.	3 1]	Point.	3 ¾ I	3 3 Point.		ints.
Diftance.	Lat.	Dep.	Lat.	Dep.	Lat.	Dep.	Lat.	Dep.
1	08	00	08	06	07	07	07	07
2	10	1 2	15	1 3	1 5	1 3	1 4	1 4
3	2 4	1 8	2 3	19	2 2	2 0	2 1	2 1
4	3 2	29	3 1	2 5	30	2 7	28	28
5	4.0	30	30	3 2	3 7	3 4	3 5	2000 200
5 6	4 8	3 6	46			40	4 2	4 2
8	56	4 2 4 8	5 4	4 4	4 2	4 7	49	4 9
8	6 4	4 8	6 2	5 1	5 9	5 4	5 7	5 7
9	7 2 8 0	5 4	70	2 7	67	60	6 4	6 4
10	and the second s		77	6 3	7 4	6 7	7 1	7 1
15	12 0	8 9	11 6	9 5	11 1	10 1	106	10 6
20	16 1	11 9	15 5	12 7	14 8	13 4 16 8	14 1	14 1
25		- 4 9	-9 3	15 9	18 5	10 0	17 7	17 7
30	24 1	17 9	23 2	190	22 2	20 1	21 2	21 2
35	28 1	20 8	27 0	22 2	25 9	23 5	24 7	24 7
40	32 I 36 I	23 8 26 8	30 9	25 4	29 6	26 9	28 3	28 3
45	40 2	29 8	34 8	28 5	33 3 37 0	30 2 33 6	31 8	31 8
1-				3. /	3/0	33 6	35 3	35 3
55	44 2	32 8	42 5	34 9	40 7	36 9	38 9	38 9
60	48 3	35 7	46 4	38 1	44 5	40 3	42 4	42 4
70	56 2	41 7	61 8	44 4	51 9	47 0	1 49 5	49 5
90	72 3	53 6	69 6	51 7	59 3 66 7	53 7	56 6	56 6 63 6
-								
100	80 3	59 6	77 3	63 4	74 1	67 2	70 7	70 7
200		119 2	154 6	126 8	148 2	134 4	141 4	141 4
	321 2	238 4	231 9 309 2	190 2 253 6	222 3	201 6	212 1	212 1
	401 5	298 0	386 5	317 0		268 8 336 o	282 3 353 5	232 8
-							222 2	353 5
an.	Dep.	Lat.	Dep.	Lat.	Dep.	Lat.	Dep.	Lat.
Diftan.	131	Point	121	Point	4 ! D	oins		
-	4 3 Point. 4 2 Point. 4 4 Point. 4 Points.				oints.			

89. A TABLE, shewing the Comets Nodes, their Perihelia, and Log. of their Perihelion Distances from the Sun; the Radius of the Earth's Orbit being 1, &c.

1	Noc	les .	Pe	er.	Diff	. a 💿	
Yrs.	S.	0	S.	ø	L	og.	Motions.
1337	2	24	I	8	9	609	Retrog.
1472	9	11	I	15	9	734	R.
1531	í	19	10	1	9	753	R.
1532	2	20	3	21	9	706	Direct.
1556	5	25	9	9	9	666	D.
1557	0	25	4	9	9	263	R.
1580	0	25 18	3	19	9	775	D.
1585	1	7	0	9	10	038	R.
1590	5	15	7	7	9	760	R.
1593	5	14	5	26	9	950	D.
1596	12	12	7	18	9	710	R.
1607	1	20	10	2	9	768	R.
1618	2	16	0	2	9	580	D.
1652	2	28	0	28	9	928	D.
1661	2	22	3	26	9	651	D.
1664	2	21	4	11	10	011	R.
1665	7	18	2	12	9	027	R.
1672	9	27	1	17	9	843	D.
1677	7	26	4	17	9	448	R.
1680	9	2	8	22	7	787	D.
1682	1	21	10	3	9	765	R.
1683	5	23	2	25	9	748	R.
1684	5 8	28	7	29	9	982	D.
1686	11	20	2	17	9	541	D.
1698	8	27	9	1	9	839	R.
1699	4	21	7	2	9	871	K.
1702	6	9	4	19	9	810	D.
1706	0	13	2	12	9	629	D.
1707	I	22	2	20	9	934	D.
1718	4	8	4	1	10	011	R.
1723	0	14	1	13	10	000	R.
1729	10	10	10	23	10	629	D.
1737	7	16	10	26	9	348	D.
1739	0	27	3	12	9	828	R.
1742	6	18		7	9	884	R.
1743	2		7 3 6	3	9	921	D.
1744	1	.15		17	9	347	D.
1747	10	27	9	7	10	342	R.
1748	7	22	7	. 5	9	924	R.

Nate, If a Comet of any Year has its Node, Perihelion, Perihelion Distance from the Sun, and Motion, nearly agreeing with another Comet of another Year, it is probable that they are one and the same, and therefore may be expected to return again, after the fame Interval of Time.

Thus have I introduced among the aforegoing Tables, fuch as are absolutely necessary in every Treatise of the Globes themselves, in order to come to any nice Determination by them, notwithstanding which all fuch Treatifes are filent on this Head, as if those Instruments wanted no fuch Correction, whereas nothing is more certain than that if the Globe be constructed ever so truly, the Days of the Month must be fixed for some particular Year, and toward the End of every four Years, it shews the Sun's Place erroneous by a Degree very nearly, and all the other Particulars depending thereon proportionably falle. And therefore our Planispheres must have been subject to the same Ertors, had we not substituted these Tables, not of Necessity, but as a necessary Supplement for using both of these Instruments in an exact Manner; the Constructions, Properties and Uses of which, will be readily understood by such as shall have duly considered the aforegoing geographical and aftronomical Definitions, and are tolerably well verfed in the Doctrine of the Sphere. Wherefore we shall not here exemplify the Uses of these Tables, seeing they appear at Sight, but proceed to fuch Particulars as may be most necessary to be observed in this Work. And here it is to be noted, that as the Projection of our Universal Planispheres is on the Plane of the Meridian after the stereographic Manner, the Eye being supposed to be placed in the Superficies of the Sphere, at the equinoctial Points, Aries and Libra; fo the Peculiarity of our Invention confifts wholly in the Discovery of a transparent Slider, Index or Projector, capable of forming inftantaneously, of Arches circular, all Kinds of spherical Triangles, such as are usually drawn by Scale and Compasses, in the Science of spherical Geometry. This Slider, Index or Projector may be rendered transparent a great Variety of Ways, the best that have occurred to us, and whereby, also, should the Transparency be lost, it may be soon restored, is as follows.

Take a Spoonful of Nut-oil, Linseed-oil, or Sweet-oil, and make it warm over the Fire. Place the Slider, on a smooth clean Table, or with a Sheet of clean Paper under to keep off the Dust. Dip a soft Brush, clean Feather, clean Linnen Cloth, or your Finger into the Oil, and oil the Slider over on both Sides, until no whitish Spots be seen. If too much Oil be laid on, wipe it off with a clean Linnen Cloth. Then warm the Slider well before the Fire, and set it aside, either between two Sheets of clean soft Paper, or without, until it is dry.



A DESCRIPTION of the EASTERN TERRESTRIAL PLANISPHERE.

The Line is drawn from one Side of the Circumference to the other; representing the Axis of the World, about which the diurnal Motion of the Earth is performed from West to East, or the apparent Motion from East to West in 24 Hours. The Extremities of this Line represent the North and South Poles of the Earth and Heavens, and on this Line are marked the Degrees of Latitude from the Centre to each Pole. This Line also represents half the equinoctial Colure, passing through the Beginning of Aries.

At right Angles to this Axis, and through the Centre, a right Line is drawn from one Side of the Circumference to the other, representing half the terrestrial and celestial Equator. The Circumference of the Planisphere is divided into 360 equal Parts for Degrees of Latitude, and numbered on one Side with 10, 20, &c. from the Poles to the Equator; this Circle is a first Meridian from which the Longitude of Places may be reckoned, and is placed here for giving a continued View of

the great Continent of Europe, Afia and Africa.

91. At the Distance of $23\frac{1}{2}$ Degrees on each Side of the Equator, both in the Circumference and polar Axis, half the Tropics of Cancer and Capricorn are drawn, shewing the Sun's greatest North and South Declination or approach toward the Poles; in like Manner, at the Distance of $23\frac{1}{2}$ Degrees from either Pole, the two Semi polar Circles are drawn, shewing those Places where there is either no Day or no Night when the Sun is in either of the Tropics.

Through the Centre, and from one Extremity of the northern Tropic to the opposite Extremity of the southern One, a right Line is drawn, representing half the Ecliptic, containing the three southern Signs, Capricorn, Aquarius and Pisces; also, the three northern Signs, Aries, Taurus and Gemini, with the Day and Month corresponding to each Degree and Sign, for

shewing the Sun's Declination, and to what Places he is vertical on any Day at Noon, with his right Ascension, and that Part of the Equation of Time depending on the Difference between

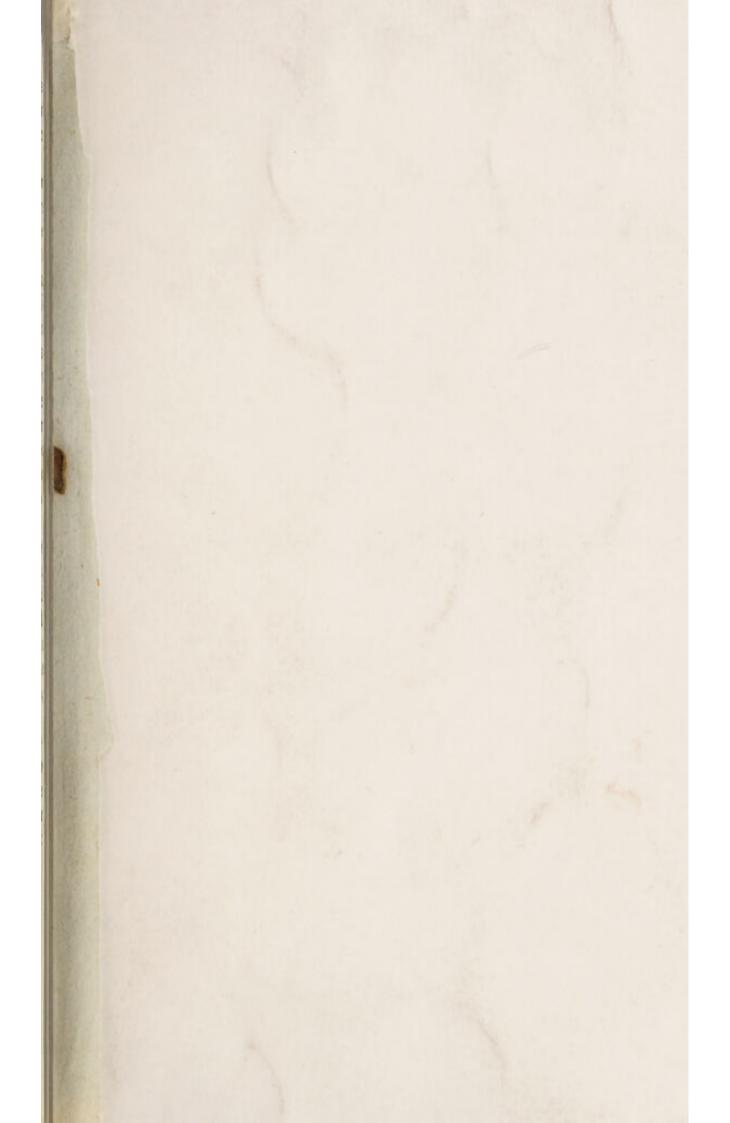
his right Ascension and Longitude.

at the Distance of a Degree from each other. And from one Side of the Planisphere to the other, at the Distance of a Degree from each other, are drawn all the Semi Parallels of Latitude from the Fquator to the Poles; the former for shewing the Difference of Longitudes or Meridians, and the latter the Latitudes of Places, and Zenith Distance of the Sun at any given Place on the Planisphere at Noon. At the North Pole is placed a Semi bour Circle, shewing the Difference of Meridians in Hours from London, and at the South Pole is placed another Semi hour Circle, shewing the Difference of Meridians in Hours from the first Meridian. These Hours are also placed at the Equinottial. Near the Side of the Planisphere, is placed a Scale of the Months and Days corresponding to those Degrees of Declination, opposite to them in the Side of the Planisphere.

93. On the Surface of this Planisphere is delineated the great Continent of Europe, Asia and Africa; the several Empires, Countries, Kingdoms, Islands, Rivers, Mountains, Cities, and remarkable Productions of this Part of the known World. Within the Indian or eastern Ocean, are shewn the Directions of the Monsoons or periodical Winds in those Seas, the Months when they blow, also the Trade-winds in the Ethiopic Ocean,

according to the Direction of the Darts.

Over the Ethiopic and Indian Oceans, are drawn the Variation Lines of the magnetic Needle, as they were observed, Anno 1700, Anno 1744, and Anno 1756, the former of which Lines are marked with Italic Letters, the next with Roman, and the latter with Italic Capitals, for shewing the Variation of the magnetic System. In the Side of the Planisphere are Dashes, shewing the Boundaries of the Climates; and at the South Pole, are several Conclusions relating to the Earth, the Heavens, and the Universe.





A DESCRIPTION of the EASTERN CELESTIAL PLANISPHERE.

94. ** SES** HROUGH the Centre of this Planisphere, a right Line is drawn for the Axis of the World, the Extremities of which represent the North and South Poles in the Heavens, the Earth's Globe being supposed to be in the Centre, and no

greater than a Point compared with the Extent of the Heavens; and about this Axis the apparent diurnal Motion is performed, whereby the Stars appear to move from East to West, and rise and set, under the Direction of the Equator and Parallels of Declination. As a Circle, this Line represents the Semi equinostial Colure. At right Angles to the Axis of the World, the Semi celestial Equator is drawn through the Centre, for shewing the Sun's right Ascension as in the terrestrial Planisphere. The Circumference of the Planisphere hath the Degrees of Latitude or Declination from the Equator, and represents the solsticial Colure. This Circle shews what Stars will be vertical, or directly over the Inhabitants of any Degree of Latitude, when they come to the Meridian.

95. At the Distance of 23½ Degrees from the Equator, the two Semi Tropics of Cancer and Capricorn are drawn, and at the same Distance from the Poles, are drawn the two Semi polar Circles. Through the Centre, and to the Extremities of the two Tropics, is drawn a right Line for the Semi Ecliptic, containing the Signs, Capricorn, Aquarius, Pisces, Aries, Taurus, Gemini, and also the Days of the Month when the Sun is in those Signs. From Pole to Pole, are drawn the Semi Meridians, and from one Side of the Planisphere to the other, the Semi Parallels of Declination for shewing the right Ascensisher, between the Tropics, is placed a Scale of Declination, for shewing the Declination of the Sun on any given Day of the Year.

96. On the Surface of this Planisphere, are laid down the fixed Stars of the eastern Hemisphere, or those on the East-side of the solftitial Colure, with the Names of the Constellations to which they belong, and the Greek Letters of Bayer placed at each Star. The Figures of the Constellations, are drawn according to the Antients.

A

A DECRIPTION of the WESTERN TERRESTRIAL PLANISPHERE.

97. ** L. L. the Semi Circles delineated in the eastern terreA & strial Planisphere are continued, and the other
Halfs of them delineated in the same Manner, and
for the same Uses in this. Besides, which, on the
Surface of this Planisphere, is delineated the great Continent of
North and South America, connected by the Istomus of Darien,
with the Countries, Kingdoms, Islands, Rivers, Trade-winds,
Variation Lines, &c. of this Part of the known World, as in
the eastern terrestrial Planisphere. So that whatever hath been
said in the Description of that, may also be applied to this. And
these two Hemispheres being placed to the Back of each other,
do compleatly represent the whole Globe of the Earth and Sea.

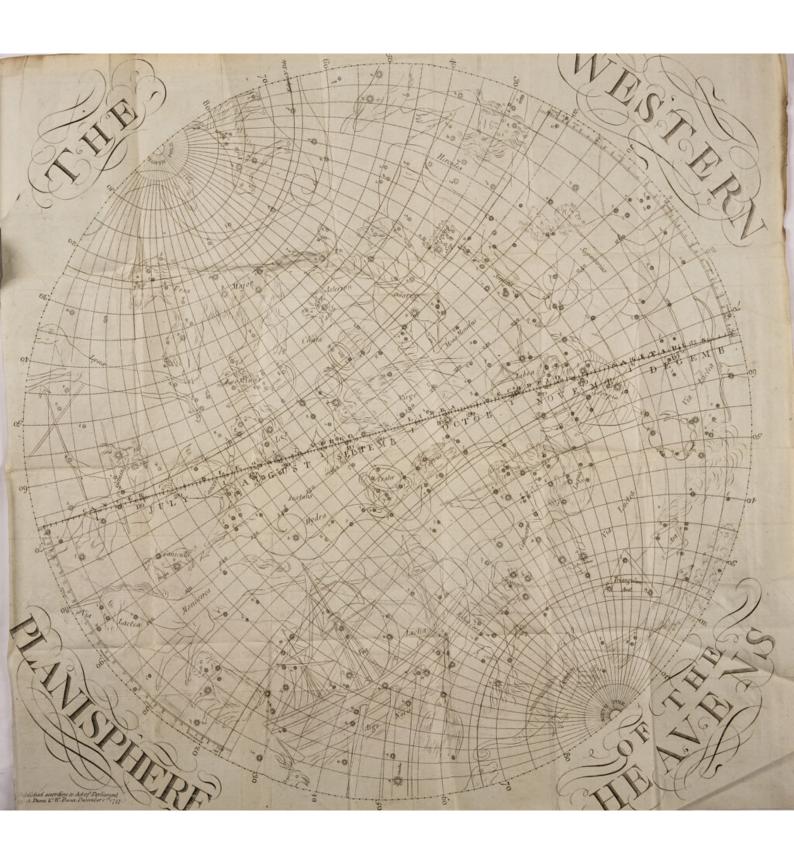
A DESCRIPTION of the WESTERN CELESTIAL PLANISPHERE.

98. OCO HE several Parts of the western celestial Planiof Sphere, are laid down in the same Manner, and for the same Purposes as those of the celestial eafern Planisphere; the solstitial Colure being the

Circumference of the Planisphere. A right Line passing through the Centre, at right Angles to the polar Axis, containing the other half of the celestial Equator. And a right Line passing through the Centre, to the Extremities of the northern and southern Tropics, containing the other six Signs, Cancer, Leo, Virgo, Libra, Scorpio, Sagittarius, and the fixed Stars of the western celestial Hemisphere being laid down in this: So that what hath been said of the eastern celestial Planisphere, may also be understood of the western One, the one being the Remainder or Complement of the other. And these two being placed to the Back of each other, do represent the whole Convex Surface of a Globe of the Heavens.









· The PRINCIPLES of GEOGRAPHY.

99. \$\overline{\text{D}} \overline{\text{D}} \overline{\text{EOGRAPHY}}\$ is a Science whereby we under\$\overline{\text{D}} \overline{\text{G}} \overline{\text{S}} \overline{\text{ftand}}\$ the Situation, Extent, Bounds, Air, Soil,
\$\overline{\text{D}} \overline{\text{D}} \overline{\text{S}} \overline{\text{C}} \overline{\text{S}} \overline{\text{C}} \overline{\text{Mommerce}}, Laws,
\$\overline{\text{Cultoms}} \overline{\text{C}} \overline{\text{C}} \overline{\text{C}} \overline{\text{C}} \overline{\text{C}} \overline{\text{C}} \overline{\text{C}} \overline{\text{Momerce}}, Laws,
\$\overline{\text{Cultoms}} \overline{\text{C}} \overline{\t

And here it is to be observed, that the usual Way of introducing the Situation of Places, is by their Latitudes and Longitudes; the former being easily obtained by celestial Observations, and the latter much better adapted for this Purpose than any other Position that could have been made use of in its Stead, on Account of the Uniformity and Regularity which subsists

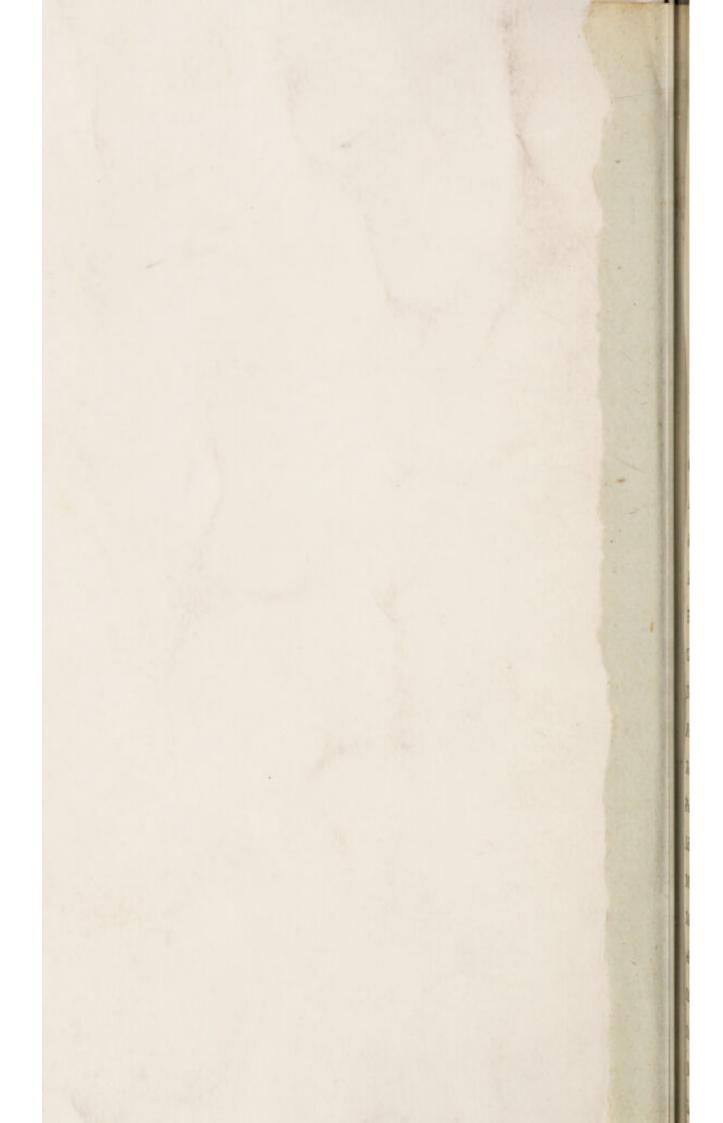
among the Circles of the Earth and Heavens.

Nor is it any Matter of Consequence, whether the Geographer begins to reckon Longitude from the Meridian of London, Paris, or any other Place; seeing, in all Cases, a Person with whom he converses, is supposed to know where he begins the same; and should any Ambiguity of Expression arise, the Word Difference of Longitude, instead of the Word Longitude itself, will obviate the same.

In the following Table the first Column contains the principal Divisions, Empires, and Countries of the Earth's Globe; the second Column expresses the Temperature of the Air for each respective Country, or its extreme Parts North and South; the third Column expresses the Soil, which if various and not uniform is marked different; the fourth Column expresses the natural Productions, Trade and Commerce, the fifth, sixth and seventh Columns express the general Religion, Language, Manners, and Properties of the Natives.

Coun-

Countries	Air.	Soil.	Productions, Trade & Commerce.
Spain	hot	dry	S Woods and Mountains, Fruits, Wines, Earths, Metals.
France	temperate	fruitful	S Corn, Flax, fine Wool, Wines, Silks, Linen, Iron.
Italy	temperate	fruitful	Corn, Wines, Fruit, Sattins, Minerals, Metals.
Hurgary	unwholfome	fertile	SCorn, Fruits, Copper, Iron, Salt,
Poland	cold	fertile	S Oak, Deals, Salt-petre, Cordage, Me-
Prussia	cold	woody	S Corn, Cattle, Oak, Deals, Hemp, Linen, Metals.
Germany	different	rugged	Metals, Allum, Corn, Linens, Ar-
Holland	temperate	watry	Cattle, Herbs, Fish, Linen, Earthen- Ware.
Denmark	cold	fertile	S Fish, Furs, Skins, Deals, Armour, Pitch, Iron.
Sweden	cold	rocky	Metals, Hydes, Furs, Rofin, Deals,
Norway	cold	rocky	Fish, Train-oil, Furs, Masts, Pitch, Copper, Hides.
England	temperate	fertile	Corn, Wool, Oak, Butter, Cheefe, Hard-ware, Armory.
Muscowy	hot and cold	woody	S Furs, Hemp, Soap, Leather, Train-
Little ?	temperate	different	Slaves, Leather, Skins, Furs.
Greece	temperate	fertile	Silks, Oil, Leather, Galls, Cotton, Wines, Mohair, Grograms.
Great ?	different	different	Flax, Musk, Rhobarb, Silk.
China	temperate	fertile	S Gold, Silver, precious Stones, China- ware, Spices, Perfumes.
India	hot	fertile	5 Drugs, Gums, Spices, Silks, Metals, Minerals.
Perfia	different	different	Silks, Gold, Silver, Skins, Fruits, Alabaster, Metals.
Arabia	hot	different	S Coral, Pearl, precious Stones, Gums, Perfumes.
Euphratia	temperate	fertile	Fruits and Grain, Pitch, Silk.
Egypt	unwholfome	flimy	Fruits, Sugar, Linen, Drugs, Flax,
Barbary	hot	fertile	Honey, Wax, Oil, Sugar, Flax, Hemp, Dates.
The Defert	hot	barren	Wild Cattle, fome Corn and abun-
Negroland	hot	fertile	Indian Corn, Herbs, Metals Gums, Elephant's Teeth, Gold, Slaves.
Comp	1		Guinea



Religion.	Language.	Manners and Propert. of Inhabit.
Romanists	Spanish	zealous, jealous, lascivious, impotent.
Romanists	French	gay, sprightly, generous, ingenious.
Romanists	Italian	polite, eloquent, cunning, revengeful.
Romanists	Hungarian	strong, valiant, cruel, adicted to war.
Romanists	Sclavonic	Scivil, brave, fincere, strong and long livers.
Lutherans	Sclavonic	faithful, strong, patient, valient.
Romanists	German	Sfincere, hospitable, valiant, patient, ingenious.
Calvinifts	Low Dutch	industrious, ingenious, civil, brave.
Lutherans	Teutonic	mistrustful, strong, valiant.
Lutherans	Teutonic	opinionated, vigorous, referved, flrong,
Lutherans	mix'd	hardy, laborious, courageous, plain, just and civil.
Protestants	mix'd	S fincere, industrious, cautious, ingeni- ous, valiant.
Grecians	Sclavonic	Sconceited, barbarous, long livers, ro- buft, fhort.
Mahometans	Stytian	Food Horse-stell and Mares-milk.
Mahometans	Greek	S zealous, charitable, vicious, robuft,
Mahometans	Tartarian	E barbarous, ugly, fierce, thievish, eaters of human Fleih.
Pagans	Chinese	conceited, crafty, faithless, ingenious.
Idolaters	Indian	Scivil, ingenious, and pretty just int Dealing.
Mahometans	Persian	Sproud, passionate, revengeful, Liber- tines, Flatterers.
Mahometans	Arabic	different, fome ignorant and barbar-
Armenians	mix'd	different, but generally like those of Arabia and Greece.
Mahometans	Arabic	Streacherous, cruel, Cowards, luxuri- ous, Diviners.
Mahometans	African	Sunfaithful, impatient, covetous, crafty, tawny.
Mahometans	African	wild, barbarous, thievish, tawny.
dolaters	African	S barbarous, rude, ignorant, luxurious, vicious, black.
	The state of the s	Pagans

Countries.	Air.	Soil.	Productions, Trade & Commerce.
Guinea	hot	fertile	Gold, Ivory, Ambergreafe, Grain, Fruit, Wax, Civet, Slaves.
Nubia	hot	fertile	Sold, Ivory, Sugar, Drugs, Croco-diles.
Southern }	hot	different	S Gold, Silver, Metals, Gems, Corn, Cattle. Wines, Salt.
Mexico	warm	fertile	S Cotton, Sugar, Silk, Cochineal, To- bacco, Drugs.
Florida	temperate	fertile	SGold, Silver, Pearls, Furs, Indigo, Wild-beafts.
Canada	temperate	fertile	Fish, Grain, Deals, Iron, Tar, Beasts, Beavers, Deers, Tobacco.
Terra Firma	hot	fertile	Gold, Silver, Metals, Gums, precious Stones, Drugs.
Peru	different	rich	Waft Quantities of Gold and Silver, with Drugs, Pearls, Cotton.
Amazonia	temperate	fertile	Gold, Silver, Sugar, Cocoa, Ebony, Tobacco, Wild-beafts.
Brazil	temperate	fertile	S Brazil Wood, Sugar, Amber, Tobacco, Train-oil, Rofin, Gold, Diamonds.
Paraguay	temperate	fertile	S Gold, Silver, Brass, Iron, Sugar, Wild-beasts.
Chili	different	different	Gold, Silver, Corn, Honey, Offriches, Metals.

There are also many Islands dispersed throughout the several Oceans and Seas, and also Places on the Continents of great Trade and Commerce, some of which are subject to European Nations, and others neutral with them. Such as the Azore, Madeira, and Cape-Verd Islands, possessed by the Portuguese; Newfoundland, belonging to the English; the Isles of Cape-Breton and St. John, now taken from the French by the English, and which, on Account of Situation, may be esteemed a Bar to the River St. Laurence, and all the Settlements thereon; the Sand-Banks in these Seas being general Fisheries for Codfish, Stock-fish and Lyng, which those Seas produce in great Abundance, and from which they are exported to Portugal, Spain, France, England, and other Nations. Greenland, both East and West, the general Whale-fishery for the English. Dutch, Danes, and other European Nations, the Coldness of the Climate permitting but few Inhabitants, a few wandering San vages, with Plenty of large, white Bears, icy Seas, and Mount tains of Snow. Iceland, inhabited by a few Danes and Nor wegians, the Air cold, the Inhabitants healthy, long livers, and thu

ment,

Religion.	Language.	Manners and Propert. of Inhahit.
Pagans	Zungai	naked, proud, thievish, black.
Pagans	Nubian	cunning, warlike, fierce, wealthy.
Pagans	Ethiopic	Sin general ignorant, treacherous, fierce Savages.
Papists & Idolaters	Spanish and Indian	civil, honest, faithful Natives.
Ditto	French and Indian	Ditto with painted Bodies almost na- ked, Warriors.
Ditto and English	Ditto and English	Ditto and English.
Idolators	various	tawny, robust, healthy, almost naked
Idolators	Spanish and Indian	Svicious, fome ignorant, other ingeni- ous.
Pagans	not known	fierce Savages, eating human Flesh.
Scarce any	various	cruel, revengeful, thievish.
Idolaters	Jargon	flout, tall, nimble, laborious.
Idolaters	Fargon	tall courageous, warriors, valiant.

the Produce of the Country, Leather, Tallow, Fish-oil, Fish and Skins. Jamaica, well inhabited by the English, producing Sugar, Rum, Ginger, and otherwife a great Mart for the Trade and Commerce of the West Indies. St. Christophers, Antegoa, Monferrat, possessed by the English, producing great Quantities of Cotton, Sugar, Tobacco. Barbadoes, extremely populous, and producing great Quantities of Sugar, Rum, Cotton. Most of the other Caribbee Islands inhabited by the Spaniards and French. Fuego, an uninhabited Ille fouth of America. Madagascar, a wild defert Island, having but few, if any, European Inhabitants. St. Helena, belonging to the Englift, where our East India Ships take fresh Supplies of Water. Sumatra, producing Gold and other Metals, Drugs and Spices. Java, inhabited by the Dutch, producing Rice, Honey, Sugar. Japan, producing Gold, Silver, Wheat, Rice, Pearls. Alfo, on the Continent of America; the Coasts of Hudson's Bay and New Britain, belonging to the English, producing great Quantities of Beavers and Furs, but few Natives. New Scotland, belonging to the English, a Country capable of great Improvement, and producing Timber, Hemp, Cordage, Wheat, Fruits. New England, numerously inhabited by the English, a fertile Country, the Inhabitants healthy, robust, sincere, courageous; also New York, Pensilvania, Maryland, Virginia, North and South Carolina, Georgia, abounding with fine Rivers, plenty of Fish, great Quantities of Tobacco, and many other mercantile Commodities, the Natives chiefly British Inhabitants, whose Geniuses for Trade and Commerce have led them to these Parts of the World, where they have established Colonies, built spacious Towns, settled Factories, and under the British Government export the Produce of those Countries, import that of others, a People apt for Improvements in Arts, Sciences and Learning, generous, strong and brave, equal to any Enterprize, and biding fairly to equal, if not outvy any other Part of the known World.

On the Coast of Africa and Corimandel, Bengal, Malabar, and several other Places in East India, the English have several Places of great Consequence for securing the Trade of these Parts of the World, as have likewise the Spaniards, Portuguese, Dutch and French, the Prerogative of which is generally retained for the immediate Benefit of the several Crowns and Princes which have been able to make Acquisitions in those Parts of the World, or for Companies of Merchants impowered by Royal Charters for carrying on Courses of Traffic with the same.

100. An Estimate of the Magnitudes of Countries compared with each other.

England and Scotland - 1	Sweden and Norway - 3:
	Muscovy - 12
France — 31	Poland — 3
	Turkey in Europe 5
Italy — — 1\frac{1}{2}	Turkey in Afia - 6
Germany — 3	East India — 20
Holland and Flanders -	China and Tonquin - 18
	Arabia — — — 15

101. An Estimate of the Number of Inhabitants in the following Countries.

England - 7 Millions.	Turkey in Europe 8 Millions.
Scotland - 1 1 Mill.	China — 50 Mill.
Ireland - 2 Mill.	Persia — 12 Mill.
Holland — 2 Mill.	India - 20 Mill.
Flanders — 2 Mill.	British 7
France - 16 Mill,	Empire in { 1 Mill. Whites.
Spain and Portugal 6 Mill.	America 3
Italy - 7 Mill.	French in)
Germany - 20 Mill.	Canada and \ 3 Mill. Whites.
Prussia and Poland 5 Mill.	Florida
Denmark, Sweden & Mill.	Spaniards & 3 Mill. Whites.
and Norway 55 William	in America \$ 3 will. willies.
Muscovy — 12 Mill.	Paragona to 1 con 12 a s
and the same of th	

in Time of Peace in the following Countries.

England — 40 Holland — 30	Thousands.	Queen of Bungary 3 120 Thousands,
France - 160		Prussia - 100 Thous.
Spain — 60	Thouf.	Poland — 40 Thouf.
	Thouf.	Saxony — 20 Thous.
The Pope's 2 g	Thouf.	Bavaria — 20 Thous.
Dominions 3 8	I moun.	Hanover — 20 Thouf.
Naples - 20	Thous.	Denmark — 30 Thous.
Tuscany — 8	Thouf.	Sweden — 35 Thouf.
Venice — 20	Thouf.	Muscovy — 150 Thous.
Sardinia — 30	Thous.	of Gilders of The Skindalles at

103. An Estimate of the Proportion of Shipping, including those of War and Commerce, belonging to the following Countries.

England - 1	Germany and Pruffia ;
and the second s	Denmark 3
France — — -	Sweden and Russia
Spain and Portugal -	

Par of Exchange, and the real Species of the Chief trading Countries of Europe.

HE Countries and Cities with which England has the greateft Commerce, are Holland or Amsterdam, Brabant or Antwerp, France or Paris, Germany, Hamburgh, Spain, Portugal, Venice, Genoa and Leghorn.

In England, Books and Accompts are kept in Pounds, Shillings, Pence and Farthings. A Pound is 20 Shillings, a

Shilling 12 Pence, and a Penny 4 Farthings.

The Exchange between this and other Countries is negotiated

in Pounds and Pence.

The real Species of Gold is a Guinea (or 21 Shillings) and its Sub-divisions: That of Silver is a Crown, or 5 Shillings, its Half is 2s. 6d. The next of Silver is a Shilling, and the next is a Sixpence: Of Copper a Halfpenny, and a Farthing.

In Holland, Books and Accompts are kept in Gilders,

In Holland, Books and Accompts are kept in Gilders, or Florins, Stivers, and Penningens. A Gilder is 20 Stivers:

A Stiver is 16 Penningens.

The Exchange between this and other Countries, is by allowing a certain Number of Pounds, or Groots Flemish, for the respective Species of these Countries. A Flemish Pound contains 20 Flemish Skillings: A Skilling contains 12 Groots, and a Groot Flemish is half a Stiver, consequently a Gilder is 40 Groots, and a Pound Flemish is 6 Gilders. — Sometimes the Exchange is by Rixdollars, one of which is 2 ½ Gilders.

The real Species of Gold, are the Ducat or Ducatoon of 20 Gilders. The Sovereign is 15 Gilders or Florins. The Rose Noble is 11 Gilders. The real Species of Silver and Copper are, the Ducatoon of 3 Gilders 3 Stivers. The Drie Gulden is 3 Gilders. The Rixdollar is 2 ½ Gilders. The Crown is 2 Gilders. The Dollar is 1½ Gilder. The Gold Florin is 1 Gilder 8 Stivers. The Skilling is 6 Stivers, or 12 Groots. The Stiver is 2 Groots, or 16 Penningens or Deniers. The Male Skilling is 5 Stivers 8 Deniers. The Doublekee is 2 Stivers.

The intrinsic Value of a Florin or Gilder is 1s. 9d. -6 Eng-

lish, and consequently a Groot is +5% of an English Penny.

In Brabant, Antwerp, &c. Books and Accompts are kept in Pounds, Skillings and Groots Flemish, and Exchange negotiated in the same Species. The Rest as in Holland, only 96 Groots Flemish are here 4s. 6d. English, and consequently I Groot Flemish is -2 of an English Penny. In

In France, Books and Accompts are kept in Livres, Solz, and Deniers. A Livre is 20 Solz, and a Sol is 20 Deniers.

The Exchange is negotiated by Ecus, or Crowns of 3 Livres,

or 60 Solz Tournois.

The real Species of Gold, are double and fingle Louis d'Ors. Of Silver a Crown, or Ecu, and its Sub-divisions: Une Petite Piece is $3^{\frac{1}{2}}$ Solz, or three Pence English. — Of Brass a Sou. Of Copper a Liard, or Farthing, and a Denier $\frac{1}{12}$ of a Sol.

The intrinsic Value of a Louis d'Or is 17s. od. of a Crown

in Specie 4s. 6d. of a Sol 73 of a Penny English.

The current Value of the several Species in France alters at the King's Pleasure: The Crown in Specie lately passed there for 72 Solz Tournois, and therefore then the intrinsic Value of the Ecu of Exchange of 3 Livres or 60 Solz Tournois was 45 Pence English. The Value of the Ecu of Exchange is according to the current Value of the Ecu in Specie.

In Hamburgh, Books and Accounts are kept in Rixdollars or Marks, Skillings and Pfennings. A Rixdollar is 3 Marks, a Mark is 16 Skillings, or Stivers Lubs, and a Skilling is 12

Pfennings.

The Exchange with London is by Pounds, Skillings, and Groots Flemish. A Groot Flemish is ½ a Skilling, or Stiver Lubs: A Skilling Flemish is 6 Stivers Lubs, and a Pound Fle-

mish is 7 1 Marks, or 120 Stivers Lubs.

The real Species are a Gold Ducat, worth 6 Marks, and 8 or 9 Stivers Lubs. A Silver Ducatoon is 3 Marks 14 Stivers Lubs. An Albertus, or Cross Rixdollar is 3 Marks 4 Stivers Lubs. The current Rixdollar is 3 Marks. A Danish Crown is 2 Marks. The Dollar is 2 Marks. The Mark is 16 Stivers Lubs. The Skilling is 6 Stivers Lubs, and the current Stiver is 2 Groots current.

The intrinsic Value of the Mark is 1s. 6d. English, by which

the Value of all the Rest may be estimated.

In Spain, Books and Accompts are kept in Rials and Ma-

ravedies. A Rial is 34 Maravedies.

The Money of Exchange is a Piastre, or Piece of Eight, of 8 Rials, or 272 Maravedies. A Ducat of 375 Maravedies, and a Pistole of 1088 Maravedies.

The real Species are, the Pistole of Gold of 4 Pesoes. The half Pistole. The Peso, Piastre, or Piece of Eight, is 8 Rials.

The Rial is 34 Maravedies.

The intrinsic Value of the Peso is 4s. 6d. English, consequently a Maravedie is 27 of an English Farthing. In

In Portugal, Books and Accompts are here kept in Rees, the Thousands being separated from the Hundreds by this Mark—, and the Millions from the Thousands by a Dot .

Exchange is negotiated by allowing a certain Number of Rees for the Species of other Countries, viz. the half Pistole, or Mill-

ree of 1000 Rees, and the Crusado of 400 Rees.

The real Species are, the broad Ducat of Gold of 10,000 Rees. The double Pistole is 40000 Rees. The Pistole is 2000 Rees. The half Pistole is 1000. The stamped Patacoon is 600. The current Patacoon is 500. The stamped Crusado is 500. The current Crusado is 400. The stamped Pezzo of Portugal is 480, and the Testoon is 100 Rees.

The Value of a Millree is 84d. English, consequently a Ree

is 336 of an English Farthing.

At Venice, Books and Accompts are kept in Lires, Soldi, and Denari. A Lire is 20 Soldi: A Soldo is 12 Denari.

The Exchange is by Ducats and Groffes Banco; a Ducat

Banco being 24 Groffes Banco.

The real Species are, a Piftole of 29 Lires; a Sequin 17 Lires; a Silver Ducat 6 Lires 4 Soldi; a Teston 2 Lires 14 Soldi; a Jule is 5 of a Teston, or 18 Soldi; a Lire is 20 Soldi; a Soldo is 12 Pichioli; a Gross is 2 Soldi 8 Pichioli, or 32 Pichioli.

The intrinsic Value of a Lire is 7d. \(\frac{1}{2} \) English, and of a Silver Ducat 44 d. 64 English: But the Ducat Banco being 20 per Cent. better than the currant Ducat, and 40 per Cent. better than the Pichioli, is 53d. 57 English, consequently the Gross

Banco is 2d. 23 English.

At Genoa, Books and Accompts are kept in Pezzi, Lires, Soldi, and Denari. A Pezzo is 5 Lires. A Lire is 20 Soldi; and a Soldo is 12 Denari.

The Money of Exchange is the Crown, of 4 Lires; and the

Pezzo, of 5 Lires.

The real Species are, the Pistole, of 18 Lires; the Pezzo, of 5 Lires; the Lire; the Soldo; the Croisade, of 7 Lires 10 Soldi; and the Teston, of 1 Lire 12 Soldi.

The intrinsic Value of a Lire, is 10d. 4 English.

At Leghorn, Books and Accompts are kept in Pezzi, or Dollars, Soldi, and Denari: A Pezzo, or Dollar, is 6 Lires: A Lire is 20 Soldi: A Soldo is 12 Denari.

The Money of Exchange is the Pezzo, or Dollar.

The real Species is the Pezzo, of 6 Lires.

The intrinsic Value whereof is 4s. 6d. English. 105. A

105. A TABLE, shewing the Par of Exchange with the Chief trading Places in Europe.

			The second second	
HOLLAND OF AMSTERD.	20 Shillings for 37s. od. & Flem.		tool man 1941	nos neómuro a nos comos sin e
BRABANT Or ANTWERP	20 Shillings for 35s. 6d. 3 Flem.	201. od. Flem. for 191. 2d. 3 Flem.	Antwerp.	Squar Banco la 10 6 Groote (c. 12) Pagges
FRANCE OF PARIS.	45 Pence for 1 Ecu of 60 Sols Tournois	83 TGroots for 1 Ecu.	80 Groots for 1 Ecu.	PARIS.
GERMAN. Or HAMBUR.	20 Shillings for 355.6d.2 Flem.	20 s. o d. Flem. for 19s. 2d. ² / ₅ Flem	96 Groots for 48 Stivers Lubs.	Ecu for 40 Stivers Lubs.
SPAIN.		137 59 Groots for a Ducat of 375 Maravad.	132 To Groots for a Ducat.	1 Ecu for 226 $\frac{2}{3}$ Maravad.
PORTU-	6s. 9d. Sterling for a Millree.	60 Groots for a Crusadoe of 400 Rees.	57 3 Groots for a Crufado of 400 Rees.	1 Ecu for 555 \$ Rees.
VENICE.		99 Toots for a Ducat Banco.	95 -95 Groots for a Ducat Banco.	Ducats Banco.
GENOA	54 Pence for a Pezzo.	100 Groots for a Pezzo.	96 Groots for a Pezzo.	Ecu for 83 ½ Soldi.
LEG- HORN,	54 Pence for a Pezzo or Dollar.	100 Groots for a Pezzo.	96 Groots for a Pezzo.	100 Ecus for 83

96 THE UNIVERSAL

HAMBURGH.

for a Ducat.	SPAIN.			
57 3 Groots for a Crusadoe.	100 Ducats for 299 53 Crufad.	PORTUGAL.		
	269 103 Mara- vedies for a Ducat Banco.	661 3 Rees for a Ducat Banco.	VENICE.	
96 Groots for a Pezzo.	272Maravedies for a Pezzo.	666 Rees for a Pezzo.	119 1 foldi for 4 Lires.	GENOA.
96 Groots for a Pezzo.	272Maravedies for a Pezzo.	666 % Rees for a Pezzo.	Banco for 99 1/3 Pezzi.	Pezzo for Pezzo.

Note, Exchange is very variable, and the Course or Rate at any particular Time, is only to be known from Merchants residing at the Places.

106. A TABLE, shewing the foreign Gold and Silver Coins.

Gold Coins.	Weight.	Value.
Gold Collis.	The second secon	
- Paris District	pross. gr.	1. s. d.
C Panish Pistole	4 8	0 16 6
French Piftole	4 8	0 16 6
French Lewis -	- 48	0 16 6
Hungary Ducat	_ 2 0	0 9 8
Holland Rider — — —	6 10	1 5 9
Flanders Albertus — — — —	- 3 8 6 3 ±	0 13 0
Double Sovereign of Flanders	6 3 1	
Gilder of Prothono — — — —	- 2 3	0 6 10
Gilder of Matthias Emperor	- 2 3	0 7 2
Gilder of Nuremburg	2 3	0 7 1
Italian Pistole — — —	4 6	0 16 4
Chickeen of Venice	2 7	0 9 7
Half Cardinal Angel — — —	- 1 20	0 9 7
A Danish Coin with a Crown on one Side of it -	- 3 20	0 15 4
Moidore with their Halves, Quarters and Eighths	- 6 II	1 7 0
Johns, with their Halves, &c	9 6	1 16 0
Double Johns, of	18 12	3 12 0
Silver Coins.	oz.pwts.gr.	s. d. f.
Ducatoon of Flanders -	I 0 22	5 4 0
Holland Dollar — — —		4 4 0
Lyon Dollar —	0 18 5	T
Rixdollar of the Empire		3 4 2
Seville Royal —		4 5 3
otting stoyal	0 17 12	4 4 3
		Mexico

	02	pw.	s.gr.	5.	d.	f.
Mexico Royal -	0	17	12	4	4	2
Old Cardecue — —	0	6	31	I	6	I
French Lewis	0	17	11	4	4	1
St. Mark of Venice —	0	10	4	2	6	a
Single Milrez of Portugal -	0	7	2	1	9	0
Double Dutch Stiver	0	1	0	0	1	3
Zealand Dollar — —	Ò	10	0	2	9	2
Crofs Dollar	0	18	0	4	2	I
Old Philip Dollar ——	.1	2	0	5	0	0
Ferdinand Dollar, 1623	0	18	6	4	3	1
Prince of Orange Dollar, 1624 -	0	18	6	4	3	3
Danish Dollar, 1620 -	0	13	0	2	11	1
New Quarter of French Lewis -	0	4	9	I	1	0
Portugal Testoon	0	4	9	1	1	0
THE PART OF THE PARTY OF THE PA		0,000	17006	1	1000	100

107. A TABLE of the Climates from the Equator to the Polar Circles.

A STATE OF THE PERSON NAMED IN COLUMN TWO IS NOT THE PERSON NAMED IN COLUMN TWO IS													
Clim.	Longest Day.	Latitu	ide	Bre	adth.	Clim.	Longest Day.	Latit	ude	Bre	adth.	-	
E-B	h.	0	1	0	-,		h.	0	1	0			
1	121	- 8	34	8	25	13	18:	59	59	1	33	1	
2	- 13	16	43	8	28	14	19	61	18	1	19	1	
3 4	131	24	11	7		15	192	62	25	1	7		
1 4	14	-30	47	6	36	16	20	63	23	0	58		
-	141		70	-	40		20	64	16	-	-		
5		36	30	5	43	17	20 1	64	H-332	0	53		
7	15	45	29	4	7	19	211	65	55 25	0	39		
7 8	16	49	1	3	32	20	22	65	47	0	22	-	
-	MANUAL PROPERTY.		-	-	-	-			17			1	
9	161	51	58	2	57	21	221	66	6	0	19	1	
10	17	54	29	2	31	22	23	66	20	0	14		
11	17:	56	37	2	12	23	231	66	28	0	8	1	
12	18	58	26	- 1	49	24	24	66	31	0	3	-	
7.0	Climate	s beta	veer	the th	e Pol	ar Circ	eles and	the I	Poles		10 83		
1	Months	0	1	10	1	1	Months	0	1	0	,	1	
I	1	67	31	1	0	4	4	78	20	5	a	-	
2	2	69	31	2	- 0	5	5	84	0	5	40	-	
3	3	73	21	3	30	6	6	90	0	6	0	1	

108. A TABLE of the Equality of Weights, at the most considerable Places in Europe.

Hamburgh.	102	901	87 12		82 10	-		100		40 2		-	61 3			* 3	- 1
London.	4. 58	113 14											66 11	-			104
Geneva.	Same.	92 9		-					-	-			53 6			-	-
Province or Marfeilles.	123 8	128 8		24 1	-	128 4	13	121	26	17 1			72 2	-	-		117 5
Tholouse.	00	122 11	10	6	000	32	7 1	1 61	20	12	8		0 0	1 6	-	33	2
Rochel.		93 13			80 15												94
Lions,	911	120 8	100		98 5	30+		13					60				
Rouen.	9	4 1001	2 1	7	- ox	201	00	4	00		4	0 1	57 4		10	84 4	1
Brabant or Antwerp.	105 8	100	90 12		9 63					001	70 5	72 12	62 12	62 0	00 12	92	100 3
Holland or Amsterdam.	100	94 12	98	101	84 12		8 10		102 6	95 %	663		26 8	9 5		87 8	95
DE STORY	1_				222		ot	Ist	nba	91	B						
	CAmflerdam,	Antwerp,	Lions	Rochel	Tholoufe	Geneva.	London,	Hamburgh,	Francfort,	Leipfick,	Genoa,	Leghorn,	Milan,	Nanles,	Spain.	Portugal,	Leige.
	1						3E	q	0	01			*				

- 80	A STREET, STREET	100				1									12						1		~
	Leige.	105 4	-		8 06										-			62	99 14	92	001	l compre-	100. 4
	Lifbon.	114 8			8 86	-		-	28	104 13	12	-	100	76 5			-	6 29		100	108 12	they are all	**
	Sevil, Cadiz.	901			91 3		-	85 13		97	103 13		101	70 11	73	63	63 91		100	92 12		Amsterdam; th	1
.,	Naples.	691		170 12	145 6	170 11	144 4	-	189 14	-	167 10	72	161	-	6 911	00		1001	159 12	148	101 091	of.	cfort.
ghts, &c.	Venice.	991			142 3	67 1		134 8		152	-	69	28	110 11		-	1001	98	-	145 4	112 251	nal to those	mhera are comprized under those of Francfort.
of Wei	Milan.	891		-	144 8	-		36	88 1	-	64 1	11			911		-	99 2	8 1	147	16 651	almost equal	1 under th
the Equality of Weights,	Leghorn.	146	1	- 0	124 11	9	2 1	1		2 1	4	147 14	00	111 96	100	86 4	7	10	- 1	131 4	2	being	comprize
fo	Genoa.	1001	142 2		129					137 4		153	142 13		#100	89 4			-	131 4		ces of France,	
ATABLE	Leipfick, Naumburgh, Hall.	100			00			87 8	611	1 96	102 15	107	100	70	72 8		63	62	4 66	91 13	-	and other Place	Weights of Nure
,	Franckfort.	1 80	02 13	102	84 4	66	000		0	89 71	. 96	001		64 5	100	28 2	-	57 13	92 10	85 12	93		the Weigh
	4. 43	ī			1	THE REAL PROPERTY.		1		01	len	ba	Э.	g								ts of	nder the
		C Amflerdam.	Antwerp.	Rouen.	Lions,	Rochel.	Tholoufe,	Marfeilles,	Geneva,	_	_	~		-	Leghorn,	Milan.	Venice.	Naples.	Spain.	Portugal,	[Leige,]	Note, -The Weights of Paris,	hended under those of
	The same of the same	1		1		14	-	-		16	4	0	001	1		-						1	

For the like Reafon, the Weights of Nuremberg are comprized under more of francjorn

20	Canes of Mar 100 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
urope.	Ells of Geneva. 0 0 0 0 0 8 4 5 5 5 6 5 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
Places in Europe.	Ells of St. Gali 2 10 - 14 - 14 - 15 - 16 - 16 - 16 - 16 - 16 - 16 - 16
Place.	Ells of St. Gall 9 2 4 2 9 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2
derable	Ells of Sweden 78 5 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2
a confe	Ells of Berguen. 0 = 88 6 0 6 2 6 8 8 5 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6
Long Meafures, at the most confiderable	Ells of Dant- 14 4 2 6 6 6 6 7 9 6 7 9 6 7 9 6 7 9 6 7 9 6 7 9 8 8 7 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9
s, at	Ells of Breflaw. 22 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2
Teafure	Ells of Hamb 0 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2
Ong N	Ells of England, 70 0 0 4 4 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5
	Ells of Brabant, 200 9 8 8 7 8 8 9 2 8 2 8 2 8 6 8 8 9 9 8 8 7 8 8 9 9 8 8 9 9 8 8 9 9 8 8 9 9 8 8 9 9 8 8 9 9 8 8 9 9 8 8 9 9 8 8 9 9 8 8 9 9 8 8 9 9 8 9 9 8 9
quality	Ells of Holland. 8 28 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8 8
the ?	тэке
ACO. A TABLE of the Equality of	Ells of Holland, Ells of Brabant, Ells of Brabant, Ells of Brabant, Ells of Brabant, Ells of Breflaw, Ells of Berguen, Ells of Berguen, Ells of Berguen, Ells of Sweden, Ells of Spain, Sards of Portugal, Sarces of Portugal, Braces of Florence, Braces of Florence, Braces of Milan.
ATA	of Holland, of Brabant, of Brabant, of Breflaw, of Breflaw, of Bergland, of Bergland, of Bergland, of Berglaw, of Berglaw, of Gerglaw, of Sweden, of Geneva, of Genev
.60°	Ells of Holland, Ells of Brabant, Ells of Brabant, Ells of Brabant, Ells of Breflaw, Ells of Breflaw, Ells of Berguen, Ells of Berguen, Ells of Sweden, Ells of Sweden, Ells of Geneva, Canes of Touloufe, Ditto, of Genoa, 9 P Canes of Touloufe, Ditto, of Genoa, 9 P Canes of Touloufe, Ditto, of Genoa, 9 P Canes of Fortugal, Yards or Barra's of S Barra's of Portugal, Braces of Portugal, Braces of Florence, Braces of Florence,
	100 00 00 00 00 00 00 00 00 00 00 00 00

	Braces of Mi-	11 10	00	N +	-	49	4 4	3672	50	39	50	0 0	26	N	100	Braces
	Braces of Florence, Leghorn.	1162	1995	935	1001	135 5	1933	3335	381	353	1453	161	114	1003	001	
	Braces of Bergam, Bologn.	1051	179	9333	944	122	1724	301	3445	319	1312	1723	1054	1001	95	and Mon
	Braces of Ve- nice.	102 1034	1743	- 0	- 15	00	200	2913	- w	90	27	9	0 0	97	875	illes
, Edc.	Cavido's of Por- tugal.	1001	171 833	6.8	98	911	1662	286	327	303	125	164	086	95	1000 L	Canes of M
Meafures,	Barra's of Por- tugal,	613	5043	00 4	n n	0	4 -	4 0	0		9	0	- 0	100	1 13	fore,
Long Me	Yards and Bar- ra's of Spain.	% 8 8	1364	64 715	72 684	924	33	2284	63	242\$	001	1311	784	,94	684	and Franc
of	Yards of Eng- land.	75.	128± 62±	60年	674	87	1243	2142	2454	2275	933	123	73	724	653	0 0
the Equality	Canes of Rome.	33.4	564	264	2 20 2	282	291	944	108	001	414	545	33	 	100 H	西田
the E	Canes of Ge- noa.	3049	25.5	777			0	P	. ()	60 6	1 00	O	\circ		(V) (I)	of the
SLE of	Canes of Tou- louse.	382	314	0 %	10 co co		332	1001	10 (1133	464		37/1			
ATABLE	77						ə	Mean								mee ar
A	andigon and property of the control	of		100 Ells of Breflaw, 100 Ells of Dantzick,	100 Ells of Berguen,	-	100 Ditto, for woollen,	100 Canes of Marfeilles,		0,	100 Yards or Barra's of Spain,	100 Barra's of Portugal,	100 Cavido's of Portugal,	Acres 1		

110. A TABLE of English Coins, Weights and Measures.

Money.	Long 1	Measure.								
Farthings.	Barley Corns.									
4 1 Penny.	3 1 Inch.									
48 12 1 Shilling.	36 12 1 Foo	it.								
960 240 20 1 Pound.	108 36 3	I Yard.								
Europe Description of the Control of	108 36 3 594 198 16½	5½ I Pool.								
Troy Weight.	23760 7920 660 22	o 40 1 Furlong.								
	190080 63360 5280 176	50 320 8 1 Mile.								
Grains.		. A day house								
24 1 Pennywt.	Land	Meafure.								
480 20 1 Ounce.	¥1. ¥									
5700 240 12 1 Pound.	Links.									
Superficial Measure.	625 1 Pole.									
Superficial Wienjure.	25000 40 100000 160									
Inches.	64000000 102400 256									
144 I Foot.	04000000 102400 230	0 040 1 21110.								
1296 9 1 Yard.	Tin	ve.								
4 6 8 6 8 8 8 8 8 8 8										
Solid Measure.	Seconds.									
S Common war and work	65 1 Min									
Inches.	3600 60									
1728 1 Foot.	86400 1440									
46656 27 Yard.	31556937 525949 876	00 3054 1 Year.								
Avoirdupoize We	ght. Apo	thecaries Weight.								
500000										
Drams.	Grains.	C1- 2								
16 1 Ounce.	20 1	Scruple 3								
256 16 1 Pc 28672 1792 112 1	und. 60 3 Hundred. 480 24	1 Dram 3 8 1 Ounce 3								
573440 35840 2240 20	1 Ton. 5760 288	96 12 1 Pound.								
3/3440 33040 2240 20	3/00 200	90 12 1 1041141								
Wine Meafure	7-1	Beer Measure.								
Solid Inches.	Sol. Inches.									
231 8 Pints 1 G	allon. 282 8 P	282 8 Pints 1 Gallon.								

2538

5076

18

288 36 4

72

144

1 Firkin.

2 1 Kilderkin.

504 63 1 Hogshead.

1018 126 2 1 Pope.

2016 252 4 2 1 Tun 10152

14553

29106

58212

III. A TABLE of the GENERAL POST.

BY a Statute made in the 9th Year of Queen Anne, the following Rates are appointed to be paid for Post Letters:

England, or South-British Inland Letters and Packets.

6 3 8 6 9 8 8 8 8	Si	ng.	D	ou.	T	re.	Oun.		
From Landon to on from any Place not ex-	s.	d.	5.	d.	5.	d.	s.	d.	
From London, to, or from any Place not exceeding 80 Miles Distance. And to and from any Place above 80 Miles from London.	0	3	0	6	0	9	1	0	
from London.	0	4	0	8	I	0	1	4	

North Britain Letters and Packets.

From London to Edinburgh, Dumfreys, or Cock- burn-speth, and the Contrary.	s.	d.	5.	d.	15.	d.	5.	d.
burn-speth, and the Contrary. From Edinburgh, not exceeding 50 Miles in	0	6	I	0	1	6	2	0
From Edinburgh, not exceeding 50 Miles in					1			
		2	0	4		1	0	8
From any Place above 50 Miles, and not ex-	10				30			
From any Place above 50 Miles, and not exceeding 80 Miles in Scotland. And if to, or from any Places above 80 Miles	0	3	0	6			I	0
And it to, or from any Places above 80 Miles				0				
from Edinburgh in Scotland, &c.	0	4	10	8	1		11	4

Ireland Letters and Packets.

From London to Dublin, and the Contrary.		6	1	0		2	0
From Dublin, to, or from any Place not ex- eeding 40 English Miles.	0	2	0	4		0	8
*** ***	0		-	8	2	1	4

Foreign Letters and Packets.

From any Part of Holland, France, or Flan-	s.	d.	5.	d.	5.	d.	s.	d.
ders, and from all Parts whatfoever, coming by								
Way of Holland to London; and also from Lon-				To de				
don to Hamburgh only.	0	10	1	8	2	6	3	4
To or from London, and any Part of Spain and								
Portugal through France, or by Packet Boats.	1	6	3	0	4	6	6	0
To or from London, and any Part of Italy, Si-						74		
cily, Turkey, Switzerland, Alface and Nancy, Bar-				.,				172
leduc, and Luneville in Lorrain, thro' France.	I	3	2	0	3	9	5	a
From London to any Part of Italy, Sicily, Ger-			1					
many, Switzerland, Denmark, Sweden, Russia,	1		1				1	
and all Parts of the North through Holland; and	1		1				1	
from each of these Places coming by Way of		0	1.	-	1.	0	1.	-
Flanders.	1 4	0	1.4	0	13	0	14	0

The Postage of Letters, to any of the following Places, must be paid for before they are sent, viz.

For all Letters or Paquets to any Part of Spain or Portugal.	5.	d	1 5.	d.	5.	d.	5.	đ.
or Portugal.	I	6	3	0	4	0	6	0
For all Letters or Paquets to any Part of Italy			-	+ 24	195			
or Sicily, by Way of Lyons, or to Turkey, by Way of Marseilles.								
Way of Marseilles.	I	3	2	6	3	9	5	0
For all Letters or Packets to any Part of Italy								
or Sicily, Germany, Switzerland, Denmark, Swe-								
den, and all Parts of the North, passing through			-	740		-		
the Spanish Netherlands, or United Provinces.	1	0	2	0	3	0	4	0
For all Letters and Pacquets to Hamburgh.	0	10	1	8	2	6	3	4

The Post Days from London, to any Part of England and Scotland, are Tuefdays, Thursdays, and Saturdays; and the Returns on Mondays, Wednesdays, and Fridays: But to Wales and Ireland, the Post goes only twice a Week, viz. Tuesdays and Saturdays, and returns from Wales every Monday and Friday; but from Ireland the Return is uncertain.

To France, Spain and Italy, Monday and Thursday.

To Holland, Germany, Denmark and Sweden, Tuesday and Friday.

To Flanders, and from thence to Denmark and Sweden, Monday and Friday. To Spain, by the Way of Falmouth from London, every other Tuesday, and from Falmouth by Packet-boats to the Groyne.

The Mails are due from foreign Parts as follows:

From Holland, Monday and Friday.
From France, Wednesday and Saturday.
From Flanders, Monday and Thursday.
From Ireland, Monday, Wednesday and Friday.
From Spain, Portugal, &c. once a Week.

By a late Order from the Postmaster-general, Packets and Letters are forwarded monthly from London to his Majesty's Islands in the West-Indies, and also to and from the several Colonies on the Continent of North-America, viz.

From London to the Islands of Barbadoes, Antegoa, Montserrat, Nevis, St. Christopher's, and Jamaica: And,

From London to New York, New-England, Virginia, Maryland, and all other Colonies on the Continent of North-America.



Geographical and Astronomical

PROBLEMS.

112. PROBLEM I.

*O find the Latitude of a Place given.

Example.

What is the Latitude of London?

Operation.

Guide your Eye either to the right or left Side of the Planisphere, under the Parallel of London;

and you have 51° 1, the Latitude required.

Examples for Practice.

Places.

Latitudes.

tuo Place given. To find the Difference of Latitude between

Example. What is the Difference of Latitude between Lon-

don in England, and Cape Frio in South America?

Operation. The Latitude of London is 51° ½ N. the Latitude of Cape Frio is 22° ½ S. whose Sum is 74°, the Difference of Latitude required. If the Places had been one in North, the other in South Latitude, their Difference would have been the Difference of Latitude required.

Examples for Practice.

Places. Places.

Diff. of Latitude.

Example. What is the Longitude of a Place given. Example. What is the Longitude of London, supposing the first Meridian to be that of the Planisphere?

Operation. Guide your Eye from thence to the Meridian of

London, and you have 17° 3, the Longitude East required.

Examples for Practice.

Places.

Longitudes.

the Meridian of London is the first Meridian.

Example. What is the Longitude of Prague, supposing Lon-

gitude to begin at the Meridian of London?

Operation. Count from the Meridian of London to that of Prague, and you have the Answer.

Note, If one Place be in one Planisphere, and the other Place in the other Planisphere, the Degrees from one Place to the other must be counted, and this is the Longitude either East or West as the Case shall be.

Examples for Practice. Supposing Longitude to begin at the Meridian of London.

Places.

Longitudes.

116. Problem V. To find the Difference of Longitude between two Places.

Example. What is the Difference of Longitude between, the Land's End, or West Point of England, and the Cape Good

Hope?

Operation. Count from the Meridian of the Land's End, to the Meridian of Cape Good Hope, and you have for the Difference of Meridians or Difference of Longitude required.

Ex

Places.

Examples for Practice.

Places. Diff. of Longitude,

117. Problem VI. To find a Place whose Longitude and Latitude are given.

Example. What is that Place whose Longitude is 21° 1/2 East

of London, and Latitude 45° 1 North?

Operation. Guide your Eye along the given Parallel of 45° ± North until you come to a Place 39° of Longitude East of London, and you have a Place near Belgrade, which was required.

Examples for Practice.

Latitudes. Latitudes. Places.

118, Problem VII. To find what o'Clock it is at one given a Place, when it is a given Hour at another given Place.

Example. What o'Clock it is at Canton in China, when it is

5 o'Clock Afternoon at London?

Operation. Count from the Meridian of London to the Meridian of Canton, and it is 113°, or 7 Hours 32 Minutes later, or 12 Hours 32 Minutes, the Hour required.

If one of the Places had been in the western Planisphere, you must have counted Westward, and it would have been sooner

than at London.

Examples for Practice.

Hours. Places. Places. Hours.

in a Country. To find the Number of square Miles

Example. How many square Miles are in Portugal?

Operation. Measure its Length, and apply it to the Meridian near that Place, it measures 5° ½, Measure the Breadth, apply it to the same, and you have 1° ½. Multiply the Length and Breadth together, and you have 8° ½, which multiplied by 3600, the square Miles in one square Degree, gives 29700 square Miles, the Answer. Irregular Countries should be divided into long Squares, and each measured.

Examples for Practice.

Places. Square Degrees. Square Miles.

gree of Longitude in any Parallel of Latitude.

Example. How many Miles go to a Degree of Latitude in

the Parallel of London?

Operation. Take 6 Degrees of Longitude near London, and apply it to its Meridian near the same Place, and you have 3 \(\frac{3}{4}\)
Degrees of the Meridian answerable thereto, which is so many Times Ten, or 37 \(\frac{1}{4}\) Miles, the Answer.

Examples for Practice.

Places.

Miles.

tude go to a Degree of Latitude or Distance in any Parallel.

Example. How many Degrees of Longitude make a De-

gree of Latitude or Distance in the Parallel of the Lizard?

Operation. Take 6 Degrees of Latitude, and apply it to the Parallel of the Lizard, both near the fame, and you have 9;

an-

109

answering thereto, which is so many Times Ten, or 93 Minutes, the Answer.

Examples for Practice.

Places.

Degrees.

122. Problem XI. To find how many Miles an Hour a Perfon moves in any Parallel of Latitude, by the Motion of the Earth round its Axis.

Example. How many Miles an Hour doth London move,

by the Motion of the Earth round its Axis, in a Day?

Operation. Take 15 Degrees of Longitude at London, and apply it to the Meridian there, so you have the Number of Degrees per Hour, which multiplied by 60, gives the Miles required.

Examples for Practice.

Places.

Miles per Hour.

Example. What Inhabitants are Antaci of a given Place.

Example. What Inhabitants are Antaci to the Arabians?

Operation. Guide your Eye to the same Degree of South
Latitude, as is the Latitude of the Arabian Inhabitants North,
and upon the same Meridian, and you will come to Madagascar
Isle, the Place required.

Examples for Practice.

Places.

Antæci.

THE UNIVERSAL

124. Problem XIII. To find the Periæci of a given Place. Example. What Inhabitants are Periæci to the Persians?

Operation. Guide your Eye along the Parallel of Persia, until you come to a Place 180 Degrees of Longitude therefrom, and you come to California, whose Inhabitants are those required.

Examples for Practice.

Places.

Periæci.

125. Problem XIV. To find the Antipodes of a given Place.

Example. What Inhabitants are Antipodes to China?

Operation. Guide your Eye along the Parallel of China, until you come to a Place 180 Degrees of Longitude therefrom, then guide your Eye to the same Degree of South Latitude, as is the Latitude of China North, and you come to Peru, whose Inhabitants are those required.

Examples for Practice.

Places.

Antipodes.

126. Problem XV. To find what Climate a given Place is in.

Example. What Climate is London in?

Operation. Guide your Eye along the Parallel of London, to the Side of the Planisphere, and you have the Beginning of the Ninth Climate required.

Examples for Practice.

Places.

Climates.

127. Problem XVI. Given the Month and Day to find the

Sun's Place in the Ecliptic.

Example. What is the Sun's Place in the Ecliptic, May 1st? Operation. Guide your Eye along the Ecliptic, and opposite to May 1st you have Taurus, 11° ½, the Place required. Otherwise you have the same in the Circumference of the Planisphere. Otherwise in the aforegoing Tables.

Examples for Practice.

Times.

Place in Ecliptic.

128. Problem XVII. Given the Month and Day to find the Sun's right Ascension.

Example. What is the Sun's right Ascension, April 25?

Operation. Guide your Eye along the Ecliptic until you tome to April 25 thereon, then guide your Eye downward along the Meridian of that Place, and you come to 33° ± on the Equator, the right Afcension required.

Examples for Practice.

Times.

Right Ascensions.

129. Problem XVIII. Given the Month and Day to find the Sun's Declination.

Example. What is the Sun's Declination, May 10?

Operation. Find May 10 in the Scale of Declination near the Side of the Planisphere, and on the graduated Meridians you have 17 \(\frac{3}{4}\) North, the Declination required. Otherwise it may be found in the aforegoing Tables.

Examples for Practice.

Times, Declination,

will be vertical on any Day at Noon.

Example. To what Parts of the Earth and Sea will the Sun

be vertical, June 21 at Noon?

Operation. Find June 21 in the Scale of Declination near the Side of both Planispheres, and guiding your Eye along the Parallel of Declination, you pass over Barbary, Arabia, India, Bengal, also the Atlantic Ocean, Mexico, and the Pacific Ocean, the Places required.

Examples for Practice.

Times.

Places.

131. Problem XX. To find where the Sun is vertical on a given Day, and at a given Hour.

Example. Where is the Sun vertical, or in the Zenith on

August 1st, when it is 9 o'Clock in the Morning at London?

Operation. First find the Parallel of Declination for August 1st, which is near 18° North; then, because the Sun is 3 Hours short of the Meridian of London at the given Hour, guide your Eye 3 Hours or 45 Degrees of Longitude Eastward on the said Parallel, and you come to Arabia Felix, the Place required.

Examples for Practice.

Times.

Places.

132. Problem XXI. To find how many Degrees the Sun will be North or South of the Zenith of a given Place, on a given Day.

Ex-

Example. How much will the Sun be South of the Zenith

of London, on July 3d?

Operation. For July 3, on the Scale of Declination, you have 23° North, the Sun's Declination, from which, counting on the Meridian of London to London, you have 28° ½, the Zenith Distance required.

Examples for Practice.

Times. Places. Zenith Distance.

133. Problem XXII. To find the longest Day and shortest Night at a given Place.

Example. What Length is the longest Day at London?

Operation. Find the Latitude of London, and it is 51° ½ North, then counting 51° ½ from the North Pole, stretch the Silken-thread to that Point in the Circumference, and observe where it cuts the northern Tropic; count from this Point to the Side of the Planisphere, and you have for one Side the half Day, and for the other Side the half Night, reckoning 15 Degrees of Longitude for an Hour, and a Degree of Longitude for 4 Minutes of Time; so you have 8 Hours 12 Minutes and ½ for half the longest Day, and 3 Hours 47 Minutes and ½ for the shortest Night required.

Examples for Practice.

Places. Longest Day. Shortest Night.

or fet at any given Place on a given Day.

Example. What Hour will the Sun rife and set at Lisbon,

May 30th?

at a given Place in the Problem

Operation. Find the Sun's Declination, May 30, and it is near 22° North; find the Latitude of Lisbon, and it is 39° North.

North. Count 39° from the North Pole of the Planisphere in the Circumference, and stretch the Silken-thread thereto, obferving where it cuts the Parallel of 22° North, from which Point, count to the Side of the Planisphere, so you have 4 Hours 44 1 Minutes the Time of Sun-rifing, and 7 Hours 15 1 the Time of Sun-fetting required.

Examples for Practice.

Times. Places. Rifing. Setting.

135. Problem XXIV. To find the Length of the Day or Night, at a given Place on a given Day.

Example. What is the Length of the Day, and also of the

Night at Copenhagen, May 20?

Operation. Find the Latitude of Copenhagen, and it is 55° \$ North. Find the Declination of the Sun for May 20, and it is 20 North: Find the Hour of Sun-rifing, as in the last Problem, it is 3 Hours 51 Minutes, which doubled, gives 7 Hours 42 Minutes the Length of the Night. Find the Hour of Sun-fetting, it is 8 Hours 9 Minutes, which doubled, is 16 Hours 18 Minutes, the Length of the Day.

Examples for Practice.

Places. Day Length. Night Length. Times.

126. Problem XXV. To find what Day the Sun begins to appear above the Horizon, at a given Place in the Frigid

Example. On what Day doth the Sun begin to appear at

Trinity-Isle near Greenland?

Operation. Find the Latitude of Trinity-Ifle, it is 710 1 N. Count 90 Degrees South therefrom on the Meridian, and you come to 18° 1 South, which Declination answers to January 27, the Answer.

Examples for Practice.

Places.

Times.

200 the Sun rifes or fets at a given Place, and on a given Day.

Example. On what Point of the Horizon will the Sun rife

and fet at London, January 30th?

Operation. Find the Sun's Declination, it is 17° \(\frac{3}{4}\) South; and the Latitude of London, which is 51° \(\frac{1}{2}\) North; stretch the Silken-string to 51 \(\frac{1}{2}\) in the Side of the Planisphere, and slide the Bead on the String until it cuts the Parallel of 17\(\frac{3}{4}\) South. Turn about the Thread until it coincides with the Equator or Axis of the World, and the Bead will measure of from the Centre, which is the Amplitude from the South required.

Examples for Practice.

Places. Times. Amplitudes.

of Time which depends upon the Sun's Place in the Ecliptic, and his Right Ascension.

Example. What is the Difference between the Sun's Place

in the Ecliptic, and his Right Ascension, April 20th?

Operation. Find the Sun's Place in the Ecliptic, it is \(\frac{3}{4}\) of a Degree in Taurus, hence his Longitude from Aries is 30° \(\frac{3}{4}\); guide your Eye downward on the Meridian to the Equator, and you come to 28° \(\frac{2}{3}\) of Right Ascension, hence the required Difference is 2° 5', or 8 Minutes 20 Seconds of Time.

Examples for Practice.

Times. Differences.

139. Problem XXVIII. To find the whole Equation of Time.

Example. What is the Equation of Time, April 20th?

Operation. Look in the Table of the Sun's Place, &c. and in the last Column thereof you have I Minute 18 Seconds, which good Clocks and Watches are after the Sun, the Answer.

Examples for Practice.

the state the Truesd until it commides with the trustor or

Courses which is the American Consults South required

or New Star will rife or fet; its Right Afcension and Declination, and the Latitude of the Place being given.

Example. What Hour will a Celestial Phenomenon rise and fet, whose Right Ascension is 38°, and Declination 20° South,

February 3d, at London?

Operation. Find the Sun's Declination, it is 16° 35' South, his Right Afcension is 316° 40'. Find the Sun's Semi diurnal Arch, it is 4 Hours 33'. And the Phenomenon's Semi diurnal Arch, 4 Hours 11'. But the Sun is near 5½ Hours of Right Ascension before the Phenomenon, therefore he will set about 4½ o'Clock, and the Phenomenon at about 10½ in the Evening.

Examples for Practice.

117

141. Problem XXX. To find all those Places where a Comet will be vertical, its Declination being given.

Example. Where will a Comet be vertical whose Declination

is 30° North?

Operation. Guide your Eye along the Parallel of 30° North, and you have China, India, Persia, Barbary, the Atlantic Ocean, Florida, and the Pacific Ocean required.

Examples for Practice.

on the eastern and western Planispheres.

Example. What is the Distance from London to Fort St.

George?

Operation. With a Pin or Needle, place the Centre of the Slider over the Centre of the eastern Planisphere, and turn the Slider round the Centre, until some Meridian of the Slider passeth through London and Fort St. George, then count from one of these Places to the other on the said Meridian, and you have 74 Degrees, the required Distance, which multiplied by 60, gives 4440 geographical Miles.

Examples for Practice.

Places.

Places.

Distances.

143. Problem XXXII. To find through what Countries a great Circle passeth, in leading from one given Place on the Planisphere, to another given Place on the same.

Example. Through what Countries doth a great Circle pass,

in leading from London to Jerusalem?

Operation. Place the Centre of the Index over the Centre of the eastern Planisphere, and turn round the Index until a Meridian thereon passeth through London and Jerusalem; then

will the same Meridian pass between France and Flanders, also through Part of Germany, Turkey in Europe, Archipelago, and the Bottom of the Mediterranean Sea, which are the Places required.

Examples for Practice.

one of which is in the Eastern, and the other in the western Planisphere.

Example. What is the Distance from the Lizard to Ja-

maica?

Operation. First find the Difference of Longitude between the Lizard and Jamaica, which being 72°, put a black Patch on the Eastern Planisphere at 72° East Longitude from the Lizard, and in the Parallel of Jamaica. Turn round the Index until a Meridian thereof passeth through the Lizard and the Patch, so you have 64° ½ by counting thereon from the Lizard to the Patch, for the Distance required, which multiplied by 60, gives 3870 Miles.

Examples for Practice.

Places. Places. Distances,

145. Problem XXXIV. To find those Parts of the Earth and Sea over which a great Circle will pass, in leading from one Place on one Planisphere, to another Place on the other Planisphere.

Example. Over what Parts of the Earth or Sea will a great

Circle pass, in leading from London to Cape Horn?

Operation. Find the Distance in the same Manner as the last Problem. Then, because London is 17° 4 of Longitude East from

from the Side of the Planisphere, observe how many Degrees of Longitude it is from London to where that Meridian which measures the Distance, cuts a Parallel of Latitude 17° 3 of Longitude East of London, which being the Parallel of 27° North; turn Z on the Index to 27° in the Side of the Planisphere, so the Meridian passing through London, gives the Answer for the Eastern Planisphere. In like Manner, bring Z to 27° in the Side of the Western Planisphere, and the Meridian passing thro' Cape Horn, over the Atlantic Ocean, is the other Part of the Anfwer required.

Examples for Practice. Places. Places.

Places.

146. Problem XXXV. To find the Polition of two Places with the Meridian, when both Places are on one Planisphere.

Example. Which Way doth Canton in China bear from Lon-

Operation. Turn round the Slider until a Meridian thereof paffeth through London and Canton, on which Meridian, at the Diltance of 90° from London place a Patch, count 90° Southward on the Meridian of London, and there also put another Patch; then turn the Slider round until a Meridian paffeth thro' both these Patches, on which Meridian you may count 120° from one Patch to the other, which is N. E. by E. & E. nearly. the Bearing required.

Examples for Practice. Places.

Places.

Bearings.

147. Problem XXXVI. To find the Polition of Places with the Meridian, when one Place lyeth in one Planisphere, and the other Place in the other Planisphere. ExExample. Which Way doth Barbadoes bear from Lbndon? Operation. Find the Difference of Longitude between London and Barbadoes, which being 59° West, place a black Patch at 59° Longitude, either Eastward from the Side of the Eastern Planisphere, or Westward from the Side of the Western Planisphere, and in the Parallel of Barbadoes; then counting London to be in the Parallel of 51° ½ at the Side of the Planisphere, proceed as in the last Problem, so you have the Bearing W.S. W. ½ W. nearly, which was required.

Examples for Practice.

Places. Bearings.

148. Problem XXXVII. To find where a great Circle, paffing over the Earth's Surface, will come nearest to the Pole.

Example. If a great Circle pass through London and the North Cape, where will it come nearest to the North Pole?

Operation. Turn the Slider until a Meridian thereof passeth through London and the North Cape, then will the same Meridian also pass nearest to the North Pole in Latitude 76° 3 North, and Longitude 76° East of London, which is a little North of the Middle of the Coast of Nova Zembla, the Place required.

Examples for Practice.

Places. Places. Nearest à Pole.

149. Problem XXXVIII. To draw a Meridian, or North and South Line at any Place on the Earth's Globe.

Example. I would draw a Meridian Line on the Floor of a

Room, whose Window faceth the South?

Operation. First, with a Pair of dividing Compasses, draw feveral Circles having the same Centre, but some wider than others:

others; erect a Pin or strait Wire in the Centre perpendicularly, and observe in the Forenoon, where the Shadow of the Top of the Wire falls on one of these Circles, making a Mark at the same, observe where the Shadow of the Top of the Wire falls on the same Circle after Noon, and a right Line drawn from one of the Marks to the other, is the East and West Line, from which the North and South is had by Bisection at Rightangles.

Examples for Practice.

150. Problem XXXIX. To find the Variation of the Mag-

Example. I would find the Variation of the Magnetic Needle

at London, or any other Place on Land.

Operation. First draw a Meridian Line, and a Circle round it; place a sharp Pin upright in the Centre, and the Magnetic Needle thereon, and it will shew the Variation required.

Examples for Practice.

151. Problem XL. To find a Meridian Line by the rifing and fetting of the Sun.

Example. I would find a Meridian Line by the rifing and

fetting of the Sun at London?

Operation. Observe on what Point the Sun riseth, and also on what Point it sets, the Middle of these two is the Meridian required.

Examples for Practice.

a terrestrial Object bears by the Planisphere.

Example. Suppose it was required to find the Bearing of

Greenwich from Chelsea?

Operation. Place the Planisphere horizontally with its Poles North and Southward by help of a Meridian Line, then by the Degrees or Points of the Compass in the Edge of the Planisphere, will Greenwich bear nearly East from Chelsea, which was required.

Examples for Practice.

153. Problem XLII. To find all those Places of the Earth where the Sun is either rising or setting, at a given Time.

Example. At what Place of the Earth is the Sun rifing or fetting, May 1st, at 10 o'Clock in the Morning, at London?

Operation. The Sun's Declination, May 1st, is 15° ½ North, and at 10 in the Morning, he is 2 Hours short of the Meridian of London; therefore count 30° Eastward from the Meridian of London along the Parallel of 15° ½ North, and you have the Place in Africa where the Sun is vertical. Bring the Horizon or Equator of the Slider to pass through the Place where the Sun is vertical, and count 90° on the same, so you come to a Meridian passing through New Holland, Celebes, Minando, the Sea of Corea, and North-east Parts of Tartary, which shews the Places of Sun-setting. Then proceed as in the 34th Problem, and you have a great Circle passing between Paraguay and Brazil, near the Mouth of the River Amazones, the Caribee Islands, New England, and Hudson's Bay, the Places where the Sun is setting, which was required.

Examples for Practice.

154. Problem XLIII. To find those Places of the Earth, which are in the enlightened Hemisphere, at a given Time.

Example. What Places of the Earth are in the enlightened

Hemisphere, August 20th, at 2 in the Afternoon?

Operation. Find all those Places which are within 90° of the Place where the Sun is vertical at the given Time, as in the last Problem, and you have the enlightened Part, the Rest is dark Night or Twilight.

Examples for Practice.

and Sea, where a folar Eclipse can be seen, the Day and Hour when it will happen being given.

Example. To what Parts of the World may the Eclipse of the Sun be visible, which will happen August 5th, 1766, at

7 in the Evening?

Operation. Find what Part of the Earth is the enlightened Hemisphere at the given Time, as in the preceding Problems.

Note, A folar Eclipse is not always visible throughout the enlightened Hemisphere, on Account of the Parallax.

Examples for Practice.

124

156. Problem XLV. To find all those Places of the Earth and Sea, where a lunar Eclipse may be seen, the Day and Hour when it will happen being given.

Example. To what Part of the World may the Eclipse of the Moon be seen, which will happen August 20th, 1766, at

7 in the Morning, at London?

Operation. Find all those Places which are within 90° Distance from the Place where the Moon is vertical at the given Time, supposing her directly opposite to the Sun, for those Places of the dark Hemisphere within 90° of the Place where the Moon is vertical, are the Places where the lunar Eclipse is visible.

Examples for Practice.

157. Problem XLVI. To find all those Places which have two Summers and two Winters in one Year.

Example. What are those Places of the Earth's Globe?

Operation. Find those Lands which are under the equinoctial Line, and they are the Places required.

Examples for Practice.

158. Problem XLVII. To find the Latitude of a Place of the Earth.

Example. I would find the Latitude of any Place.

Operation. Take the greatest, and also the least Altitude of the Pole-star, or any other Star near the Pole, when he is above, and also when he is under the Pole, which Time may be known by his Right Ascension, and the half Sum of these two Altitudes,

tudes, (Allowance being made at both Times for Refraction) is the Latitude required.

Examples for Practice.

on the Earth or Sea.

Example. I would find the Longitude of any Place.

Operation. Observe the Hour and Minute of either an Eclipse of the Moon, an Eclipse of one of Jupiter's Satellites, or Appulse of the Moon to a fixed Star, at the same Instant of Time the same Phenomenon is observed at London or any other where, and the Difference of Time converted into Degrees and Minutes, is the Difference of Longitude.

Examples for Practice.

160. Problem XLIX. Given the Day of the Month and Latitude of the Place, to find the Hour of Beginning and End of Twilight.

Example. I would know at what Time of the Night the

Twilight will end at London, April 15?

Operation. Turn Z on the Slider to the Latitude 51° ½ on the Planisphere, and guide your Eye downward along the Parallel of 10° North Declination on the Planisphere, until you come to 18° of the Meridian Lines of the Slider under its Horizon, so you come to a Meridian of the Planisphere 43° ½ short of Midnight, consequently the Beginning of Twilight is at 2 Hours 54 Minutes in the Morning, and the End at 9 Hours 6 Minutes in the Evening.

Places. Examples for Practice.
Times.

Twilight.

161. Problem L. Given the Latitude of a Place, to find what Day of the Year the Twilight is shortest at that Place.

Example. On what Day of the Year is the Twilight shortest

at London?

Operation. Turn Z on the Slider, to the Latitude of 51° ½ on the Side of the Planisphere, then with a Pair of dividing Compasses, or otherwise, find out the shortest Arch of a Parallel of Declination on the Planisphere, that is contained between the Horizon of the Slider, and the Parallel of 18° under the same Horizon, so you find the said shortest Arch of Declination on the Planisphere to be 20°½ North of the Equator, which answers to May 22, and July 21, the Days required.

Examples for Practice.

Places.

Shortest Twilight.

of the Place, and Altitude of the Sun, to find the Hour of the Day.

Example. If the Altitude of the Sun be taken at London, August 11th, and found 11° 15', what is the Hour of the

Day?

Operation. Subtract the Refraction from the given Altitude, and you have 11° 10' for the true Altitude. Then turn Z on the Slider, to the Latitude 51° ½ on the Planisphere, and guide your Eye along the Parallel of 15° 7' North Declination, answering to the given Day, until you come to 11° 10' Altitude on a Meridian of the Planisphere, which shews that the Hour

was either 2 Minutes before 6 in the Morning, or 2 Minutes after 6 in the Afternoon.

Places.

Examples for Practice. Altitudes. Hours. Days.

163. Problem LII. Given the Latitude of the Place, the Declination and Altitude of the Sun, to find the Azimuth.

Example. If the Altitude of the Sun be taken at London. June 21, and found 50° after Allowance for Refraction, what is the Azimuth?

Operation. Turn Z on the Slider, to the Latitude of 51° = in the Side of the Planisphere, and guide your Eye downward along the Parallel of 23° 1 North Declination on the Planifphere, until you come to 50° Altitude above the Horizon, among the Meridians of the Slider, so you come to a Meridian on the Slider 59° 3 from the Side of the same, which is the Sun's Azimuth from the South required.

Examples for Practice.

Days. Altitudes. Azimuths. Places.

164. Problem LIII. Given the Day of the Month, and the Difference between the Altitude of the Sun taken at one Time of the Day, and the Altitude taken at another Time of the fame Day; also the Difference between the two correspondent Azimuths taken at the fame Times the Altitudes were taken, to find the Latitude of the Place.

Example. On the 21st of July, the Altitude of the Sun was taken in the Forenoon, and his Azimuth taken at the same Time; some Time after the same Afternoon, the Altitude was taken, and also his Azimuth at this second Observation; and

here-

hereby the Difference of 'Altitudes were found 15° 5', and the Difference of Azimuths 32° 0'; what was the Latitude of the Place of Observation?

Operation. Turn the Slider, until the Parallel of 20° 23' North on the Planisphere passeth through two Meridians of the Slider, and such Points in them, as at the same Time do form the given Difference of Altitudes above the Horizon of the Slider, and the given Difference of Azimuths on the same; so you have Z on the Slider placed at 55° in the Side of the Planisphere, the Latitude North which was required.

Examples for Practice.

on the Planisphere. To find the Latitude of a Star given

Operation. Turn Z on the Slider to the Pole of the Ecliptic, fo the Horizon of the Slider represents the Ecliptic, and you have Syrius on the Planisphere, 39° \(\frac{1}{2}\) South of the Ecliptic.

tic, which is the Latitude required.

Examples for Practice.

Stars.

Latitudes.

166. Problem LV. To find the Longitude of a Star given on the Planisphere.

Example. What is the Longitude of Capella in Auriga?

Operation. Turn Z on the Slider, to the Pole of the Ecliptic, as in the last Problem, and you have Capella on the Planifiphere, 78° ½ from the Meridian passing through the Centre of

the

the Slider, which is the equinoctial Colure; therefore the Longitude of Capella is II 180 1, which was required.

Examples for Practice.

Stars.

Longitudes.

167. Problem LVI. Given the Latitude and Longitude of a Comet, Planet, or Star, to find its Right Afcension and Declination.

Example. What is the Right Ascension and Declination of a celeftial Body, whose Latitude is 61° 45' North, and Longi-

tude vº 11° 30'?

Operation. Turn Z on the Slider to the Pole of the Ecliptic, on the Western Planisphere of the Heavens, and counting 11° from the Left-hand Side of the Slider, among the Meridians of the fame, you come to a Place 61° 3 North of the fame Horizon, under which is a Place, whose Right Ascension is 267° 1, and Declination 38° 1 North, which was required.

Examples for Practice. Latitudes. Longitudes. Right Afcen. Declin.

168. Problem LVII. To find the Distance of two Stars that are both on one Planisphere.

Example. It is required to find the Distance of any two Stars

that are both on the Eastern or Western Planisphere?

Operation. Turn the Slider round, until a Meridian thereof passeth through both the given Stars, and the Number of Degrees, counted on the fame Meridian, is the Distance required. P

Ex-

Examples for Practice.
Stars. Stars.

Distances.

one of which is on the Eastern, and the other on the Western celestial Planisphere.

Example. It is required to find the Distance of any two Stars, one of which is in one Planisphere, and the other of those Stars

in the other Planisphere?

Operation. Place a black Patch in the same Parallel of Declination, and Difference of Right Ascension in one Planisphere, as is the Declination and Difference of Right Ascension of the other Star in the other Planisphere; find a Meridian passing thro' this Patch and the Star in the same Planisphere, and it is the Distance required.

Examples for Practice.
Stars. Stars.

Distances.

170. Problem. LIX. Given the Direction of any celeftial Phenomenon passing through any Part of the Heavens in the Arch of a great Circle, to find the Path of the same.

Example. Suppose a Comet was first seen under the Belly of the great Bear, and did bend its Course toward Arcturus in Bootes; what Constellations would it pass thro', supposing its

Trace to be nearly in the Arch of a great Circle?

Operation. Turn round the Slider, until a Meridian thereof passeth through the Places given, so will the same Meridian pass through Mons Menalus, Serpentarius, &c. the Places required.

Ex-

Examples for Practice.

171. Problem LX. Given a Star on the Planisphere, to find its Latitude, Longitude, Right Ascension and Declination, for any Year past, present, or to come.

Example. What Latitude, Longitude, Right Ascension and Declination, will the Pole-star have 1000 Years hence?

O'eration. Find how many Degrees and Minutes the Pole-star is got forward in the Ecliptic, allowing at the Rate of 50 Seconds per Annum. Then turn Z on the Slider to the Pole of the Ecliptic, and count those Degrees forward under the Star's Parallel of Latitude, so you come to a Place on the Planisphere, which shews the Right Ascension and Declination required.

Examples for Practice.

172. Problem LXI. Given the Day of the Month, to find what Hour a Star, Planet, or other celestial Body, whose Place in the Heavens is known, will come to the Meridian.

Example. At what Time will the Pleiades, or Seven-stars, come

to the South at London, August 10th?

Operation. Find the Right Ascension of the Sun for the given Day, and it is 140° ½, count from thence backward, in a Direction contrary to the Order of the Signs, until you come to the Right Ascension of the Seven-stars, and you count 53° ½, which subtracted from 140° ½, leaves 87° ¼, or 5 Hours 49' after Noon, the Time required.



175. Problem LXIV. To find the oblique Ascension, of any celestial Body.

Example. What is the oblique Ascension of the Sun, May

20th, for the Latitude of 51° 35' North?

Operation. Find that Point of the Equator which comes to the Meridian when the Sun or Star riseth, add 6 Hours thereto, and the Sum is the oblique Ascension required. Hence the Answer is 30° 15'.

Examples for Practice.

176. Problem LXV. To find the oblique Descension of any celestial Body.

Example. What is the oblique Descension of the Sun, May

20th, for the Latitude of London?

Operation. Find that Point of the Equator which comes to the Meridian when the Star fets, subtract 6 Hours therefrom, and the Difference is the oblique Descension required.

Examples for Practice.

177. Problem LXVI. To find the cosmical Rising and Setting of a given Star.

Example. What Time of the Year doth Syrius rise or set

cosmically, at London?

Operation. Find that Point of the Ecliptic which rifes, when the Star rifes or fets, and you have the Answer August 20, and February 24.

Examples for Practice.

133

134 THE UNIVERSAL

178. Problem LXVII. To find the anchronical Rising and Setting of a given Star.

Example. What Time of the Year doth Syrius rise and set

anchronically at London?

Operation. Find that Point of the Ecliptic which fets at the same Time as the Star rises or sets, and you have the Answer February 6th, and May 14th.

Examples for Practice.

179. Problem LXVIII. Given the three Sides of a spherical Triangle, to find the Angles by the Planisphere.

Example. Given the three Sides of a spherical Triangle 38°

28', 95° and 76°, to find an Angle?

Operation. Bring Z on the Slider to so many Degrees from the Pole of the Planisphere as is one of the given Sides, and count downward from Z toward the Horizon of the Slider, until you have a Point where a vertical Circle of the Slider, and a Meridian of the Planisphere, answer the other two Sides of the Triangle, then one of the Angles is shewn on the Horizon of the Slider, another on the Equator of the Planisphere, and the third Angle may be known by shifting the Sides. Hence the required Angle is 60°.

Use. By this Problem, having the Latitude, Sun's Altitude and Declination, are found the Azimuth, Angle of Position,

and Hour of the Day.

180. Problem LXIX. Given the three Angles of a spherical Triangle, to find the three Sides by the Planisphere.

Example. Given three Angles of a spherical Triangle, 107°

36', 56° 26', and 37° 55', to find an Angle?

Operation. Take the Compliments of the three given Sides to a Semi-circle, and proceed with them as if they were Sides finding the Angles for the Sides required. Hence the required Side is 74° 50'.

Use. By this Problem, having the Hour of the Day, Azimuth, and Angle of Position, are found the Latitude, Sun's

Declination and Altitude.

Examples for Practice.

181. Problem LXX. Given two Sides of a spherical Triangle, and the Angle between them, to find the other Side and two Angles.

Example. Given two Sides of a spherical Triangle, 38° 28', 74° 50', and the Angle between them 56° 26', to find the other

Side and two Angles.

Operation. Bring Z on the Slider, to so many Degrees from the Pole of the Planisphere, as is one of the given Sides, and count downward from Z, under the given Angle on the Horizon of the Planisphere, so many Degrees as the other given Side; then directing your Eye on that Meridian of the Planisphere, to its Equator, you have one of the Angles, counting from thence to the Pole of the Planisphere, you have the required Side; and by shifting the Parts, you have the other Angle required. Hence the required Side is 57° 32', and one of the required Angles 107° 36'.

Use. By this Problem, having the Latitude, Azimuth, and Altitude of the Sun, are found the Sun's Declination, Hour of the Day, and Angle of Position; also, having the Sun's Altitude, Angle of Position and Declination, are found the Hour of the Day, Latitude and Azimuth; also, having the Sun's Declination, Hour of the Day, and Latitude of the Place, are found, the Azimuth, Altitude and Angle of Position.

Examples for Practice.

182. Problem LXXI. Given two Angles of a spherical Triangle, and the Side between them, to find the two other Sides and Angle.

Example. Given two Angles of a spherical Triangle 107° 36', and 56°26', and the Side between them 38° 28', to find

the other Angle and Sides?

Operation. Bring Z on the Slider, to so many Degrees from the Pole of the Planisphere as is the given Side, and count downward from the Pole on the Planisphere, and from Z on the Slider, under the two given Angles, then from the Place where they meet, count upward on a vertical Circle of the Slider, and also on a Meridian of the Planisphere, to the Pole and to Z, and you have the two Sides required; the required Angle may be found by shifting the Sides. Hence the required Angle gle is 37° 55', and one of the other Sides 57° 32'.

Use. By this Problem, having the Hour of the Day, the Latitude of a Place, and Sun's Azimuth, are found the Sun's Altitude, Angle of Position, and Declination; also, having the Azimuth, Altitude and Angle of Position, are found the Declination, Hour of the Day and Latitude; also, having the

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Angle of Position, Declination and Hour of the Day, are found the Latitude, Azimuth and Altitude.

Examples for Practice.

angle, and an Angle opposite to one of those two Sides, to find the other Side and two Angles.

Example. Given two Sides of a spherical Triangle 38° 28', and 57° 32', and the Angle opposite to the greater Side 56°

26', to find the other Angle and Sides?

Operation. Bring Z on the Slider, to so many Degrees from the Pole of the Planisphere, as is one of the given Sides, and under the given Angle, count downward from the Pole of the Planisphere, on a Meridian, until you come to the other given Side on a vertical Circle of the Slider, then the Distance on the Meridian from the Pole of the Planisphere, is the Side required; one of the Angles sought, is shewn on the Horizon of the Slider, and the other Angle required, is sound by shifting the Sides. Hence one of the required Angles is 107° 36′, and the required Side 74° 50′.

Use. By this Problem, having the Latitude of a Place, Hour of the Day, and Sun's Altitude, are found the Sun's Azimuth, Angle of Position and Declination; also, having the Latitude, Altitude, and Angle of Position, are found the Azimuth, Declination, and Hour of the Day; also, having the Sun's Altitude, Declination and Azimuth, are found the Latitude, Hour of the Day, and Angle of Position; also, having the Declination, Altitude and Hour of the Day, are found the Latitude, Azimuth and Angle of Position; also, having the Latitude, Declination and Angle of Position, are found the

Hour

Hour of the Day, Azimuth and Altitude; also, having the Latitude, Declination and Azimuth, are found the Hour of the Day, Altitude and Angle of Position.

Examples for Practice.

184. Problem LXXIII. Given two Angles of a spherical Triangle, and a Side opposite to one of them, to find the other Sides and Angle.

Example. The two Angles given, are 107° 36', and 56° 26', and the given Side 57° 32', required the other Angle and

Sides?

Operation. Turn Z on the Slider, until a Meridian on the Planisphere, with one of the given Angles, cuts off the given Side on a vertical Circle of the Slider, and the same vertical Circle shews the other given Angle on the Horizon of the Slider; then one of the required Sides is the Distance of Z on the Slider from the Pole of the Planisphere, the other required Side, is the Distance of Z from the Pole of the Planisphere, and the required Angle is found by shifting the Sides. Hence the required Angle is 37° 55', and the required Side 38° 28'.

Use. By this Problem, having the Sun's Altitude, Azimuth, and Hour of the Day, are found the Latitude, Declination, and Angle of Position; also, having the Declination, Azimuth, and Angle of Position, are found the Latitude, Altitude, and Hour of the Day; also, having the Altitude, Hour of the Day, and Angle of Position, are found the Latitude, Azimuth, and Declination; also, having the Declination, Hour of the Day, and Azimuth, are found the Latitude, Altitude, and Angle of Position; also, having the Latitude, Azimuth, and

Angle

139

Angle of Position, are found the Altitude, Declination, and Hour of the Day; also, having the Latitude, Angle of Position, and Hour of the Day, are found the Declination, Altitude and Azimuth.

Examples for Practice.

185. Problem LXXIV. To know the Latitude of the Moon, Planets, Comets, or other celestial Bodies.

Example. Required the Moon's Latitude at any given Time

of the Night when she is visible?

Operation. Observe any two Stars, that are in the Arch of a great Circle with the Moon, Planet, Comet, or other celestial Body to be observed, which crosseth the Ecliptic at Right-angles, and note those two Stars on the Planisphere, then by an Instrument (such as Hadley's Quadrant) take the Distance from either of those Stars, which added to, or subtracted from the Star's Latitude, gives the Latitude required.

186. Problem LXXV. To know the Longitude of the Moon, Planets, Comets, or other celestial Bodies.

Example. Required the Longitude of the Moon at any given

Time of the Night when she is visible?

Operation. Observe any two Stars, that are in the Arch of a great Circle with the Moon, Planet, Comet, or other celestial Body to be observed, which crosseth the Ecliptic at Right-angles, and note those two Stars on the Planisphere, then turn the Slider until some Meridian thereon passeth thro' those two Stars on the Planispheres, and the Point where it cuts the Ecliptic, is the Longitude required.

Examples for Practice.

the magnetic Needle (commonly called the magnetic Azimuth) the Latitude of the Place and Sun's Declination, to find the Variation of the Compass.

Example. In Latitude 30° North, May 30th, in the Morning, the Sun's Altitude was taken 30°, and his Bearing, by the magnetic Needle, at the same Time was due West, required

the Variation,

Operation. Find the true azimuth; Place, or suppose to be placed the true Azimuth over the magnetic Azimuth, and the magnetic North-point falling Eastward or Westward of the true North-point of the Horizon, shews the Variation required.

188. Problem LXXVII. Given the Latitude of a Place in the torrid Zone, to find the Sun's greatest Azimuth on a given Day, when the Sun is nearer the Pole than the Place given; also the Altitude of the Sun at the same Time.

Example. What is the Sun's greatest Azimuth, at Cape Three Points, on the Coast of Africa, June 20th, and his Altitude at

the same Time?

Operation. Bring Z on the Slider, to the Latitude given on the Planisphere, observing where the Convexity of the Parallel of Declination on the Planisphere toucheth an azimuth Circle on the Slider, you have the required Azimuth on the Horizon of the Slider and the Altitude of the Sun at the same Time on the azimuth Circle of the Slider, by counting upward from the Horizon.

Examples for Practice.

189. Problem LXXVIII. Given the Sun's Declination, Al-

titude, and greatest Azimuth, to find the Latitude.

Example. At Sea, in north Latitude, the 20th of June, the Sun's Altitude was taken, when he was on the greatest Azimuth from the North, and found to be 60°, and the greatest Azimuth at the same Time 75°. Quere the Latitude?

142 THE UNIVERSAL

Operation. Move the Slider until the given Altitude thereon toucheth the Convexity of the Parallel of Declination on the Planisphere, and at Z on the Slider, you have the Latitude on the Planisphere.

Examples for Practice.

190. Problem LXXIX. To draw an horizontal Dial for any Latitude by the Planisphere.

Example. Suppose an horizontal Dial was to be drawn for the Latitude of 45° North, what would be the Hour, Half-

hour, Quarter-hour, &c. Distances from the Meridian?

Operation. Turn Z on the Slider to the Latitude on the Planisphere, and observe where 15°, 30°, 45°, &c. also 7° \(\frac{1}{2}, 22° \frac{1}{2}, 37° \frac{1}{2}, \text{ of the Meridians on the Planisphere do cut the Horizon of the Slider, so you have 10° \frac{3}{4}, 22° \frac{1}{4}, 35° \frac{1}{4}, \text{ for the Hour-distances, &c. required.}

Note, The Gnomen, or Style of the Dial, must be placed in a Direction parallel with the Axis of the World.

142

191. Problem LXXX. To draw an upright fouth or north Dial, by the Planisphere.

Example. Suppose an upright fouth or north Dial was to be

drawn for Latitude 55° North?

Operation. These Dials will become horizontal Dials, the former in Latitude 35° South, the later in Latitude 35° North; therefore find the Hour Distances for these Places, as in the last Problem, and you have the Answers.

Examples for Practice.

192. Problem LXXXI. To draw a declining or reclining Dial, by the Planisphere.

Axist and must the Slow-man where they are proje

Place the Caroties in the Date, for of the?

Note, A declining Dial, is perpendicular to the Horizon, but faceth about more toward the East or West than a South or North Dial; and a reclining Dial is neither perpendicular to the Horizon, nor parallel thereto, but faceth the Heavens obliquely.

Example. Suppose a Dial was to be drawn on a Plane, declining any Number of Degrees toward the East or West, and reclining any Number of Degrees from the Zenith?

Operation. First place the Gnomen in the Direction of the Axis of the World, and then draw the Hours by the Shadow

of the Sun.

Otherwise. Project the Meridians on the Horizon of that Degree of Latitude and Difference of Longitude, answering to the Plane's, Declination and Reclination.

Examples for Practice.

193. Problem LXXXII. To draw Hour-lines on any Surface whatfoever.

Example. Suppose it was required to draw Hour-lines on the

Cieling of a Room, or any other uneven Surface?

Operation. Place the Gnomen in the Direction of the Earth's Axis, and mark the Hour-lines where they are projected by the Sun.

Examples for Practice.

194. Problem LXXXIII. Having the Latitude of your Habitation, and Meridian-line, to determine the Hour of the Day by the Sun.

Example. Suppose I would know the Hour of the Day,

having a Meridian-line in any known Place of the Earth?

Operation. Place the Planisphere North and South, and parallel to the Horizon, hang up a Plumb-line, and take the Azimuth; then bring Z on the Slider to the Latitude on the Planish

Planisphere, and guiding your Eye from the observed Azimuth on the Horizon of the Slider upward, until it cuts the Parallel of the Sun's Declination for that Day on the Planisphere, guide your Eye from thence to the Equator of the Planisphere, and you have the Hour required.

Examples for Practice.

195. Problem LXXXIV. Required the following Answers, for the Latitude of 50° North, on May 16, viz.

	Answers.	
Sun's Meridian Altitude — — —	59°	10'
Sun's Amplitude — — —	32	30
Sun's Height at 6 o'Clock	15	10
Sun's Azimuth at 6 o'Clock, North-easterly —	47	15
Sun's Height being East and West	26	20
Hour when East — — —	7h	20'
Ascensional Difference ———	26°	20'
Right Ascension — — — —	53	10
Hour of Sun-rifing — — —	-4h	15'
Hour of Sun-fetting	7	45
Length of the Day	15	30
Length of the Night	. 8	30
Longest Day in that Latitude	16	10
Shortest Day in that Latitude	7	50
Hour of Day-break — — — —	. I	59
Duration of Twilight	2	15

146 THE UNIVERSAL

196. Problem LXXXV. Required the following Answers, for the Latitude of 50° North, on November 23, viz.

Total Control of the	Answers.	
Sun's Meridian Altitude — —	19°	30'
Sun's Amplitude South-eafterly — —	33	25
Sun's Depression at 6 o'Clock ———	15	4
Sun's Azimuth at 6 o'Clock, South-easterly	77	0
Sun's Depression East and West	27	30
Hour when East — — — —	- 4h	42'
Right Afcention ———	237	35
Hour of Sun-rifing — — —	7	50
Hour of Sun-fetting — — —	4h	10
Length of the Day	8	20
Length of the Night ———	15	40
Shortest Day in that Latitude	7	46
Longest Night in that Latitude — —	16	14
Hour of Day-break — —	5	48
Duration of Twilight — —	2	18

197. Problem LXXXVI. Required the following Answers, for the Latitude of 35° South, on November 10, viz.

	Answers.	
Sun's Meridian Altitude	72°	10'
Hour of Sun-rifing — — —	5h	9'
Hour of Sun-fetting -	6	51
Sun's Afcentional Difference	0	51
Length of the Day	13	42
Length of the Night	10	18
Sun's Right Afcention — —	225°	40
Sun's Amplitude — — —	80	55
Sun's Height at 6 o'Clock — —	9	50
Sun's Azimuth at 6 o'Clock, South-easterly -	70°	0'
Sun's Height when East and West	30	0
Hour of Day-break — —	3h	36'
Hour of Twilight ending	8	24
Duration of Twilight — —	1	33
Hour when Sun is East — —	7	44
Hour when Sun is West	4	16
Length of the longest Day	14	26
Length of the shortest Night	18	20
	10	20

198. Problem LXXXVII. Required the following Answers, for the Latitude of 45° South, on July 29, viz.

THE RESERVE OF THE PARTY OF THE	Answers.	
Sun's Meridian Altitude — —	26°	261
Sun's Depreffion at 6 o'Clock ——	14	15
Sun's Azimuth at 6 o'Clock, North-easterly -	75	50
Sun's Depression East and West — —	28	30
Sun's Amplitude North-eafterly —	61	50
Sun's Right Afcension ————	129	
Sun's Afcentional Difference — — —	1h	24
Hour of Sun-rifing — — —	7	24
Hour of Sun-fetting —	4	16
Length of the Day — — —	14	48
Length of the Night — —	9	12
Hour of Day-break ———	6	24
Duration of Twilight	1	48
Hour when Sun is East ———	4	32
Hour when Sun is West ———	7	28
Length of the shortest Day	8	34
Length of the longest Day	15	25
AND THE RESIDENCE OF THE PARTY		

199. Problem LXXXVIII. To find If it be Leap-year, or

how many Years after.

Example. How many Years is it after Leap-year, Anno 1770? Operation. Divide the Year by 4, and if o remain, it is Leap-year, but if 1, 2, 3 remains, it is fo many Years after Leap-year.

Note, The Years 1800, 1900, 2100, 2200, 2300, 2500, & &c. are common Years.



203. Problem XCII. To find the Roman Indiction.

Example. What is the Roman Indiction for the Year of our

Lord 1770?

Operation. To the Year of our Lord add 3, and divide the Sum by 15, the Remainder is the Roman Indiction required.

Examples for Practice.

204. Problem XCIII. To find the Epact, New-style, for a given Year.

Example. What is the Epact for the Year of our Lord

1770?

Operation. Multiply the Golden-number by 11, and divide the Product by 30, the Remainder is the Epact, Old-style. Then from the Epact, Old-style (augmented by 30, if Need require) subtract the Days of Difference between the Old and New-style, and you have the Epact, New-style.

Examples for Practice.

205. Problem XCIV. To find the Age of the Moon?

Example. What is the Age of the Moon, January 1, 1770?

Operation. Add to the Epact for March 1, for April 2, for May 3, for June 4, &c. and also add the Day of the Month; the Sum of these three Numbers, or their Overplus above 30, or Thirties, is the Moon's Age required.

Examples for Practice.

150 THE UNIVERSAL

206. Problem XCV. To find the Time of the Moon's fouthing.

Example. What Time will the Moon be South, January 1,

1770?

Operation. Multiply the Age of the Moon by 4, and divide the Product by 5; the Quotient is the Hour of the Moon's fouthing required.

Examples for Practice.

207. Problem XCVI. To find the Time of High-water at a given Place.

Example. What Time will it be High-water at London-bridge,

January 1, 1770?

Operation. Find how many Hours the Moon is past the Meridian, when it is High-water by the Tide-table, add these to the Hour of her Southing, and you have the Hour of Highwater required.

Examples for Practice.

208. Problem XCVII. To find Easter-day.

Example. What Month and Day will Easter-day be on,

Anno 1780?

Operation. Find how many Days old the Moon is, March 20, in the given Year, if it be Leap-year, otherwise the 21st; and if the Moon is 14 Days old on the said 20th or 21st of March, find what Day of the Week she is so old, and the Sunday sollowing is Easter-day. But if she is not 14 Days old on the said 20th or 21st of March, then count forward to that Day of the Week when she is 14 Days old, and the Sunday sollowing is Easter-day.

Ex-

Examples for Practice.

209. Problem XCVIII. To conduct a Ship from one given Port to another, along the Arch of a great Circle.

Note, By this Kind of Sailing, a Ship fails to her defigned Port the shortest Way she possibly can, which, if the Place she fails from, and that she fails to, be both in high Latitudes of the same Kind, may be shorter than by Mercator's failing by a sixth Part of the whole Voyage, and the Longitude may hereby be more correctly ascertained. Besides which, this Tract being out of the usual Course, Merchant Ships may hereby often escape their Enemy. The principal Objections to this Kind of failing, have been, that the Winds will not always permit it to be practised, and that it requires a tedious Calculation of spherical Triangles, the former of which can have no Weight in the Tradewinds, and the latter may be easily removed by the Planisphere.

Example. Required the Latitudes, and Differences of Longitude by which a Ship must pass, in sailing from Latitude 32° 30' North, to Latitude 50° 0' North, the Difference of Longitude 70° 0', to keep on the Arch of a great Circle, by the Pla-

nifphere.

Operation. Find two Places on the Planisphere, whose Latitudes and Difference of Longitude agree with those of the given Ports, turn the Slider until some Meridian thereof passeth through those two Places, and you have the Degrees of Latitude and Difference of Longitude on the Planisphere, by which the great Circle passeth, of which make a Table for every 1, 2, 3, 4, 5, 10, &c. Degrees of Latitude and Longitude. Then for the Course at any Place in the Voyage; turn Z on the Slider, to the Latitude of one of the Ports on the Planisphere, and counting downward from the Pole of the Planisphere, under the Angle of their Difference of Longitude, to the Latitude of the

other given Port, on a Meridian of the Planisphere, you have the Course on the Horizon of the Slider. Hence the Answers for the Latitudes and Difference of Longitudes by which the Arch passeth, are as follows, viz.

Diff. Long.	1 at.		Diff. Long.			Lat.	
5° —	35° 5	0'	40°	-	_	49°	5
10	38 5	0	45	-	-	49	50
15	41 2	5	50	-	-	50	15
20 — —		35	55	N-	-	50	30
25 — —	45 2	5	60	_	-	50	35
30 — —	46 5	5	65			50	25
35 —		5	70	-	-	50	0

of a Place at Sea, where the Variation Lines are nearly North and South, to find the Variation of the magnetic Needle at that Place for a given Year.

Note, 1st. The Variation of the magnetic Needle, is the Declination of its North Point, either Eastward or Westward from the true North Point of the Horizon, when the Needle is touched by the Loadstone, or artificial Magnet, balanced, and freely suspended Parallel to the Horizon.

2d. This Variation is continually increasing or decreafing, at all Places on the Surface of the Earth and Sea, by a Law not yet fully discovered, whereby it is found greater at one Place than at another, and greater in one Year than another Year at the same Place, and sometimes half a Degree greater at one Time of the Day than another Time of the same Day, at the same Place.

Example. Suppose it be required to know the Variation of the Needle at Cape Good Hope for the Year 1765?

Operation. Observe how many Degrees of Longitude the Line of Variation hath moved Eastward or Westward in the given Number of Years for which it is marked; and say, by the Rule of Proportion, as the given Interval of Years, is to the given Difference of Longitude, to where the same Variation Line is removed; so is the Time elapsed from one of the Years for which the Variation Lines were delineated, to the Difference of Longitude, which the same Variation Line hath made in that Time, either increasing or decreasing, as the Case shall happen.

Examples for Practice.

211. Problem C. Given the Latitude and Variation, to find the Longitude where the Variation Lines run nearly North and South.

Example. Suppose a Ship at Sea, near the Island of St. Helena, in a given Latitude, and Place of a given Variation; how

may the Longitude be known?

Find the yearly Motion of the Variation as in the last Problem, and thereby the Place of the observed Variation in the given Parallel of Latitude, so you have the Longitude required.

154 THE UNIVERSAL

212. Problem CI. Given the Variation of the Compass, at two Places, whose Difference of Latitude and Difference of Longitude are given, to find the Latitude and Longitude of each of those Places.

Example. Suppose the Difference of Latitude between the two Places to be 38°, and their Difference of Longitude 40°, the Variation at the former Place 10° West, and at the latter 20° West; required the Latitude and Longitude of each Place?

Operation. Count among the Meridians and Parallels of Latitude, until you find two Places on the Planisphere, whose Difference of Latitude and Longitude answer to those given, and you have the Places required.

I shall here propose the following Hint, to such Persons as are astronomically inclined, and are desirous of promoting an Improvement by the Angle of the Sun's Position to the Pole and Zenith; which although the Practice of taking the same, may have been hitherto neglected, as a Thing either inconsiderable or impossible; yet that it is not so, will appear on a

due Consideration of the following Particulars.

Definition. In a vertical spherical Triangle, the Angle of the Sun's Position to the Pole and Zenith, is the Distance of the Extremities of two quadrantal Arches, in the Arch of a great Circle, one passing through the Sun and Pole, the other passing through the Sun and Zenith. And seeing that the fix Parts of a vertical spherical Triangle, do depend upon the Sun's Declination, Hour of the Day, Latitude of the Place, Sun's Azimuth, Altitude and Angle of Position, the Want of a practical Method for taking the latter, hath hitherto rendered this most useful Part of Astronomy incompleat. Wherefore,

Construction. Let there be formed an Instrument, having a Gnomen, horizontal Plane, a Pole and Zenith, with two Arches meeting, and the Point where they meet so constructed, as to be readily directed to the Sun; the Arch of a great Circle extending from the Ends of the two former Arches, at the Di-

stance of 90° from the angular Point.

Use. By a Plumb-line, the Latitude of the Place, Hour of the Day, or any other Data, place the whole Instrument in the Position of the Earth's Globe, direct the Meeting of the two Arches to the Sun, and turn them up to the Pole and Zenith; their Distance in the great Circle, from one to the other, is the Angle of Position required.

The AUTHOR, on the Receipt of proper Orders, Post-paid, will cause this Instrument to be made for Land or Sea; the Use of which will be no less than from a single Observation of the Sun, to determine the Latitude, the Azimuth and Hour of the Day; and in Composition with other Data, will be instrumental for resolving many other important Problems.

PARADOXES.

Parallel of Latitude, to which a Ship may be fupposed to sail, and to have Sea-room, to depart from that Parallel; but such is the Situation of every Point of that Parallel of Latitude, that no Ship can sail from the said Latitude, but upon one particular Point of the Compass; and the Wind, though it veer round all the 32 Points, will always blow against that Point upon which the Ship must sail from that Latitude, and the Wind, though blowing against that Point upon which the Ship may go directly before the Wind.

- 214. There are two Places on the Terraqueous Globe so situated, with respect to one another, that from each of those Places a Ship may sail, and though they have Sea-room to all the 32 Points, yet it is impossible for them to sail from those Places but upon opposite Points of the Compass; and though they sail upon opposite Points, yet they may, upon those Points, be either sollowing or meeting one another.
- which the Sun and Moon rife always in the South, and fet in the South.
- North and South of one another, and their Distance from each is 100 Miles, but the true Course from one of those Places to the other, is to sail 50 Miles due North, and 50 Miles due South.
- 217. There is a certain Place on the Terraqueous Globe, to which a Ship may fail directly upon one Point of the Compass; but there are other fourteen Points on which a Ship may fail towards that Place, but if she continue to fail upon e'er a one of the Fourteen, she will always approach to the aforesaid Place, but will never be able to come to it.

- 218. There is a certain Parallel of Latitude upon which a Ship may fail round the Globe, and may all that Time observe the Body of the Sun, without his being above the Horizon.
- Ships may depart on one and the fame Day, yea, and on the fame Hour and Minute of that Day, and upon one and the fame Point of the Compass, may continue to sail with the same Velocity for the Space of 24 Hours, and though they depart from the same Place, and at the same Time, and both sail upon the same Point of the Compass, with the same Velocity, yet, in 24 Hours, they will be distant from one another above 1000 Italian Miles.
- 220. There is a certain Parallel of Latitude upon which a Ship may fail at the Rate of 12 Italian Miles an Hour, and in 24 Hours her Velocity, measured in Degrees of the Equator, shall be equal to the Velocity of the Sun on the 10th of June.
- Stone of about 10,000 Weight could not fall, though a Pit were digged all round it, and that which is the Cause of other Bodies falling from other Places, would be the only Pillar to support a Body in that Place.
- 222. There are two Parallels of Latitude, in one of which less Gold will make a Pound Weight than in the other.
- fame Vessel will hold more Wine or Water in the one Place than the other.
- Ships may sail at one and the same Moment of Time, with respect of real Time, and with respect of apparent Time, will differ one intire natural Day; but there is another Parallel of Latitude, to which, if those two Ships sail upon one and the same Point of the Compass, with the same Velocity, they shall arrive at that certain Parallel of Latitude at one and the same Moment, both with respect to real and apparent Time.

- Stars have still the same Altitude; and notwithstanding of the diurnal Revolution of the Earth, they always bear upon the same Point of the Compass.
- one another, that lie all directly East of London, but London does not lie directly West from any of the said Places.
- rheir Sunday, the Jews the Seventh for their Sabbath, and the Turks the fixth Day of the Week for the Time of their Worship; but there is a particular Place of the Globe, to which, if a Christian, Jew, and Turk sail in one and the same Ship, they shall keep the Time for their Worship on different Days, as above, all the Time they are failing to that particular Place; but when they arrive at that Place, and during the Time they remain at it, they shall all keep their Sabbath on one and the same Day; but when they depart from that Place, they shall all differ as before.
- depart at one and the same Time, and sail on three particular different Courses till they return to the Port they departed from, and if in one of these Ships be Christians, in the second Jews, and in the third Turks, when they return to the Port they departed from, they shall differ so with respect to real and apparent Time, that they all shall keep their Sabbath on one and the same Day of the Week, and yet each of them, separately, shall believe that he keeps his Sabbath on the Day of the Week his Religion requires.
- Zone, to whose Inhabitants the Bodies of the three superior Planets, Saturn, Jupiter and Mars appear to be most enlightened, when they are actually least enlightened.
- 230. There is a Continent all lying in the North-side of the Equator, to whose Inhabitants the Body of the Sun seems to be nearest to them, when he is the farthest removed from them.

- 231. There is a certain Parallel of Latitude, on which Parallel the Star, called the North Polar-star, appears by Observation, to be 47 Degrees elevated above the Horizon, and at another Time the said Polar-star will appear, by Observation on the same Parallel of Latitude, to be at the Horizon.
- 232. There are two Places lying in the Torrid-Zone, that are not above 60 Miles distant from one another, but if a Ship sail from one to the other, on one particular Point of the Compass, the Difference of Time between those two Places, will actually be found to be above 23 Hours.
- 233. There are 24 Places on the terraqueous Globe, in which all the 24 Hours of the natural Day do always exist at the same Time, and yet the Distance of any two of those Places does not exceed 60 Miles.

CONCLUSION.

ROM the aforegoing Problems, and their Solutions, we prefume it sufficiently appears, that our Universal Planispheres, are compleat Substitutes for a large Pair of Globes, and that the capital Problems of Geography, Astronomy and Navigation, may be folved by them. That the Distances and Positions of Places on the Earth and Sea may be measured, and the Place of the Zenith in the Heavens, with the Distances and Positions of Stars, as accurately ascertained, as by the Globes themselves, with their Horizon and Quadrant of Altitude appending. Which, although fomewhat paradoxical, feeing that the Globe hath Dimensions of Length, Breadth and Depth, and a Rotation round its Axis; whereas our Planispheres have no fuch Properties, but are fixed to the Plane of the Meridian and folfticial Colure, with their Slider moving in a direct Oppofition to fuch a Motion, coincident with the Plane of the Meridian and folfticial Colure from the Zenith to the Nadir, or from Pole to Pole; yet their Operations are undoubtedly true, and may be expected to become familiar and easy in Practice, to fuch as shall have attentively considered their Construction and the aforegoing Examples with the Precepts.

MR. DUNN, the AUTHOR of this Treatise,

having been engaged, as a Conductor and Teacher,
at the most famous Academies in London;
And had the Honour to instruct Persons of Distinction,
at their own Houses, with Approbation;
also, grown Gentlemen and Ladies, at Home and Abroad,
by new, easy, and short Methods:
is removed to a very commodious and healthy Situation,
and with the Assistance of proper Masters, now carries on

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